Isfu





The NEVFAR project:

New Evaluation of v Fluxes At Reactors

DE LA RECHERCHE À L'INDUSTRIE



Revisiting the summation calculation of reactor antineutrino spectra

Lorenzo Périssé<sup>(a)</sup>

Xavier Mougeot, Anthony Onillon<sup>(b)</sup>, Matthieu Vivier

CEA/IRFU/DPHP – CEA/List/LNE-LNHB

CEA-Saclay, 91191 Gif-sur-Yvette, FRANCE

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www.cea.fr

<sup>(a)</sup>Now at ILANCE (CNRS/UTokyo), Japan <sup>(b)</sup>Now at TUM, Germany



#### 1. Introduction & motivations

- a. Experimental anomalies
- b. Modeling methods

#### 2. Revised summation method

- a.  $\beta^{-}$  spectrum calculation
- b. Nuclear data content
- c. Uncertainty budget

#### 3. Comparison to experiments and models

- a. Integral measurements
- b. Spectrum shape

#### 4. Conclusion & perspectives



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#### REACTOR ANTINEUTRINO ANOMALY (RAA)

- Systematic IBD rate deficit vs to HM
- Measured/predicted IBD rate: **0**. **936**<sup>+0.024</sup><sub>-0.023</sub> (2.5σ)
- RAA possible origins
  - Experimental bias

- Unlikely
- New physics (sterile neutrino)
- $\blacktriangleright$  Mismodeling / underestimation of  $\overline{\nu}_e$  spectrum uncertainty
- Single / multiple actinide(s) ?

#### **SHAPE ANOMALY**

- First observed by Double Chooz, Daya Bay; RENO
  - Confirmed by recent very-short baseline reactor exp. (NEOS, STEREO, PROSPECT, DANSS)
- Possible origins
  - Detector energy scale calibration
     Checked
  - Fuel composition
  - Prediction issue, single / multiple actinide(s) ?

### **FUEL-DEPENDENT IBD RATE EVOLUTION**

- IBD yield changes with fuel evolution of PWR
- Comparison between measured IBD yield evolution and predicted evolution
  - $3.1\sigma$  at Daya Bay
  - 1.3σ at RENO
- Induced by inequal fractional deficit among actinides







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#### **REACTOR DATA-DRIVEN METHOD**

- Unfolding exp. prompt IBD spectrum
  - $\succ \bar{v}_e$  spectrum + covariance matrix

### PROS الم

- Model-independent (no anomalies)
- Small uncertainties

## E CONS

- Limited to exp. range, 1.8-9 MeV
- Small number of available datasets
- No activation spectrum

Daya Bay: Total, <sup>235</sup>U, <sup>239</sup>Pu RENO, NEOS: Total

STEREO, PROSPECT: <sup>235</sup>U

#### b. Different modeling methods

#### **CONVERSION METHOD**

- Measure exp.  $\beta$  fission spectra
- Convert virtual  $\beta$  branch fit to  $\bar{\nu}_e$  branches

### 沟 PROS

- Small uncertainties ~2-3%
- Access total  $\bar{v}_e$  fission spectrum

## E CONS

- Limited to exp. range, 2-8 MeV
- No activation spectrum
- HM subject to the anomalies
- BILL data questionned  $\rightarrow$  KI exp.
- Impact of forbidden branches on fit

#### **SUMMATION METHOD**

- Fission spectrum prediction = sum of all β branches listed in nuclear databases
- +900  $\beta^-$  emitters ~ 10 000  $\beta^-$  transitions

### 沟 PROS

- Prediction ∀ energy, ∀ β emitter
   ► CEvNS
- Convenient to understand physics
- Mandatory for activation spectra

#### 

- Uncomplete/biased nuclear database
- Modeling approximations
- Uncertainties very complex to estimate

Huber-Mueller model (+ KI data)

#### ⇒ <sup>235</sup>U, <sup>239</sup>Pu and <sup>241</sup>Pu from P. Huber PRC 84, 024617 (2011)

 $\Rightarrow \frac{^{235}\text{U}/^{239}\text{Pu data from Kl}}{^{\text{PRD 104, L071301 (2021)}}}$ 

 $\Rightarrow {}^{238}\text{U from Mueller et al.} \\ \underline{PRC 83, 054615 (2011)}$ 

#### b. Different modeling methods

#### **THE NEVFAR PROJECT**

(New Evaluation of v Fluxes At Reactor)



- Revise summation method with BESTIOLE code
  - Improve β-decay modeling
    - Refine non-unique forbidden transition modeling
  - ▷ Impact of database uncompleteness and quality
    - Update nuclear database with Pandemonium-free data
    - Adjusted effective modeling for nuclides with no data
  - Build a comprehensive uncertainty budget
    - Nuclear data and modeling uncertainties

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#### ⇒ Reliable summation method required for multiple purposes



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#### **MODELING OF NON-UNIQUE TRANSITIONS**

- Disregarded in previous modeling (modeled as allowed or unique forbidden)
- Hayes *et al.* (2014) + Hayen *et al.* (2019): modelings of non-unique transitions in conversion predictions → partial explanation of shape anomaly
- Nuclear structure calculation with NuShellX
  - Very time consuming (man & cpu)
  - No general nor systematic trend
- 23 non-unique forbidden transitions contribute to  $|\sim$  27% of IBD yield

~22% of CE $\nu$ NS yield

#### $\Rightarrow$ Using NSC decreases IBD yield by (1.3 $\pm$ 0.2)%



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#### TACKLING THE PANDEMONIUM EFFECT IN SUMMATION SPECTRA

HPGe detector, high energy resolution + decreasing efficiency for increasing energies

