



UNIVERSITY OF
LIVERPOOL

Upgrade of the T2K near detector

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HEP seminar, University of Liverpool

29th January 2025

T2K

Introduction

In this seminar I will cover the following..

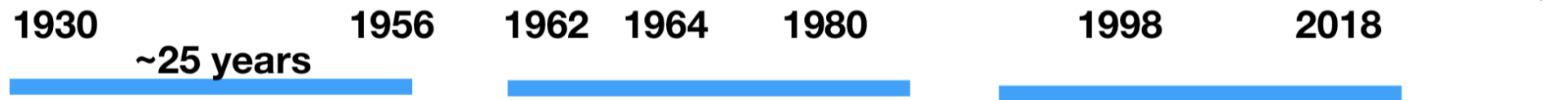
1. General overview of oscillation experiments, T2K and the importance of near detectors
2. The design, construction and installation of the upgraded near detector
3. A first look at data taken with the upgraded near detector and physics potential
4. A brief look ahead to HK era and potential further upgrades

Part One: Overview of T2K experiment and ND280

Adapted "The Growing Excitement of Neutrino Physics" by APS

- ★ 1930: On-paper appearance as "desperate" remedy by W. Pauli
- ★ 1956: Anti- ν_e first experimentally discovered by Reines & Cowan
- ★ 1962: ν_μ existence confirmed by Lederman *et al*
- ★ 1986: Existence of ν_τ was established (see Gary Feldman's talk)
- ★ 1998: Atmospheric ν oscillations discovered by Super-K
- ★ 2000: ν_τ first evidence reported by DONUT experiment
- ★ 2001: Solar ν oscillations detected by SNO (KamLAND 2002)
- ★ 2011: $\nu_\mu \rightarrow \nu_\tau$ transitions observed by OPERA
- ★ 2011-13: $\nu_\mu \rightarrow \nu_e$ observed by T2K and *anti*- $\nu_e \rightarrow$ *anti*- ν_e by Daya Bay
- ★ 2015: Nobel prize for ν oscillations, Breakthrough prize (2016)
- ★ 2018: T2K hints on leptonic CP violation

Pauli predicts the Neutrino Fermi's theory of weak interactions Reines & Cowan discover (anti)neutrino muon neutrinos discovery Solar neutrino anomaly



T2K hints on leptonic CP violation
 IceCUBE observes extragalactic ν
 Nobel prize & Breakthrough prize for ν oscillation
 T2K observe ν_e appeared from ν_μ
 Daya Bay observe anti- ν_e disappeared
 K2K confirm atmospheric ν oscillation
 KamLAND confirms solar ν oscillation
 Nobel prize for ν astrophysics
 SNO observe solar ν oscillation to active flavor
 Super-K confirms solar ν deficit and images the sun
 Super-K observes ν oscillation
 Nobel Prize for ν_μ discovery
 Kamioka-II/ IMB observe supernova ν
 SAGE/Gallex observe the solar ν deficit
 LEP shows 3 active flavors
 Kamioka-II confirms solar deficit

Taken from S. Cao's slide at "3 neutrinos and beyond" 2019

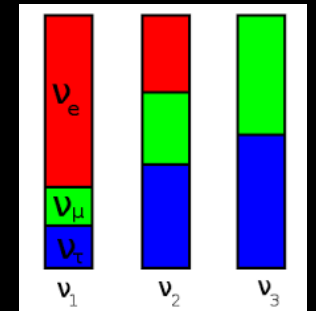
Neutrino oscillation

Flavour states



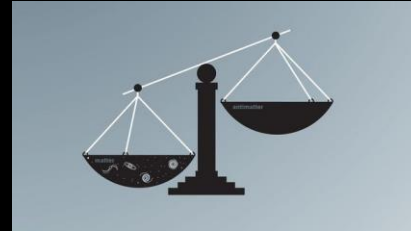
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Mass states



- Discovery of neutrino oscillation → neutrinos have mass
- Each neutrino flavour state (electron, muon or tau) is a superposition of different mass states
- The PMNS matrix describing neutrino mixing contains four parameters
 - Three mixing angles, θ_{12} , θ_{23} , and θ_{13}
 - The CP-violating phase, δ_{CP}
- The probability of neutrino oscillation also depends on
 - Mass squared differences between the mass eigenstates (Δm_{12}^2 and Δm_{23}^2)
- Experiments detecting atmospheric, accelerator, reactor and solar neutrinos have different complementary sensitivity to these neutrino oscillation parameters

Neutrino oscillation



$\delta_{CP} \neq 0$ or π implies CP violation in neutrino sector

Flavour states

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

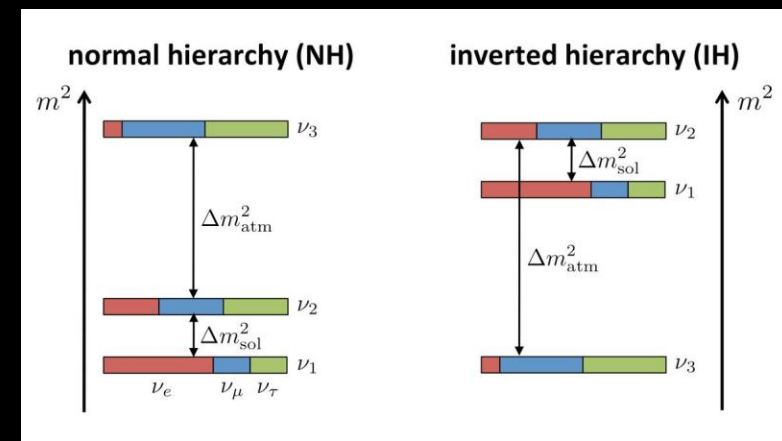
Mass states

		Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 6.1$)	
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range
IC24 with SK atmospheric data	$\sin^2 \theta_{12}$	$0.308^{+0.012}_{-0.011}$	0.275 \rightarrow 0.345	$0.308^{+0.012}_{-0.011}$	0.275 \rightarrow 0.345
	$\theta_{12}/^\circ$	$33.68^{+0.73}_{-0.70}$	31.63 \rightarrow 35.95	$33.68^{+0.73}_{-0.70}$	31.63 \rightarrow 35.95
	$\sin^2 \theta_{23}$	$0.470^{+0.017}_{-0.013}$	0.435 \rightarrow 0.585	$0.550^{+0.012}_{-0.015}$	0.440 \rightarrow 0.584
	$\theta_{23}/^\circ$	$43.3^{+1.0}_{-0.8}$	41.3 \rightarrow 49.9	$47.9^{+0.7}_{-0.9}$	41.5 \rightarrow 49.8
	$\sin^2 \theta_{13}$	$0.02215^{+0.00056}_{-0.00058}$	0.02030 \rightarrow 0.02388	$0.02231^{+0.00056}_{-0.00056}$	0.02060 \rightarrow 0.02409
	$\theta_{13}/^\circ$	$8.56^{+0.11}_{-0.11}$	8.19 \rightarrow 8.89	$8.59^{+0.11}_{-0.11}$	8.25 \rightarrow 8.93
	$\delta_{CP}/^\circ$	212^{+26}_{-41}	124 \rightarrow 364	274^{+22}_{-25}	201 \rightarrow 335
	$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.49^{+0.19}_{-0.19}$	6.92 \rightarrow 8.05	$7.49^{+0.19}_{-0.19}$	6.92 \rightarrow 8.05
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.513^{+0.021}_{-0.019}$	+2.451 \rightarrow +2.578	$-2.484^{+0.020}_{-0.020}$	-2.547 \rightarrow -2.421

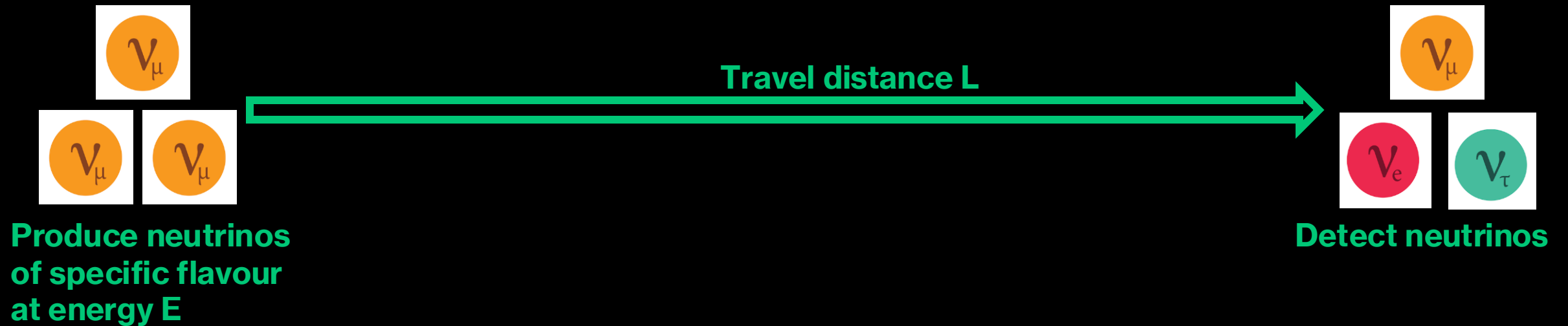
Best known values of neutrino oscillation parameters

Taken from JHEP 12 (2024) 216

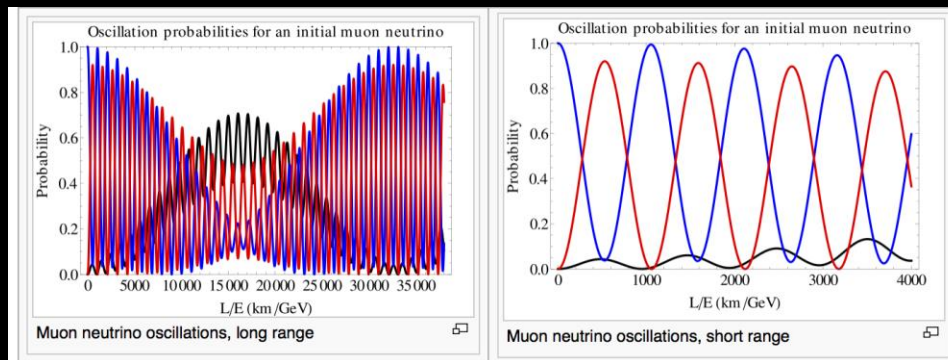
Normal or inverted mass hierarchy?



Long baseline experiments



Long baseline experiments



Produce neutrinos of specific flavour at energy E

Travel distance L



Detect neutrinos

$$\begin{aligned}
 P_{\nu_\alpha \rightarrow \nu_\beta}(t) &= |\langle \nu_\beta | \nu_\alpha(t) \rangle|^2 \\
 &= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2\left(\frac{\Delta m_{ij}^2 L}{4E}\right) \\
 &\quad + 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin\left(\frac{\Delta m_{ij}^2 L}{4E}\right),
 \end{aligned}$$

where $\Delta m_{ij}^2 = m_i^2 - m_j^2$ and we can rewrite:

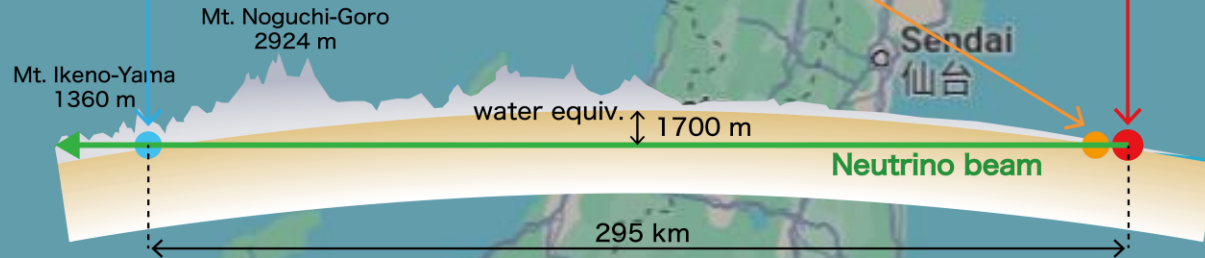
$$\frac{\Delta m_{ij}^2 L}{4E} \approx 1.267 \frac{\Delta m_{ij}^2 [eV^2] \times L [km]}{E [GeV]}.$$

T2K experiment

Super Kamiokande

Near Detector

J-PARC



Neutrino beam



Japan

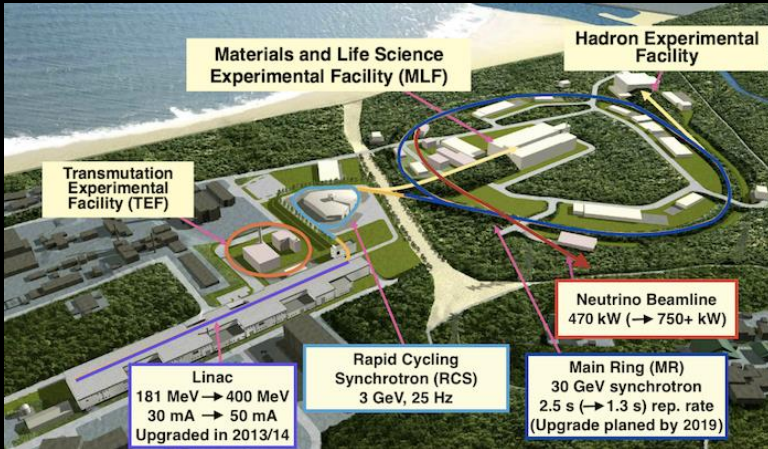
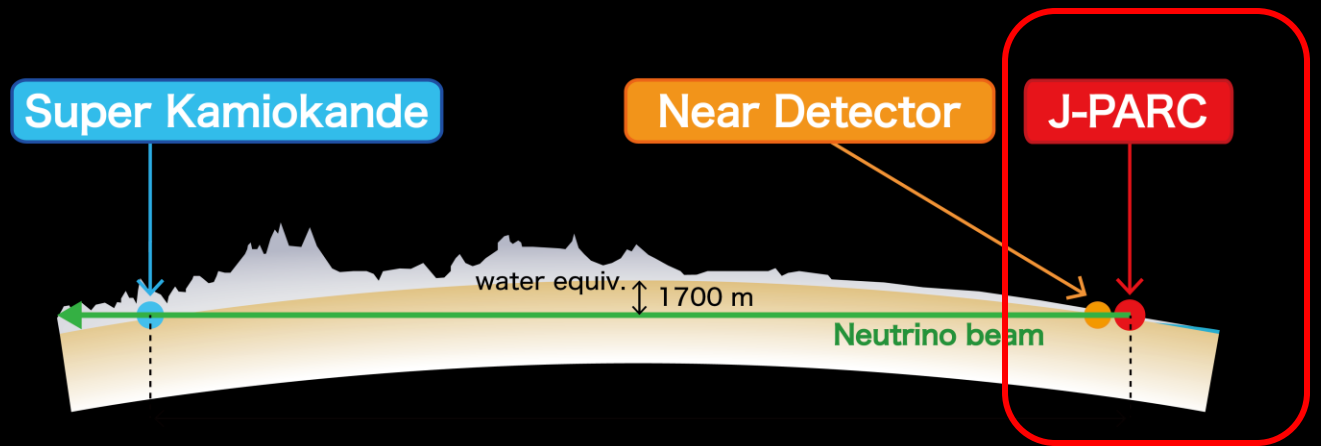
Tokyo
東京

Nagoya
名古屋

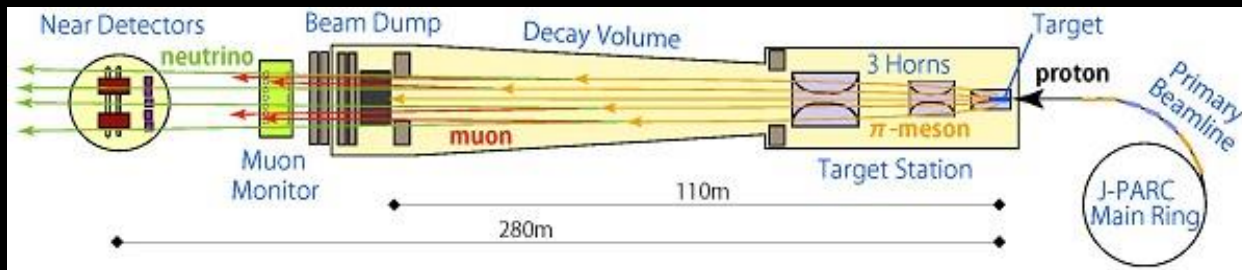
Kobe
神戸

Shikoku
四国

1. Beam production



- High energy protons collided on graphite target
- Charged pions and kaons produced which then decay inside a decay volume
- This produces **neutrinos or anti-neutrinos** depending on whether +/- charged pions are focused with magnetic horn



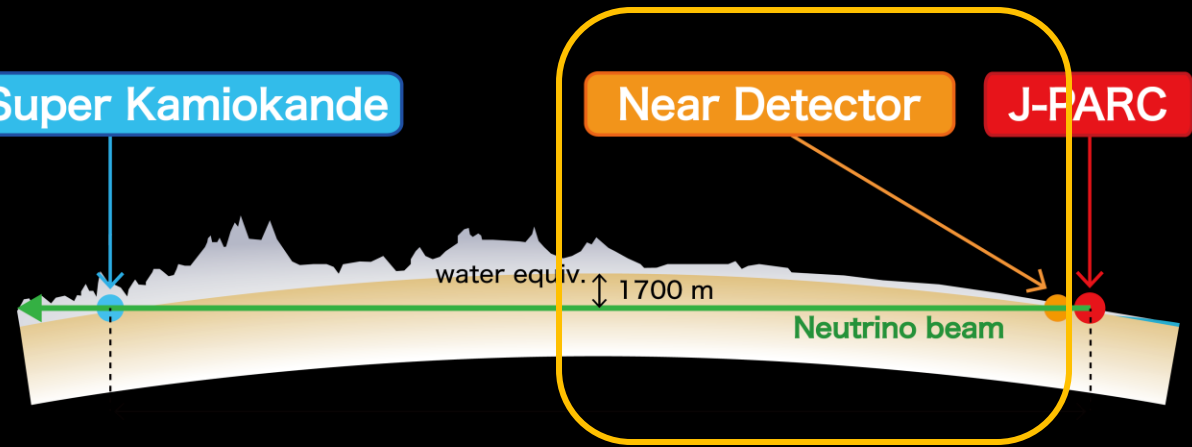
direction of Super-K

2. Near detectors

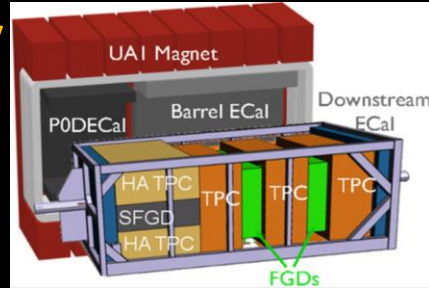
Super Kamiokande

Near Detector

J-PARC

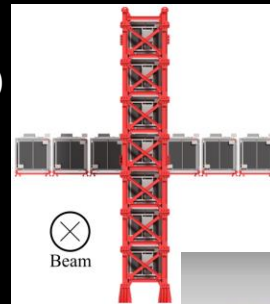


ND280



2.5 deg off axis

INGRID



0 deg off axis

Beam direction

Beam from JPARC

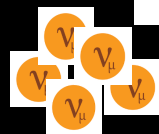
Wagasci-BabyMIND



1.5 deg off axis

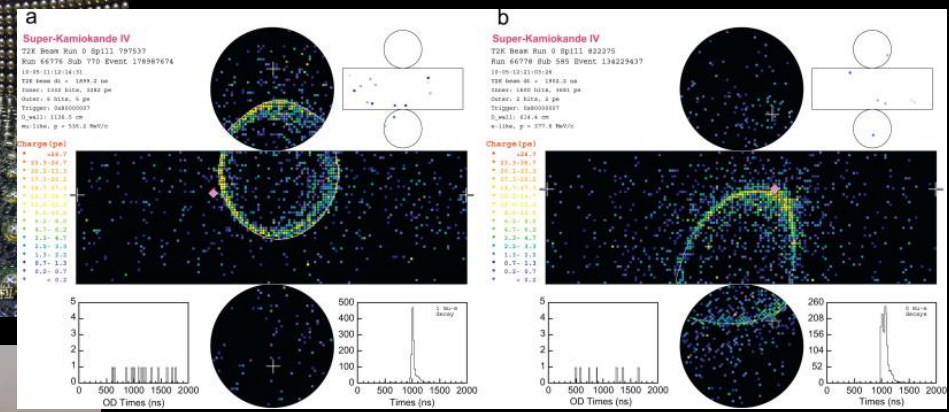
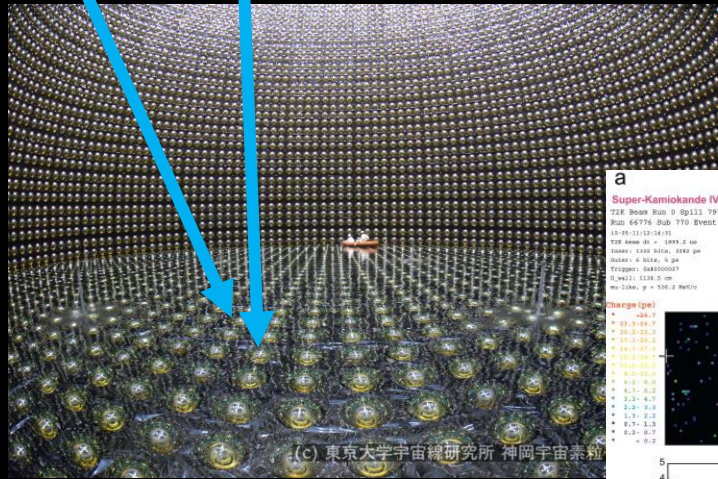
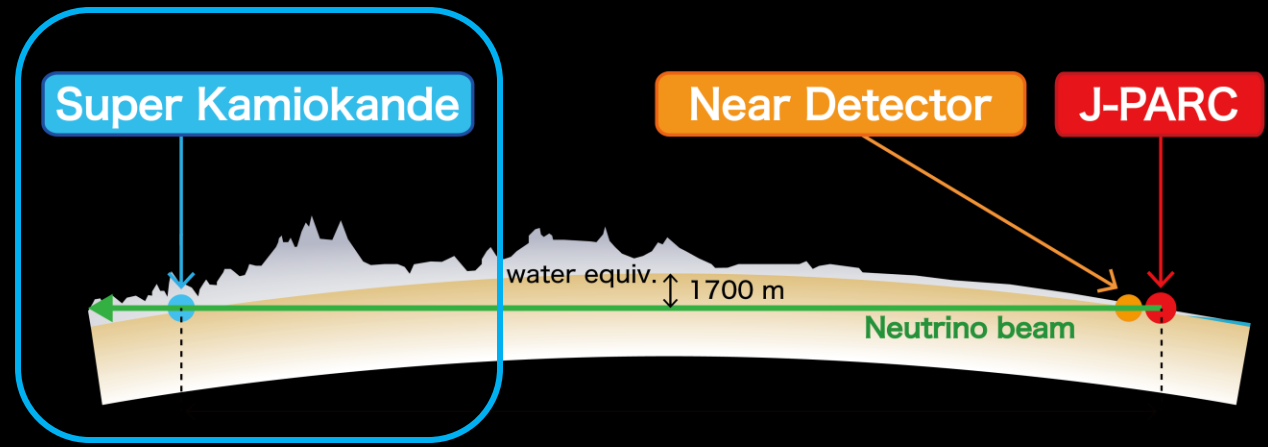
280 m

- Different targets
- Different off-axis angles give different neutrino energy spectra (peak energy and broadness)

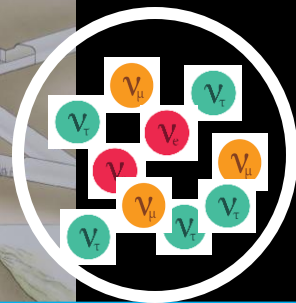
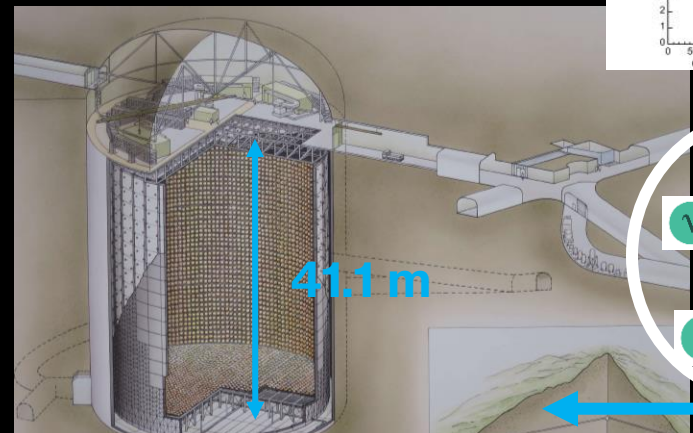


as well as neutrinos from the JPARC beam, SK also detects atmospheric and solar neutrinos

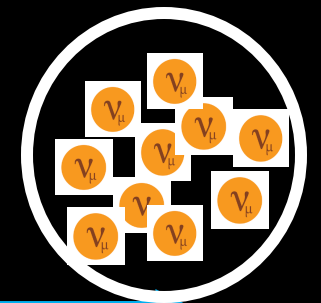
3. Far detector



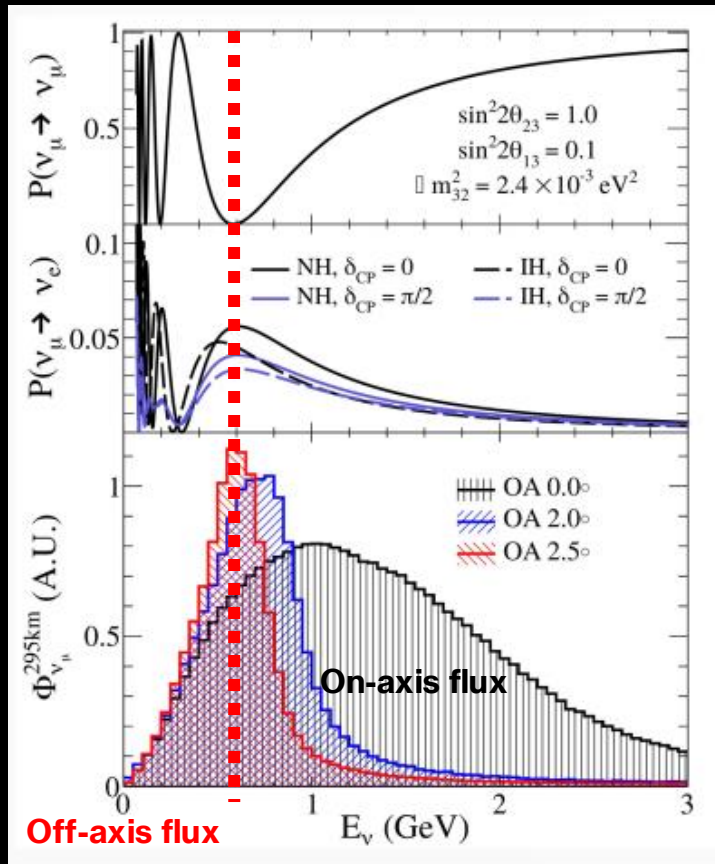
- Water Cherenkov detector (50,000 tons of water)
- Cherenkov light detected by 13,000 PMTs



295 km

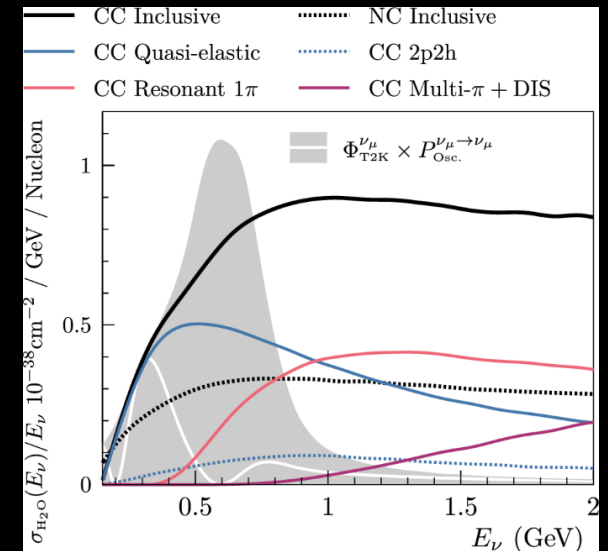


Why are our detectors "off-axis"?



Neutrino energy

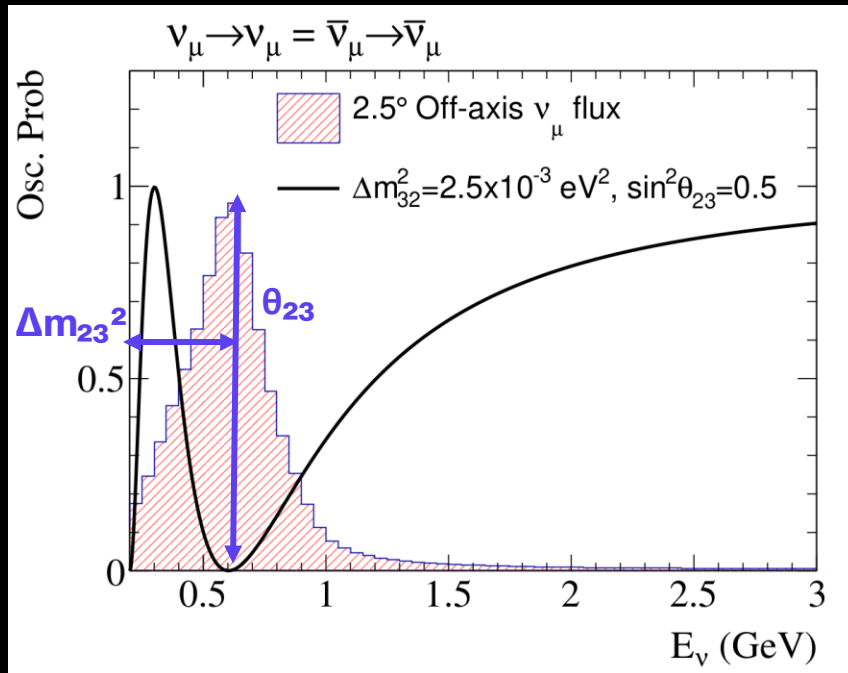
- ND280 and SK detectors sit 2.5° off-axis
- This results in a neutrino beam energy distribution which is **narrower**, with less spread in energies
- Peak of the energy distribution also shifts with the off-axis angle, to **600 MeV peak** at 2.5°
- This energy corresponds to **maximum neutrino oscillation** at 295 km
- This distribution also results in more CCQE like events and fewer high energy backgrounds