- $\beta$  feedings to low (high) energy levels are overestimated (underestimated)
- Nuclear database are biased by the Pandemonium effect
  - Estienne et *al.* (2019): including Pandemonium-free TAGS data decreases IBD yields and shape differences

• Including up-to-date Pandemonium-free data (TAGS + Direct β measurements)

- $\Rightarrow$  IBD yield decreased by (12.8  $\pm$  1.5) %
- $\Rightarrow \sim$  65% of IBD and CE $\nu NS$  yields
- Remaining isotopes potentially impacted by Pandemonium in nuclear database
  - 29 isotopes identified by IAEA
  - Apply correction for residual Pandemonium effect
    - $\Rightarrow$  IBD yield decreased by (2.2  $\pm$  2.4) %
    - $\Rightarrow \sim$  12% of IBD and CE $\nu NS$  yields



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#### c. Uncertainty budget



IBD yields (10 <sup>-43</sup> cm <sup>2</sup> /fission)				
<sup>235</sup> U:	6.25 ± 0.21			
<sup>238</sup> U:	10.01 ± 0.32			
<sup>239</sup> Pu:	4.48 ± 0.15			
<sup>241</sup> Pu:	6.58 ± 0.21			

 $\Rightarrow$  IBD yield uncertainty  $\sim$ 3%

<b>CE</b> v <b>NS yields</b> * (10 <sup>-43</sup> cm <sup>2</sup> /fission)		
<sup>235</sup> U:	1113 ± 34	
<sup>238</sup> U:	1669 ± 48	
<sup>239</sup> Pu:	882 ± 25	
<sup>241</sup> Pu:	1169 ± 33	

\* For a Ge target nucleus and 20 eV detector threshold

 $\Rightarrow$  CEvNS yield uncertainty  $\sim$ 3%

## **NORMALIZATION UNCERTAINTY**

### **FRACTIONAL UNCERTAINTY**

PWR				$\langle \sigma_{IBD} \rangle$	$\langle \sigma_{CE\nu NS} \rangle$	
[10 <sup>-4;</sup>	<sup>3</sup> cm <sup>2</sup> /fission]			6.08	1090	
	Uncertainty	Abbrev.	Method	[%]	[%]	
	Endpoint + Spin-parity	E <sub>0</sub> + Jπ	MC	0.1	0.1	
DATA	Branching ratio + $\beta^-$ intensity	$BR + I_{\beta}$	MC + Analytic	0.4	0.3	
	Residual Pandemonium	RP	Analytic	2.5	2.4	
	Direct $\beta$ measurement	Dβ	Analytic	1.5	1.2	
	Nuclides with no data	NND	Pool modeling	0.8	0.5	
	Fission yield	FY	Analytic	~0.7	~0.6	
	Fission fraction		Analytic	~0.7	~0.7	
MODELING	Weak magnetism	WM	Model comparison	0.3	0.2	
	Radiative corrections	RC	Model comparison	0.1	0.1	
	Non-unique transitions	NU	Model comparison	0.4	0.4	
	Nuclear struct. calcul.	NSC		0.2	0.1	
	<ul> <li>ξ-approximation</li> </ul>	ξ		0.3	0.3	
	Cross-section		Analytic	0.1	0.5	
	TOTAL			3.1	2.9	

\* For a Ge target nucleus and 20 eV detector threshold



# ⇒ Uncertainty budget dominated by RP and Dβ (+ NND at high energy)



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#### 3. Comparison to experiments and models

a. Integral measurements

Predictions and Bugey-4 taken from <u>Giunti et al., Phys. Lett. B, 829, 137054 (2022)</u> 1: PRL 123, 111801 (2019) 2: PRD 104, L111301 (2021) 3: PRL 125, 201801 (2020)





- DB / BESTIOLE =  $0.982 \pm 0.015$  (exp)  $\pm 0.031$  (model)
- DB / HM = 0.945  $\pm$  0.014 (exp)  $\pm$  0.024 (model)
- $\Rightarrow$  Significance at 0.5 $\sigma$  for BESTIOLE and 1.9 $\sigma$  for HM

- $\Rightarrow$  BESTIOLE consistent within  ${\sim}2\sigma$  with global rate analysis
- $\Rightarrow$  Discrepancy with HM favors RAA caused by  $^{235}\text{U}$  HM flux



 $\Rightarrow$  Impact of FY seen in upper energy range

STEREO + PROSPECT data from <u>Almazán et ak. (2022)</u> Daya Bay + PROSPECT data from <u>An et al. (2022)</u>

## **RATIO OF IBD SPECTRA**

Shape only comparison, predictions normalized to data

- Gaussian distorsion not significantly favored in 5-7 MeV
  - Gaussian bump hypothesis favored by  $\leq 2.3\sigma$

⇒ Overall good shape agreement with experimental IBD spectra within uncertainty





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## **KEY POINTS OF BESTIOLE SUMMATION PREDICTION**

#### All modeling impacts considered and quantified

- Nuclear structure calculation for 23 non-unique branches
  - ▶ IBD yield decreased by (1.3 ± 0.2)%

#### Quality of data checked for all data sources

- Correction for Residual Pandemonium
  - ▶ IBD yield decreased by (2.2 ± 2.4)%
  - Measurement needed to validate RP correction

#### **Comprehensive uncertainty budget**

Uncertainty budget of summation model for the first time ever

#### **Complete revision of summation method**

- Overall good agreement with data
- Results favors RAA caused by <sup>235</sup>U HM flux

#### Next steps for further improvement...

- Fission yield correlation matrix for data and evaluation
- Remaining non-unique forbidden branches

⇒ Article on arXiv with supplementary materials, soon to be published

# Final IBD and CEvNS yield uncertainty budget ~3%

#### Led by RP correction

⇒ more Pandemonium-free data needed



v kinetic energy [MeV]

Reach of a comprehensive summation model, needed for validation