T2K measurements

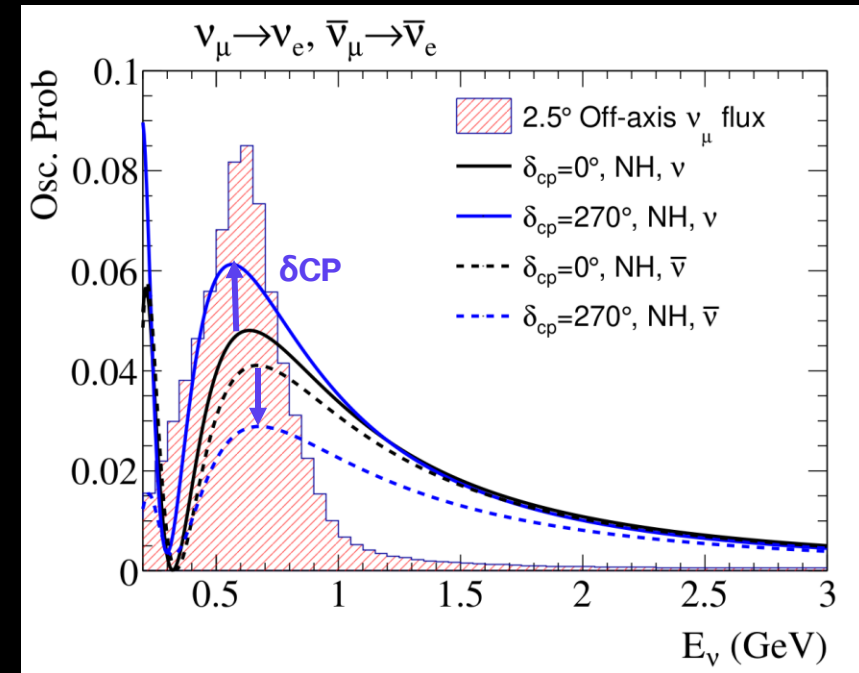
Muon neutrino disappearance

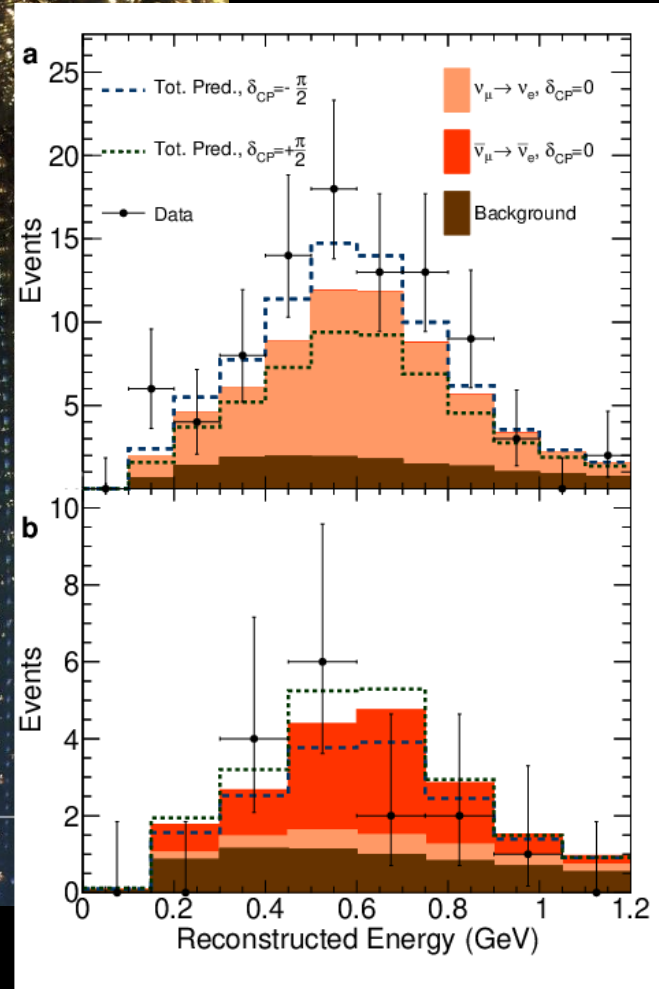
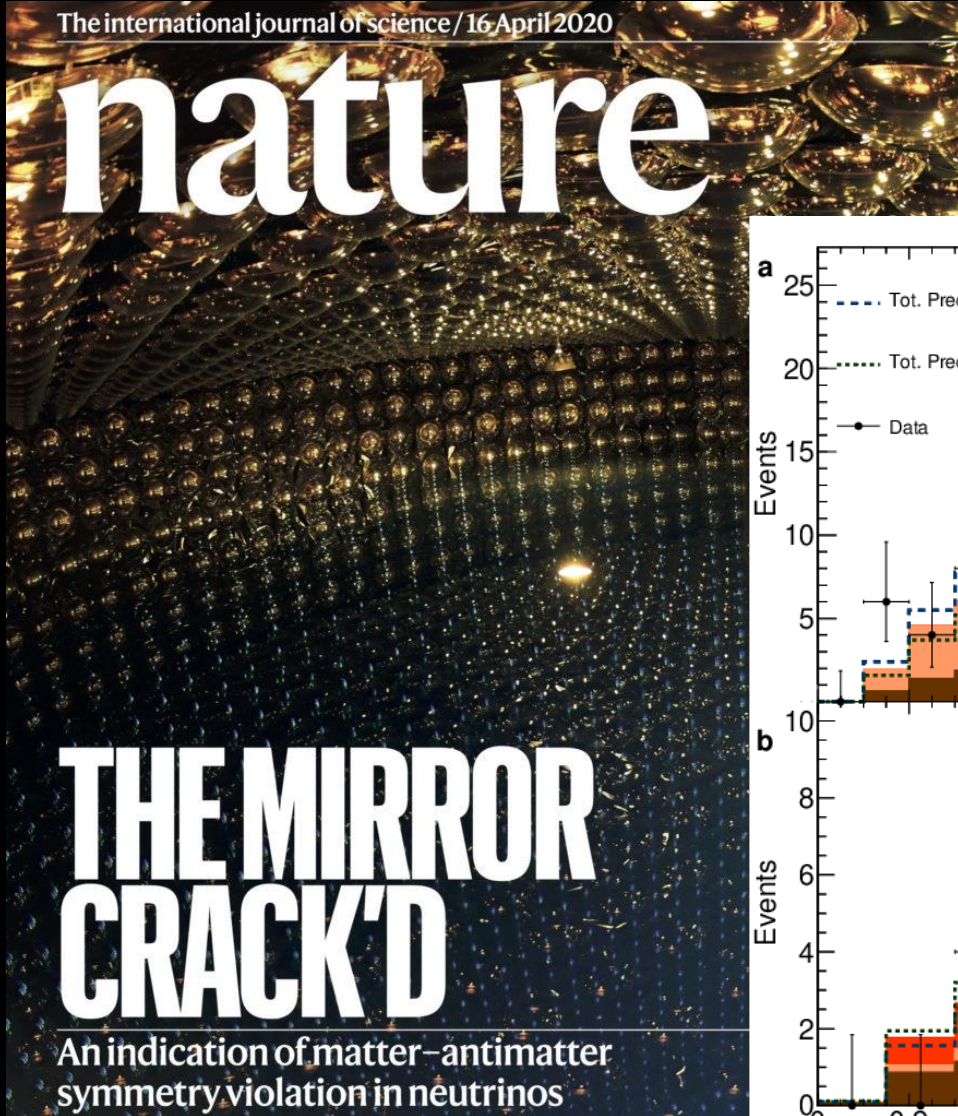
$$P_{\mu \rightarrow x} \approx 1 - \sin^2 2\theta_{23} \cdot \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E_\nu} \right)$$



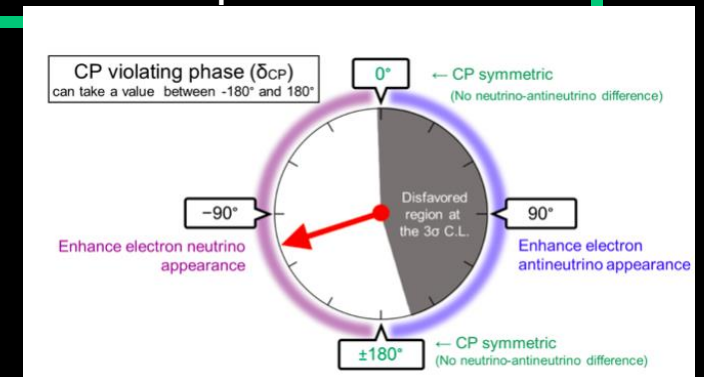
Electron neutrino appearance

$$P_{\mu \rightarrow e} \approx \sin^2 \theta_{23} \cdot \sin^2 2\theta_{13} \cdot \sin^2 \left(\frac{\Delta m_{23}^2 L}{4E_\nu} \right)$$





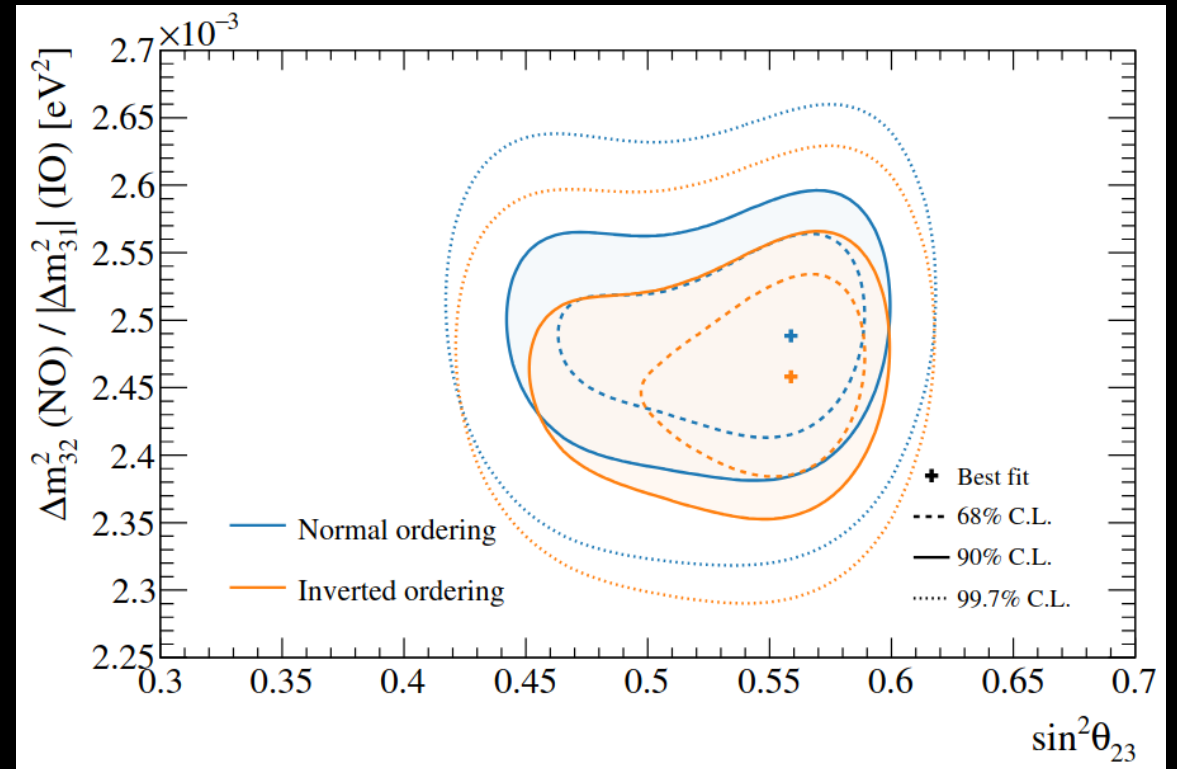
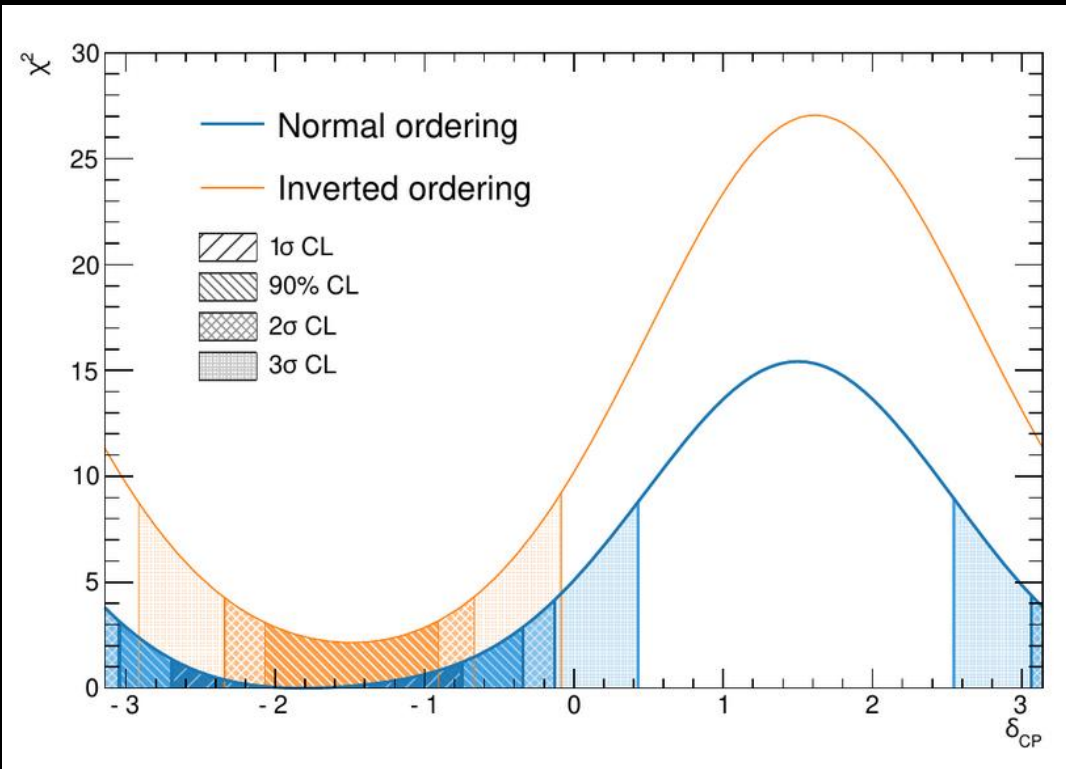
- In 2012 T2K produced the first evidence of $\theta_{13} \neq 0$ with a measurement of electron neutrino appearance in a muon neutrino beam
- By 2020 T2K were able to compare the electron neutrino / antineutrino appearance to provide a strong indication for CP violation in leptons



Recent results

Updated oscillation analysis

- 3.6×10^{21} protons on target (POT)
- Normal ordering and upper octant favoured
- $\delta_{CP} = 0$ and π excluded at more than **90% confidence level**
- Nearly maximal CP violation favoured
- Measured $\sin^2(\theta_{23})$ consistent with reactor experiments



Eur. Phys. J. C 83, 782 (2023)

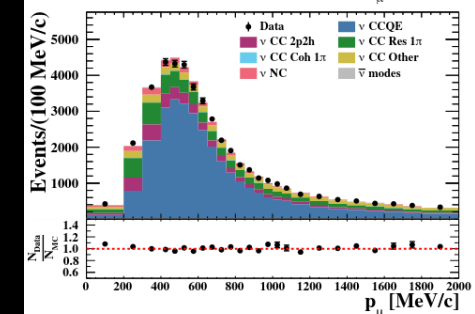
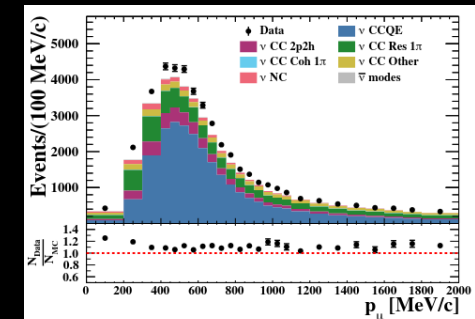
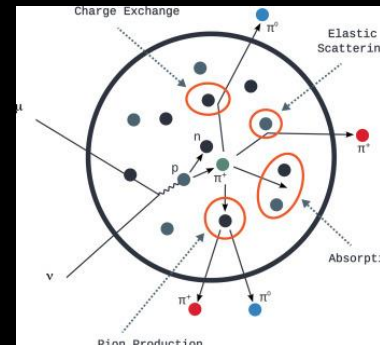
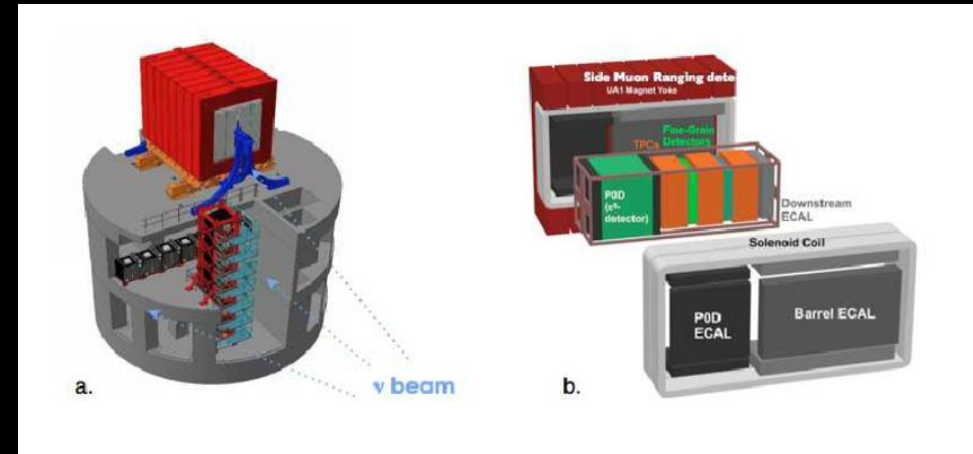
Plus recent **joint analyses**: T2K-SK and T2K-NOvA

Near detector

$$N_{\text{ND}}^{\alpha}(\vec{x}) = \Phi^{\alpha}(E_{\nu}) \times \sigma^{\alpha}(\vec{x}) \times \epsilon_{\text{ND}}^{\alpha}(\vec{x})$$

$$N_{\text{FD}}^{\alpha}(\vec{x}) = P(\nu_{\alpha} \rightarrow \nu_{\alpha}) \times \Phi^{\alpha}(E_{\nu}) \times \sigma^{\alpha}(\vec{x}) \times \epsilon_{\text{FD}}^{\alpha}(\vec{x})$$

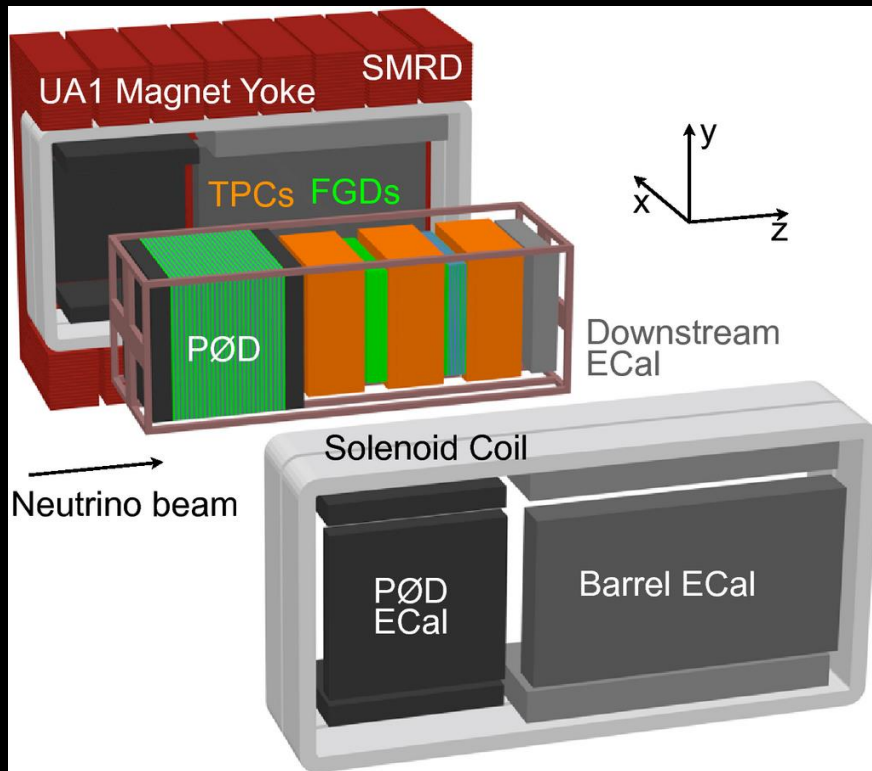
- The **beam flux** and **interaction model** are consistent between the near and far detector
- Carry out a near detector fit to constrain a model describing these + ND detector effects
- As such the near detector provides important measurements of
 - Electron neutrino contamination of the beam (important to ensure any electron neutrino appearance is actually due to oscillation)
 - Flux and energy distribution of beam
 - Neutrino interaction rates and cross sections



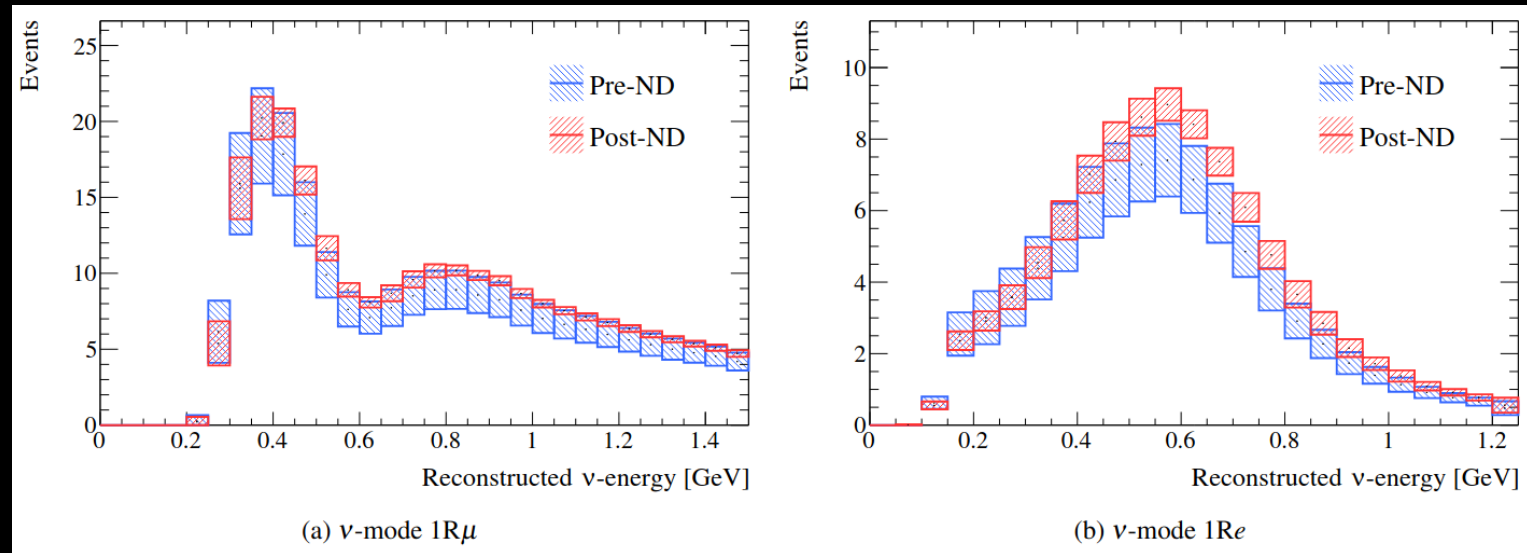
(a) FGD1 ν -mode ν_{μ} CC0 π

Pre-upgrade ND280

- The UK and Liverpool have a long history of working with ND280
- The electromagnetic calorimeter (ECal) was designed and constructed in the UK during 2007-2010
 - The barrel ECal was built between Daresbury lab and Liverpool



Systematic uncertainties



Sample		Uncertainty source (%)			Flux \otimes Interaction (%)	Total (%)
		Flux	Interaction	FD + SI + PN		
1R μ	ν	2.9 (5.0)	3.1 (11.7)	2.1 (2.7)	2.2 (12.7)	3.0 (13.0)
	$\bar{\nu}$	2.8 (4.7)	3.0 (10.8)	1.9 (2.3)	3.4 (11.8)	4.0 (12.0)
1Re	ν	2.8 (4.8)	3.2 (12.6)	3.1 (3.2)	3.6 (13.5)	4.7 (13.8)
	$\bar{\nu}$	2.9 (4.7)	3.1 (11.1)	3.9 (4.2)	4.3 (12.1)	5.9 (12.7)
1Re1de	ν	2.8 (4.9)	4.2 (12.1)	13.4 (13.4)	5.0 (13.1)	14.3 (18.7)

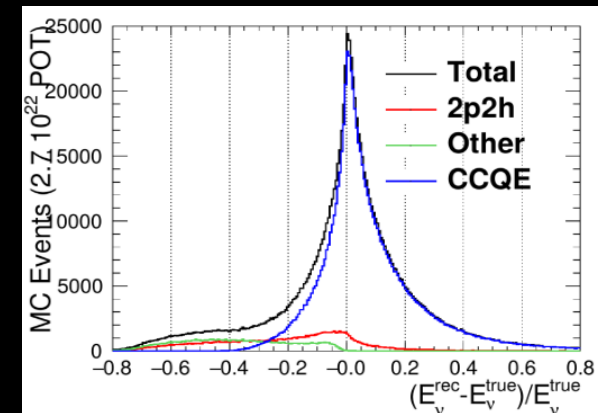
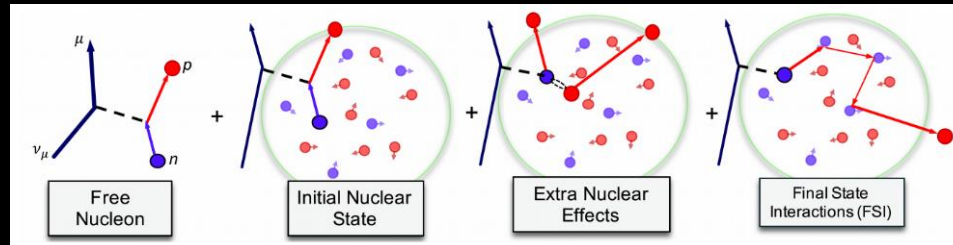
Interaction model uncertainties are the largest source of uncertainty for the oscillation analysis \rightarrow the near detector is crucial for reducing these

Total event rate uncertainty reduction from ND fit: $\sim 17\% \rightarrow \sim 5\%$

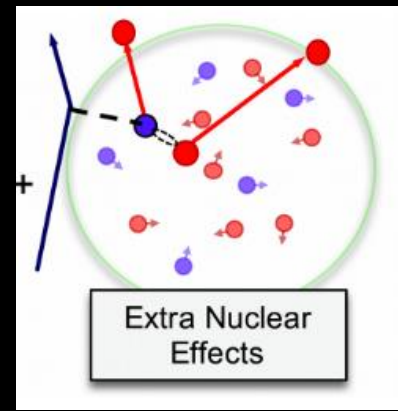
Interaction model

- Reconstruction of the neutrino energy for QE scattering on a nucleon at rest:
- In reality there are effects from
 - Nucleons may have non-zero momenta
 - Final state interactions
 - Interaction on two nucleons (2p-2h)
- Require sufficient models and uncertainties to cover any physics that could affect the oscillation analysis
- Measure interaction topologies (CC0pi, CC1pi etc) rather than interaction modes (CCQE, 2p2h, CCRES)

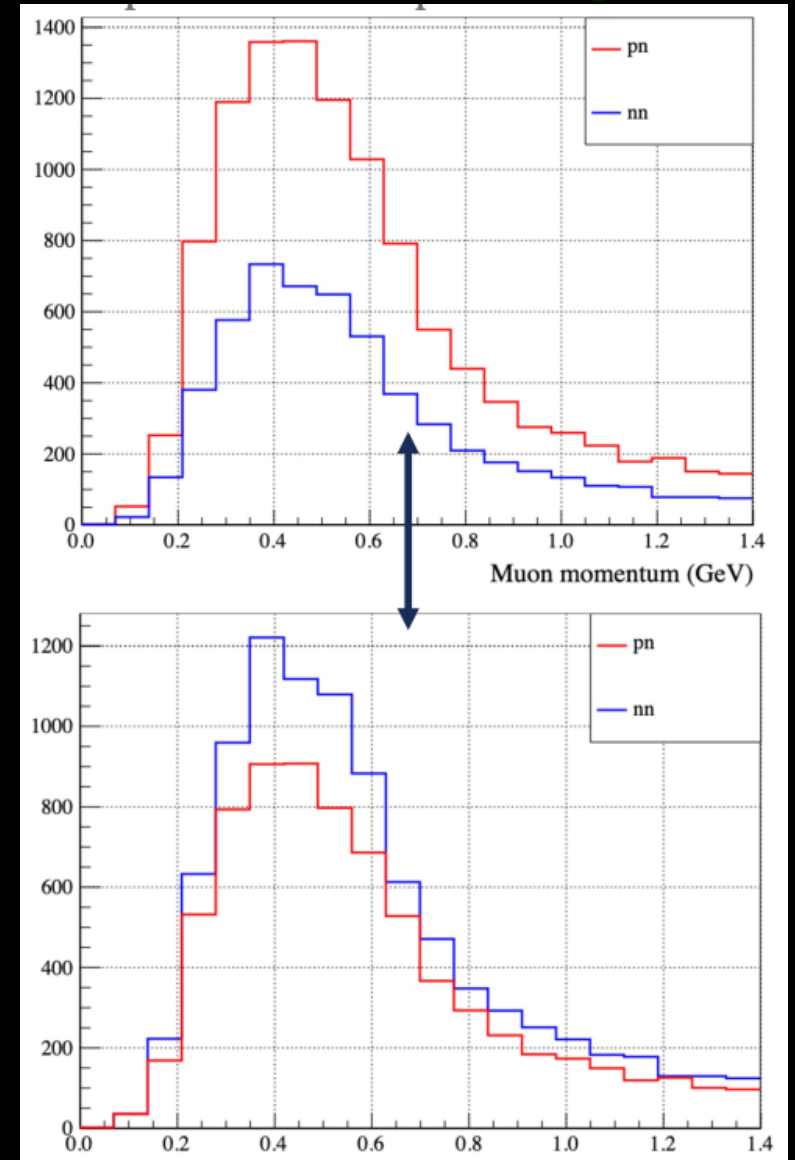
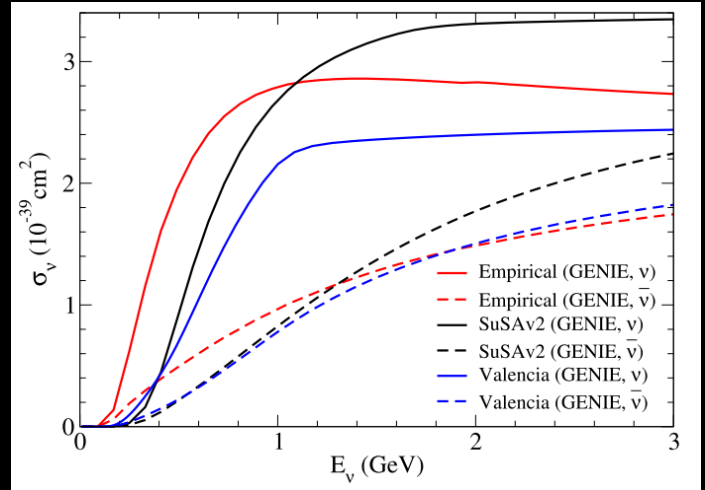
$$E_\nu = \frac{m_p^2 - (m_n - E_b)^2 - m_\mu^2 + 2(m_n - E_b)E_\mu}{2(m_n - E_b - E_\mu + p_\mu \cos \theta_\mu)}$$



Example: 2p2h



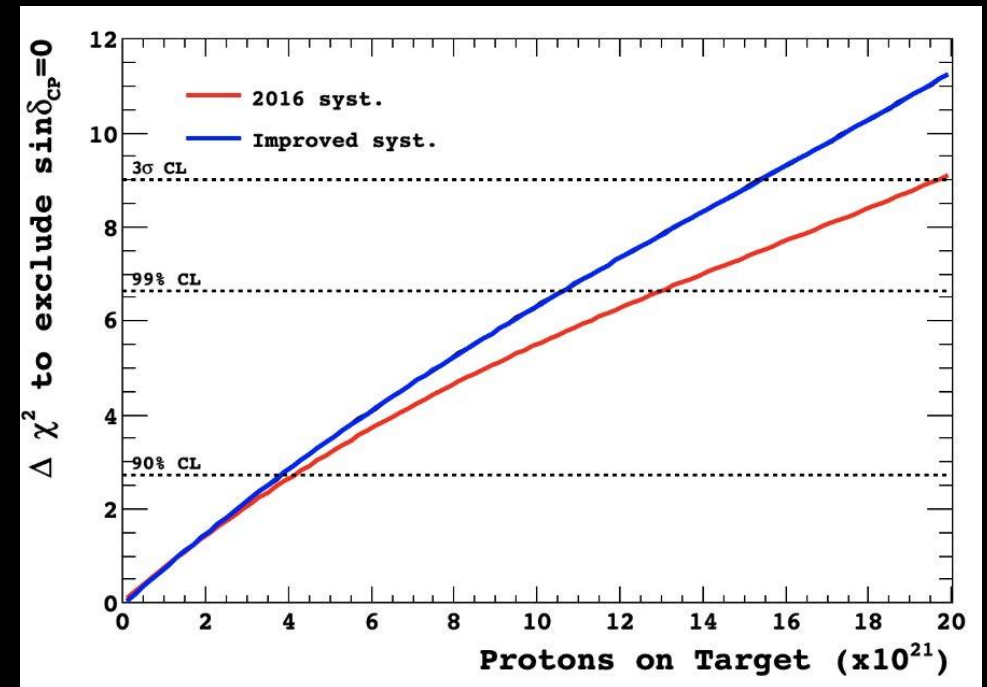
- 2p2h or **two-particle two-hole** interactions can occur when a neutrino interacts with a pair of nucleons, due to short range correlations or meson exchange currents
- Can be simulated using Nieves et al. 2p2h model
- Require uncertainties for:
 - Fraction of pn / nn pairs
 - Normalisation
 - Shape
 - Neutrino energy dependence



Part Two: Upgrade of J-PARC beam and ND280

What do we need now?

- ND280 has been extremely successful in reducing the uncertainties due to neutrino fluxes and cross-section to ~5% level
- T2K has entered **precision phase** – can the data collected at ND280 be improved even further?

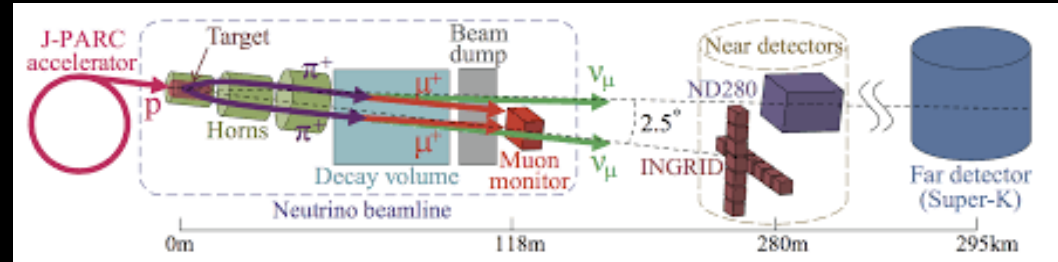
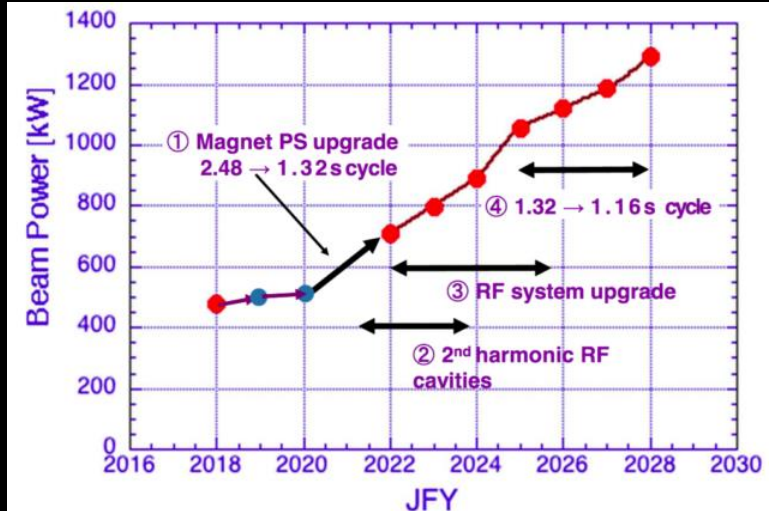


We need to

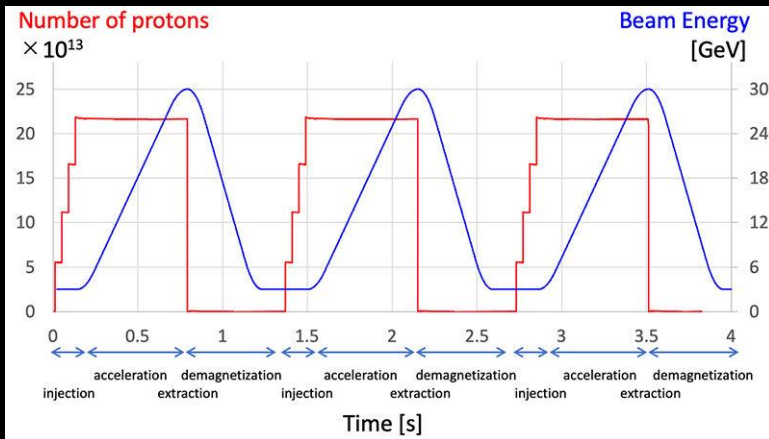
- 1. Maximise statistics AND**
- 2. Reduce systematic uncertainties**

in order to have the best chance of measuring δ_{CP} with precision

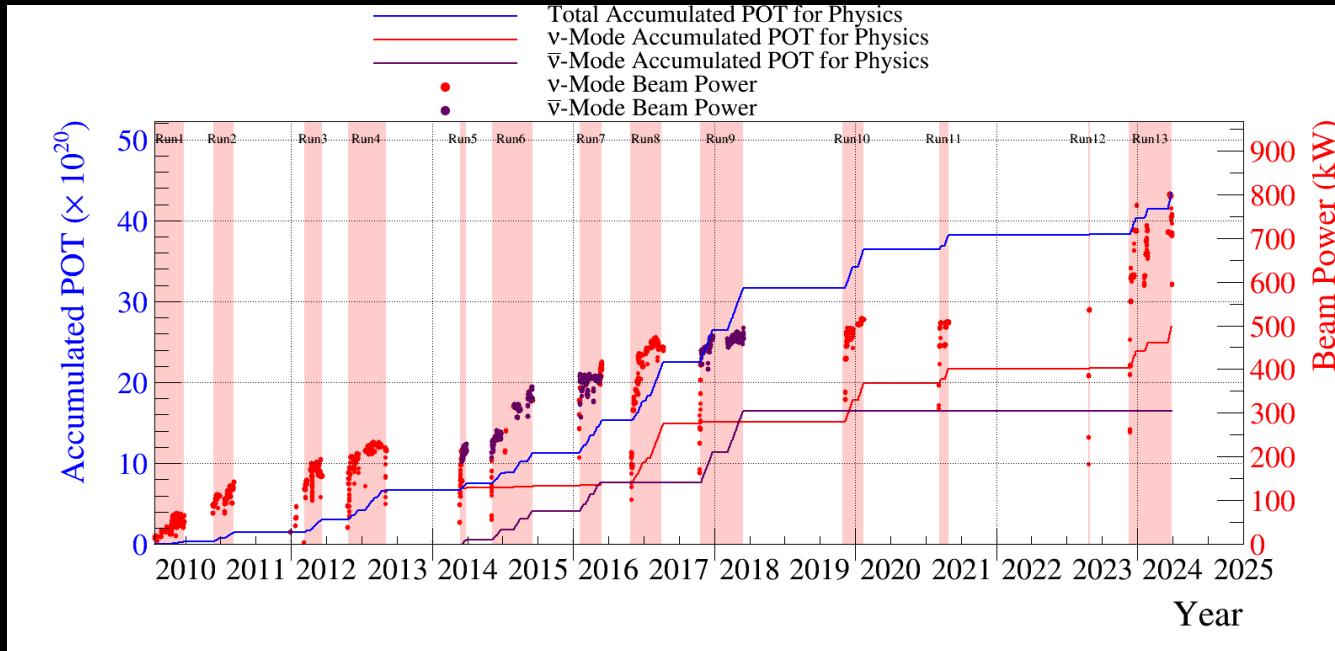
JPARC beam upgrade



- The J-PARC neutrino beam was recently upgraded
- Electromagnetic horns current increased from 250 kA → 320 kA, increasing neutrino flux
- Beam cycle decreased from 2.48s → 1.36s, increasing beam power
- Steady increase of beam power:
 - December 2023 the beam ran stably at 760kW, increased from ~500kW and **surpassing for the first time the original design power of 750kW**
 - June 2024: increased further to 800kW
 - Further upgrades planned to reach 1.3MW



JPARC beam upgrade

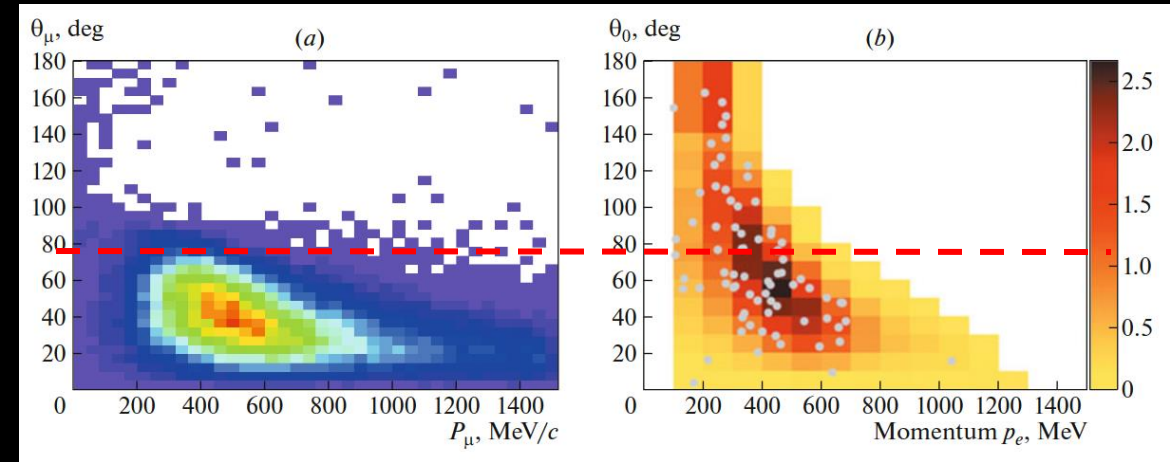


We can now collect the same POT in 4 months as we did in the whole period 2010-2022 !

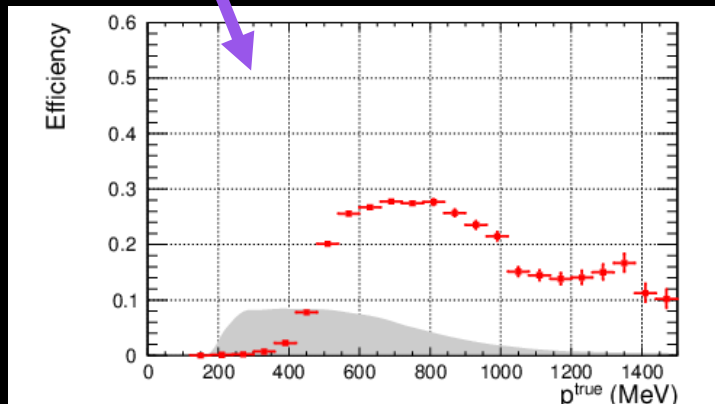
What could be improved for ND280?

Limitations:

- Tracks at high angles (would like an angular acceptance more similar to SK)
- Limited timing information
- No neutron information
- Poor electron/photon separation
- High detection threshold



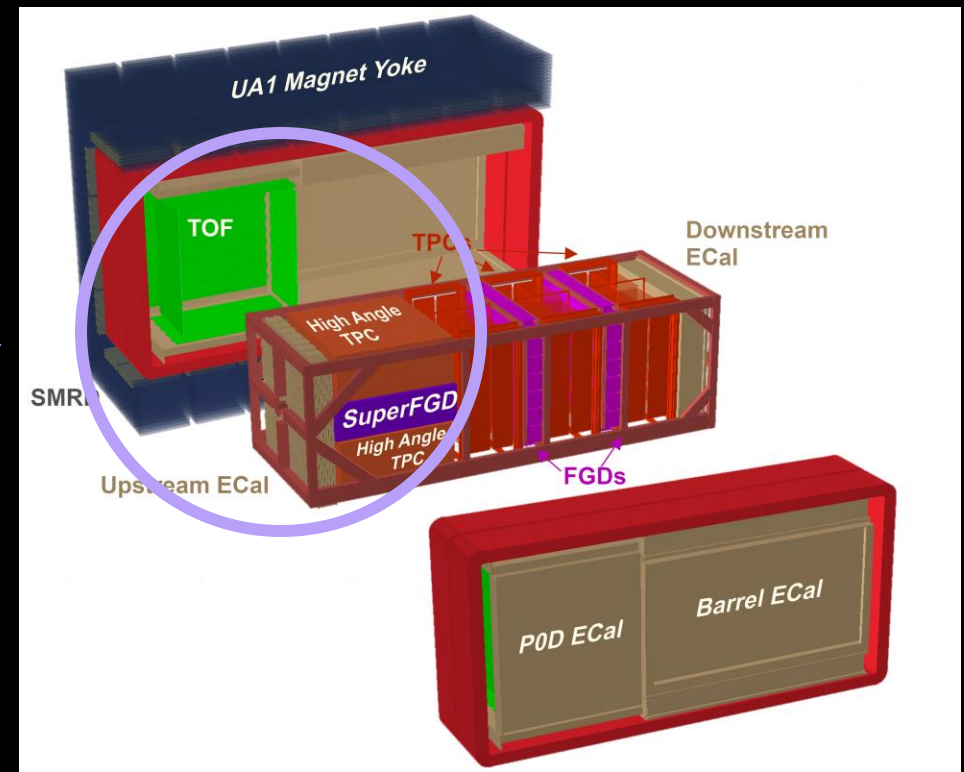
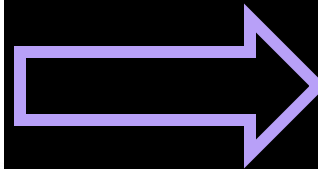
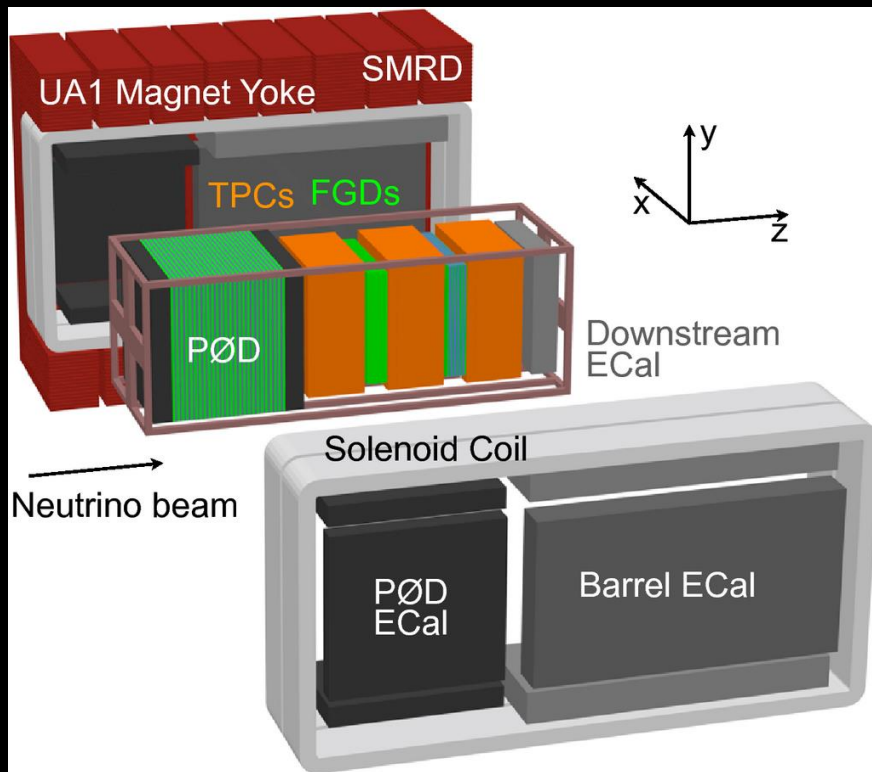
Reconstructed momentum and angle for muons selected at ND280 (left) and electrons selected at SK (right)



Proton reconstruction efficiency in ND280.

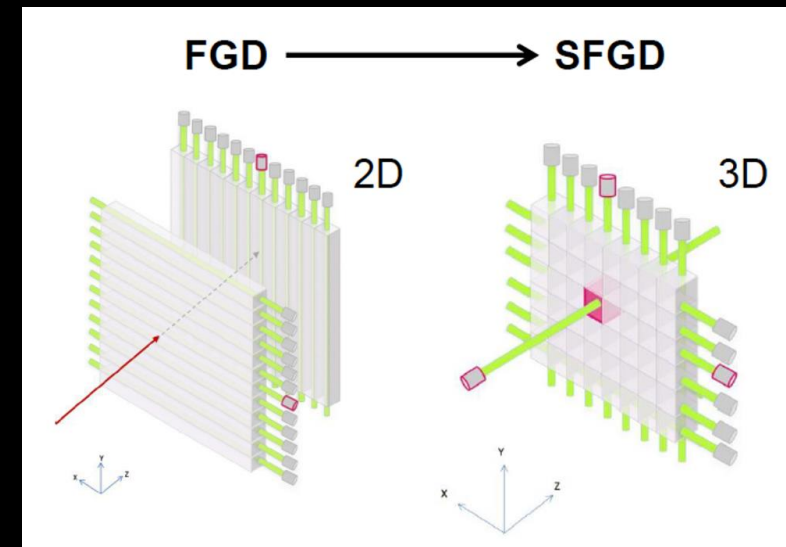
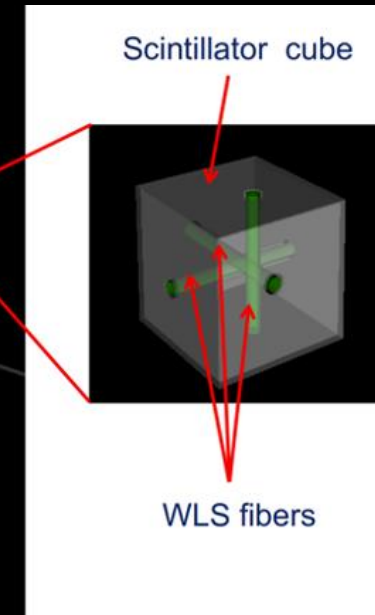
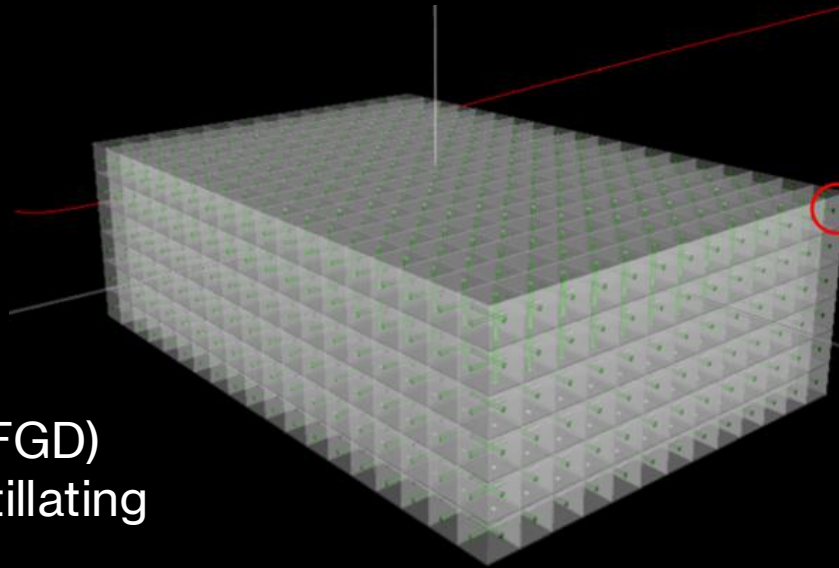
ND280 upgrade

In 2023/24 the POD detector was removed from ND280 and replaced with three new sub-detectors

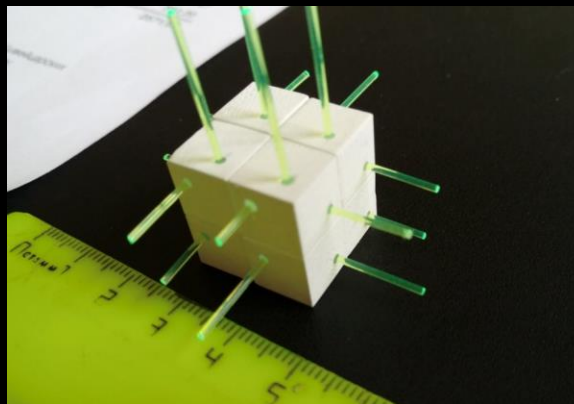
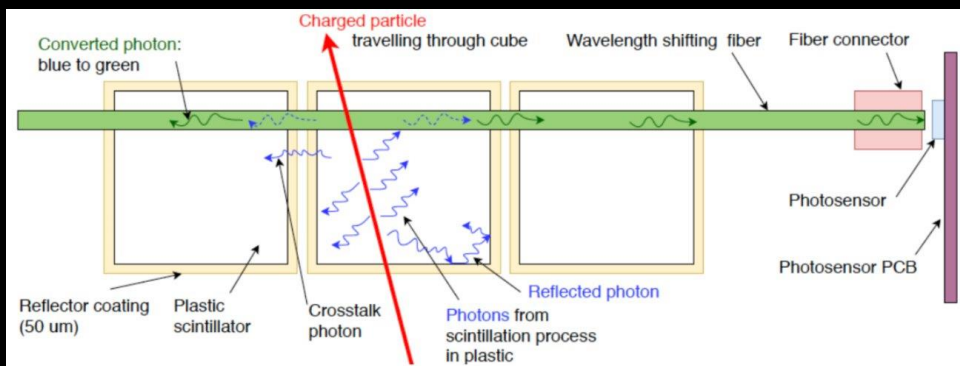


SFGD design

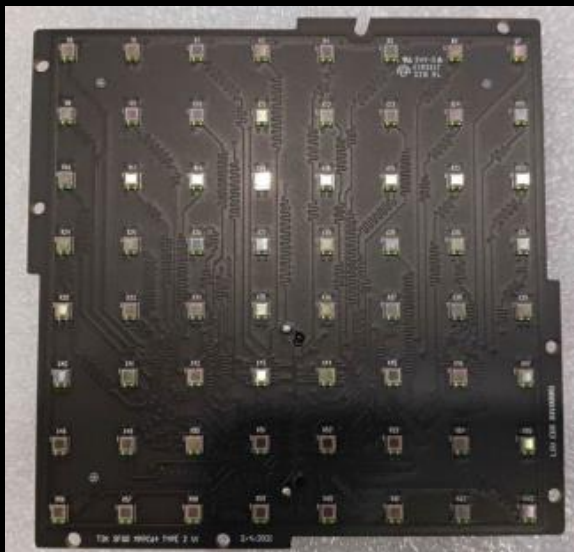
- Super Fine Grained Detector (SFGD) made up of 2 million plastic scintillating cubes
- Total fiducial mass of 2 tonnes
- Each cube is 1x1x1 cm and optically isolated, with three WLS fibers going through each one
- 3D granularity
 - High angle tracks can be reconstructed
 - Short low momentum tracks
- Sub-ns timing resolution
- Two prototypes built and tested in test beams at CERN



SFGD readout

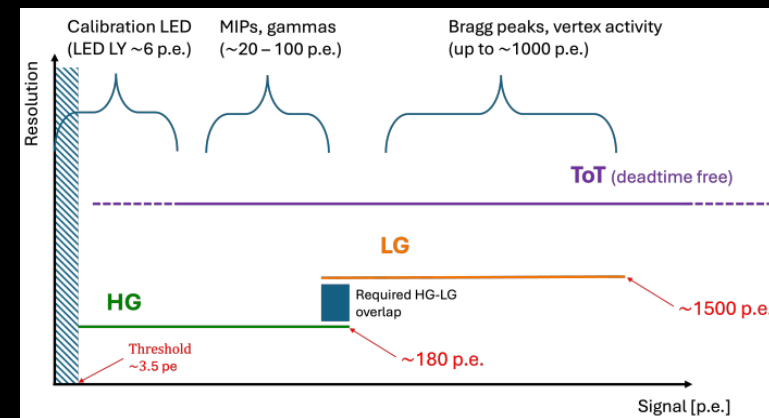


Item	Specification
Effective photosensitive area	1.3 mm x 1.3 mm
Pixel pitch	25 μ m
Number of pixels	2668 pixels
Fill factor	47%
Package type	Surface mount
Breakdown voltage (V_{BR})	53 ± 5 V
Peak sensitivity wavelength	450 nm
Photo detection efficiency	25%
Gain	7.0×10^5
Dark count	70 kcps (typ.)
Crosstalk probability	1%



PCB with 8x8 array of surface mounted MPPCs

- Three WLS fibres through each SFGD cube requires **56k readout channels**
- Single ended readout of Hamamatsu Multi-Pixel Photon Counters (MPPC)
 - High dynamic range from 3.5 to 2000 PE



(i) Support system assembly



(ii) First cube layer assembly



(iii) All 56 layers assembled



(iv) Stop panels removed



(v) Box closure



(vi) Transfer to new support



(vii) Horizontal fibers assembly



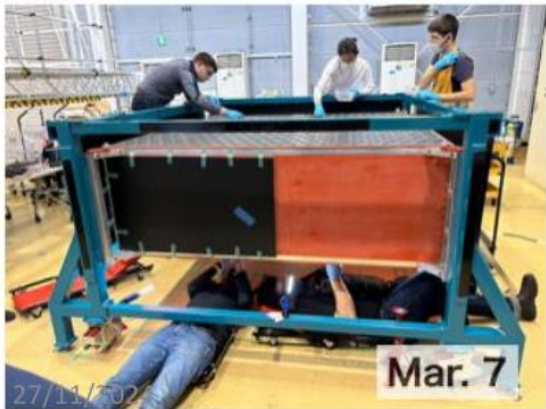
(viii) Wall MPPCs assembly



(ix) Vertical fibers assembly



(x) Top MPPCs assembly



(xi) LED calib. modules assembly

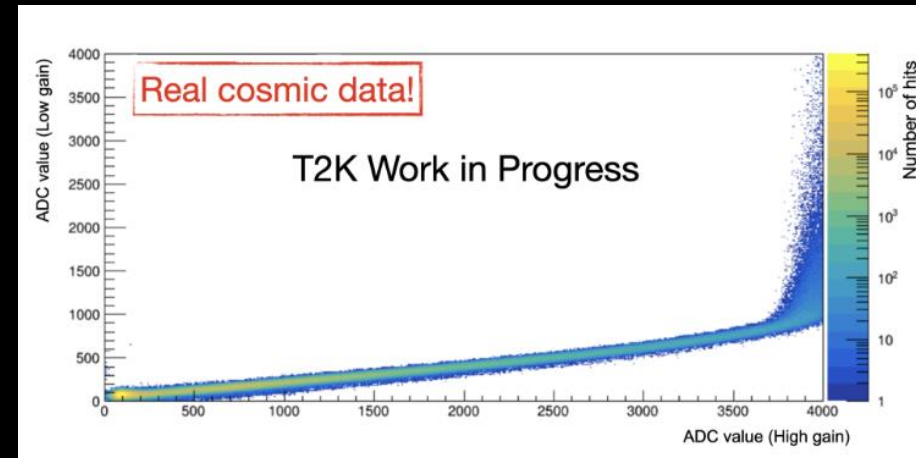
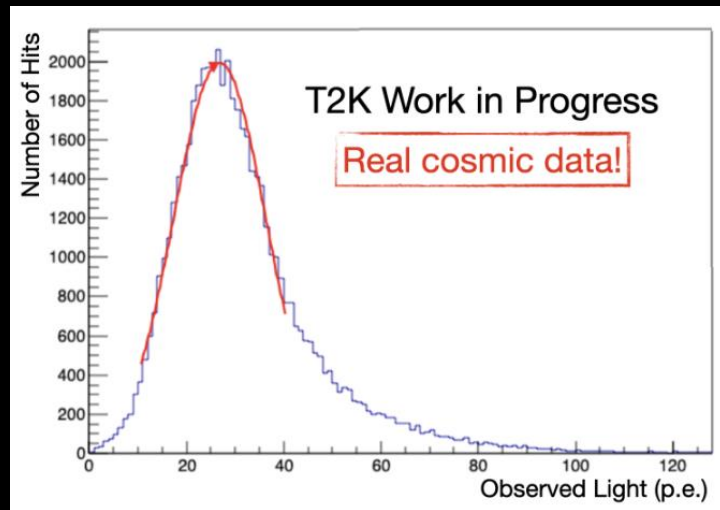
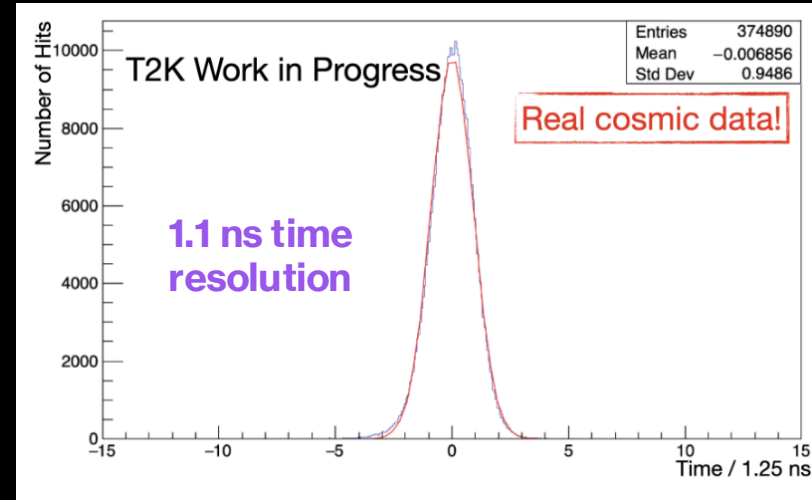
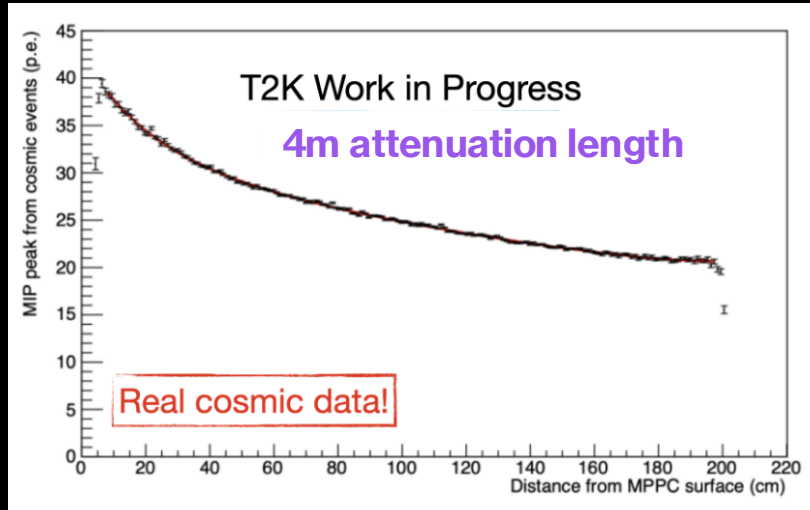


(xii) Light barrier/cables assembly



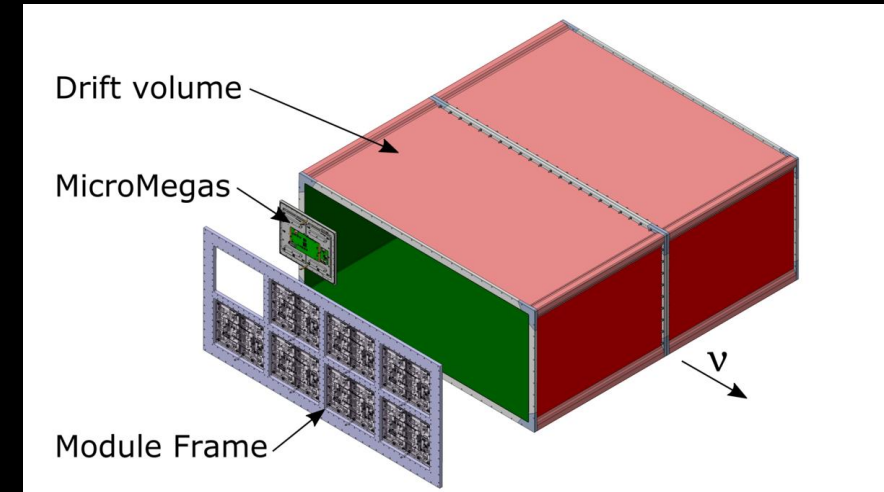
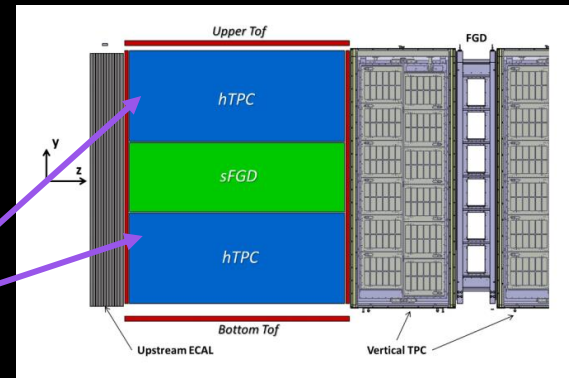
25

Detector performance with cosmics



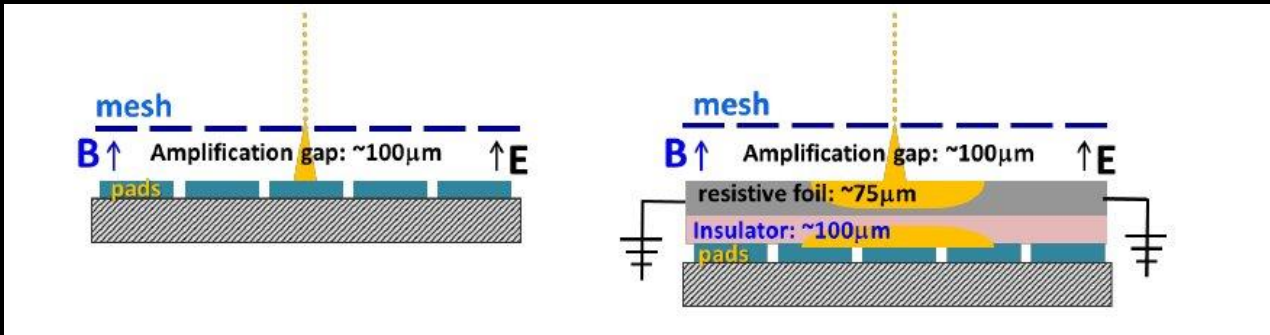
High angle TPCs

- Two new time projection chamber detectors
 - Dimensions $2.0 \times 0.8 \times 1.8$ m
 - Gaseous mix, 95% Ar
 - 275 V/cm electric field
 - Two Cu field cages joined by a central cathode producing a 90 cm drift distance
- Similar to the present ND280 TPCs, they provide high granularity 3D track reconstruction, charge and momentum measurement and PID
- Instead of downstream, these are above and below SFGD, providing information about charged particles emitted at high angles

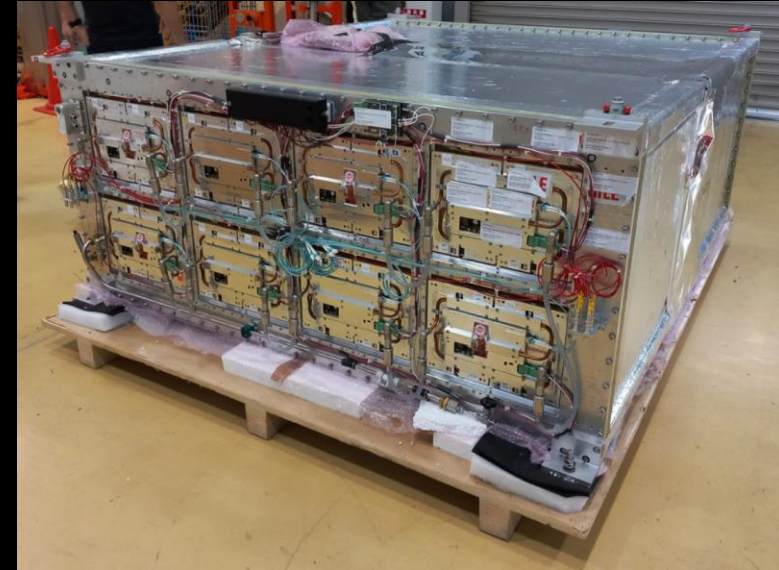


Major changes to design: novel readout technology and new field cage design to minimise dead space + position of TPCs designed for high angle tracks

Resistive Micromegas readout



Signals read out by 8 Micromegas organised in a 4x2 readout plane

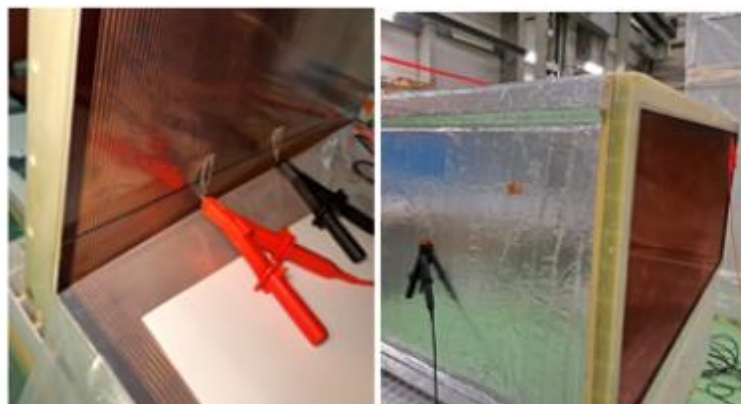
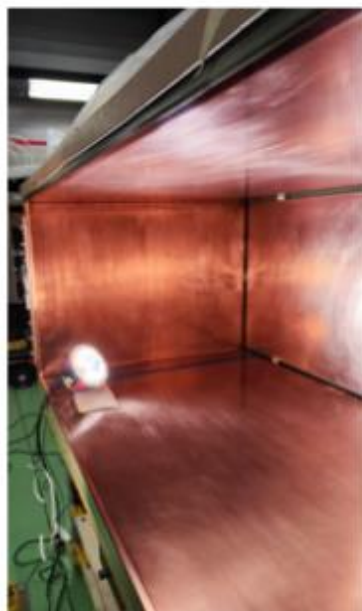


- Novel technology - Encapsulated Resistive Anode Micromegas (ERAM)
- Additional resistive layer spreads charge over several readout pads on anode plane
- Results in a much better point resolution
- Resistive layer additionally prevents sparks, allowing higher gain to be used and simplified front end electronics

Inner cage surfaces polishing

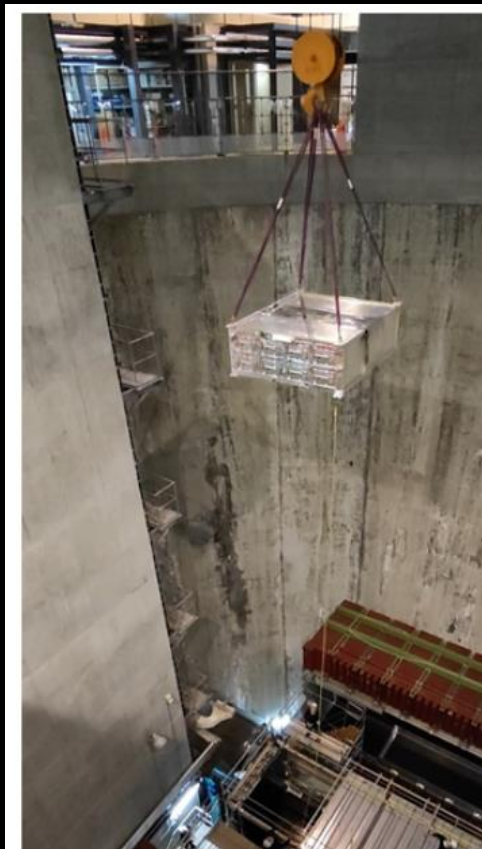


Checking grooves for o-ring and for charge labyrinth on cathode flanges



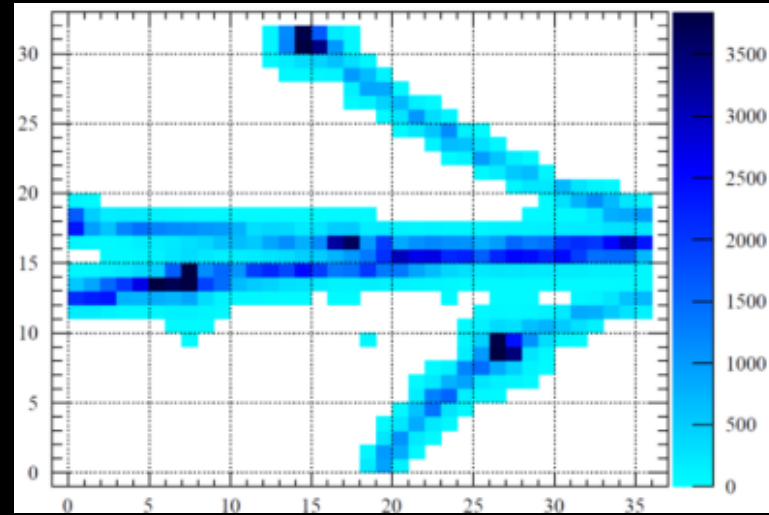
Measuring strip-strip and strip-shield insulation at high voltage

ND280 after lowering of top HATPC
2024.4.25

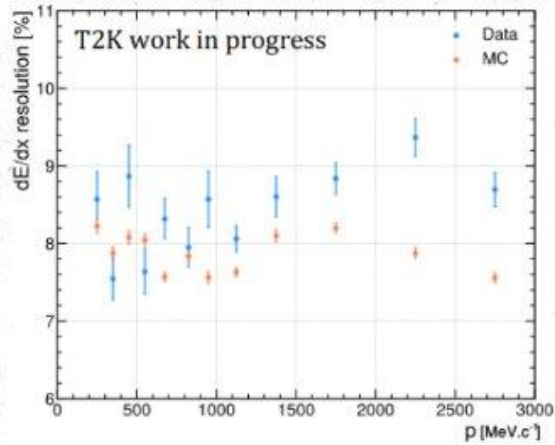


Lowering bottom HATPC
2023.9.8

HATPC performance

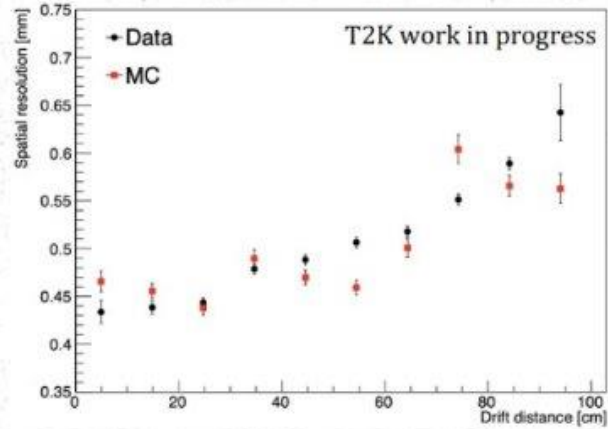


dE/dx resolution



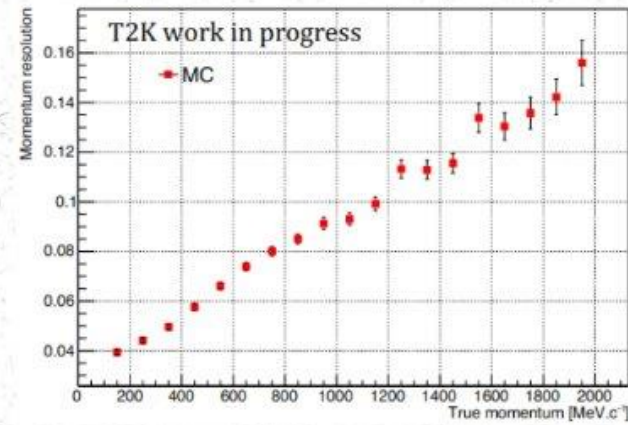
- dE/dx resolution **better than 10 %**.
- Behaves as expected.

Spatial resolution



- Spatial resolution $\approx 500 \mu\text{m}$.
- In reasonable agreement with MC.

Momentum resolution



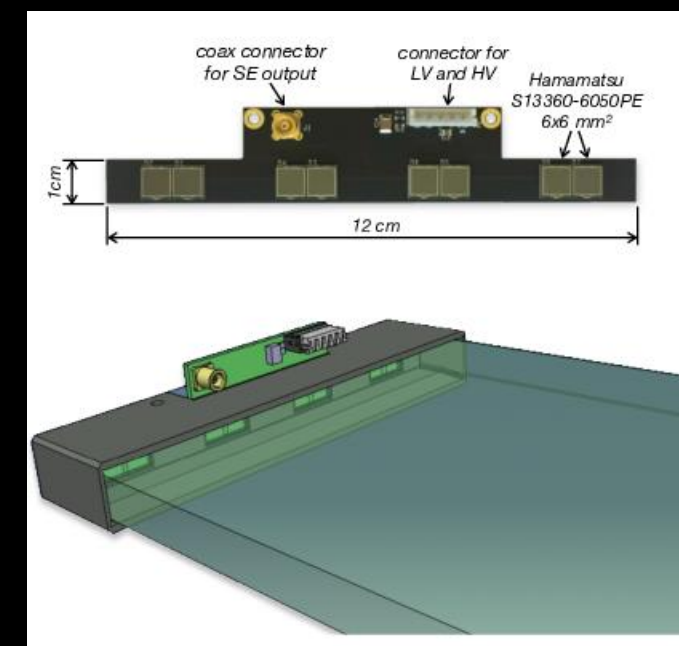
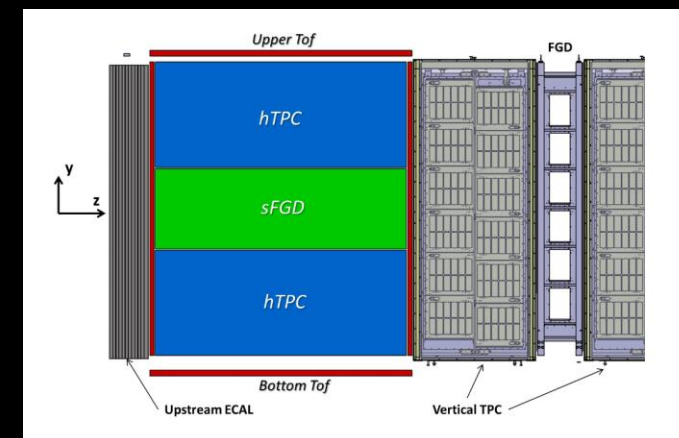
- Momentum resolution **better than 10%** for vertical muons with momenta $< 1.2 \text{ GeV}/c$ and $L > 600 \text{ mm}$.

Time of flight (ToF) detectors

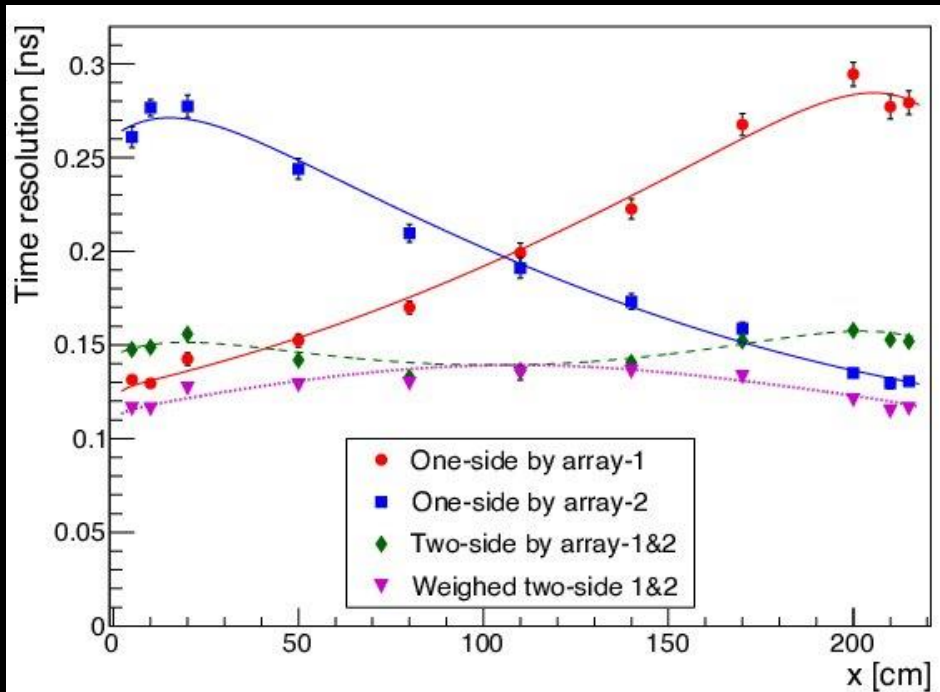
- Six time of flight modules surround the SFGD and HATPC volume
- Each plane is 2.3 by 2.5 m² and is made up of 20 1cm thick scintillator bars with high light output, stable attenuation length and fast timing
- Scintillation light propagates through the bars and is read out by double ended SiPM readout



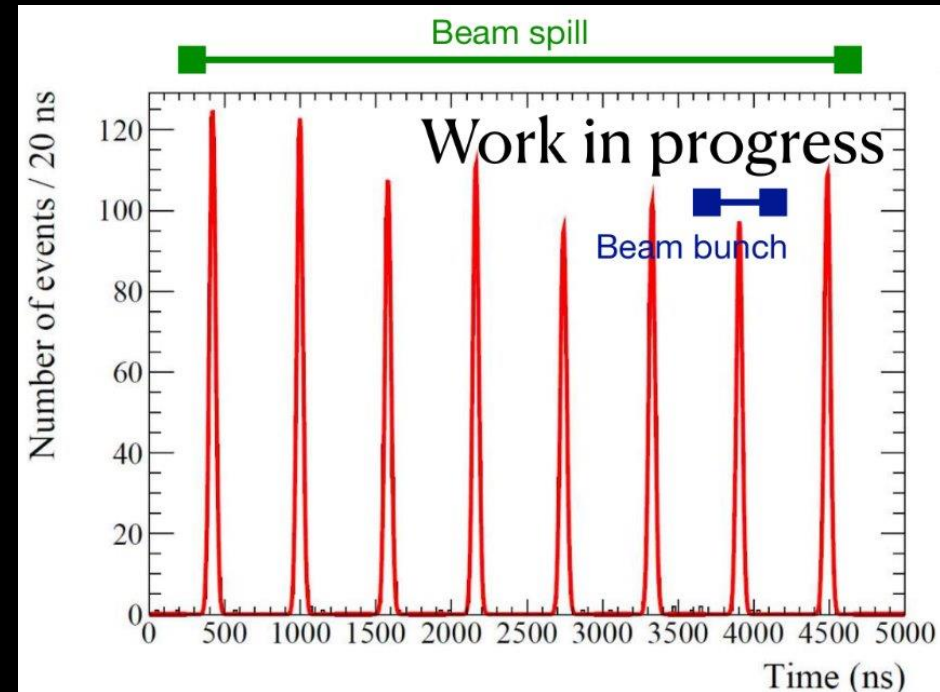
ToF modules during commissioning at CERN



ToF timing performance



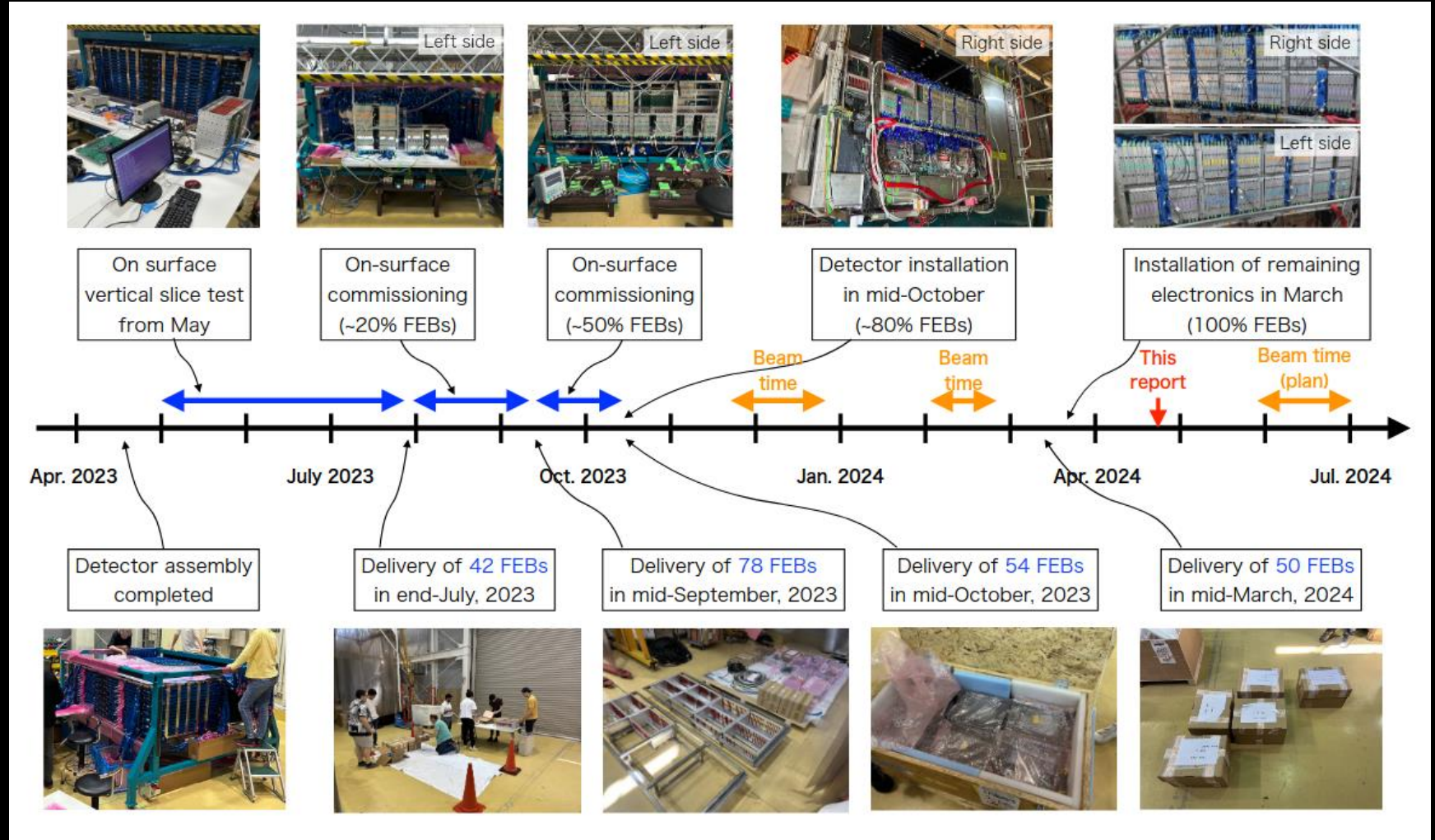
Timing resolution of 150 ps is achieved



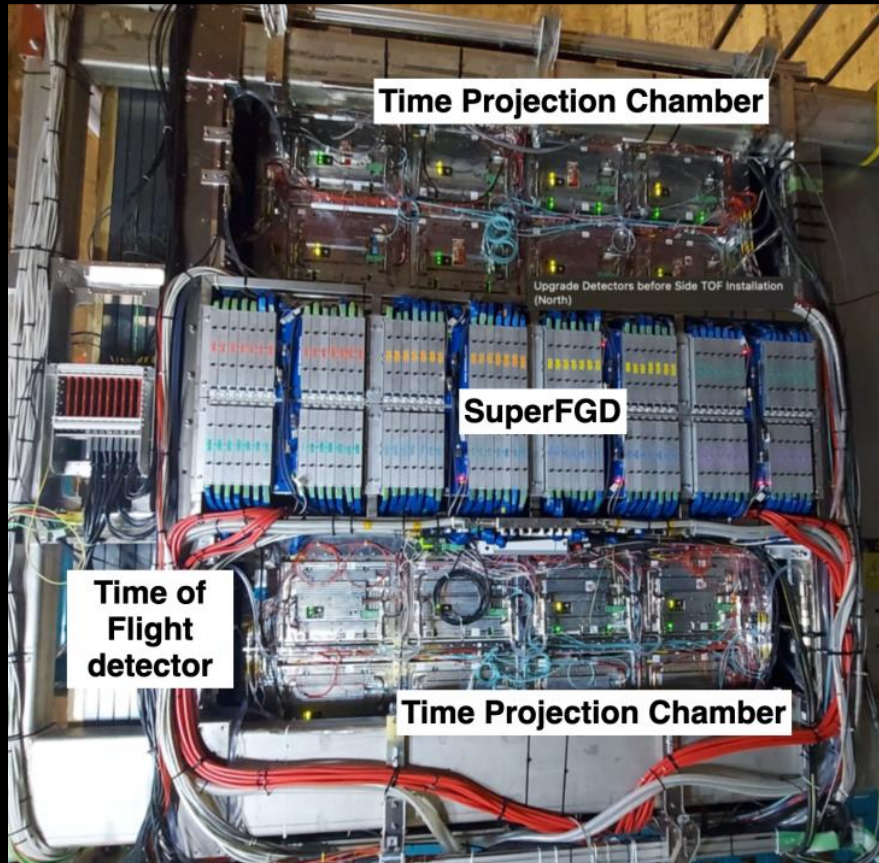
Beam structure clearly captured by ToF

Part Three: First data with ND280 upgrade!

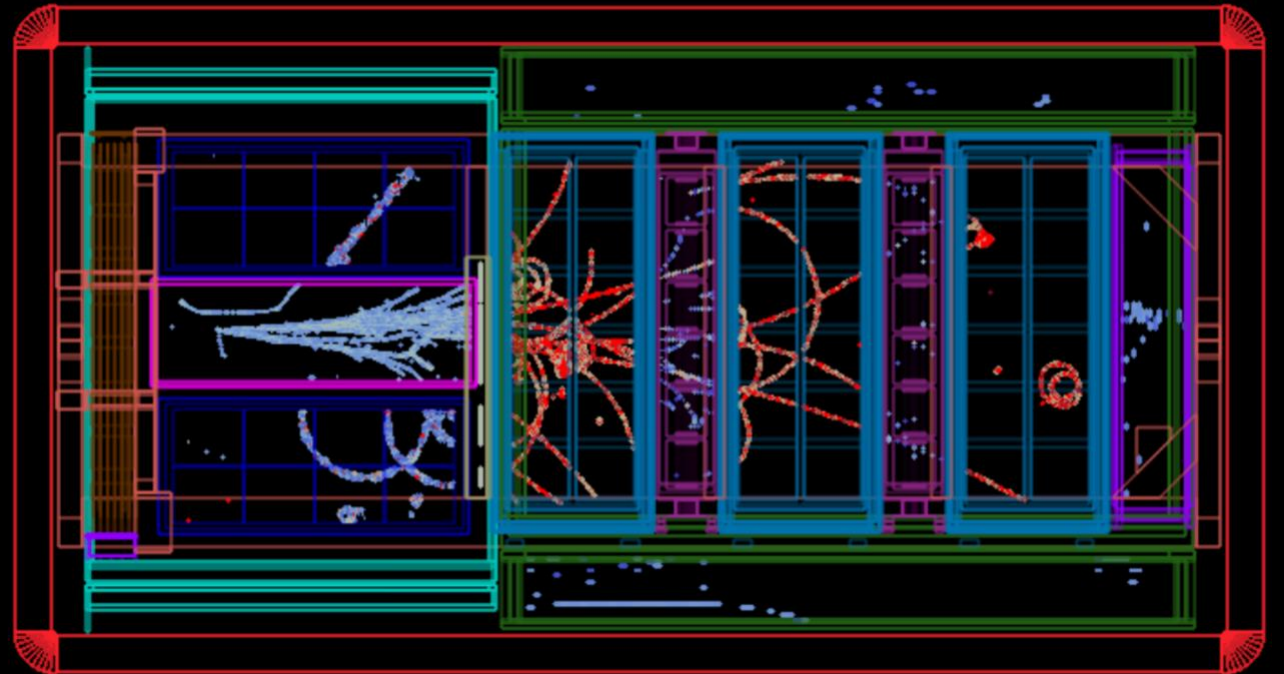
Building ND280 upgrade



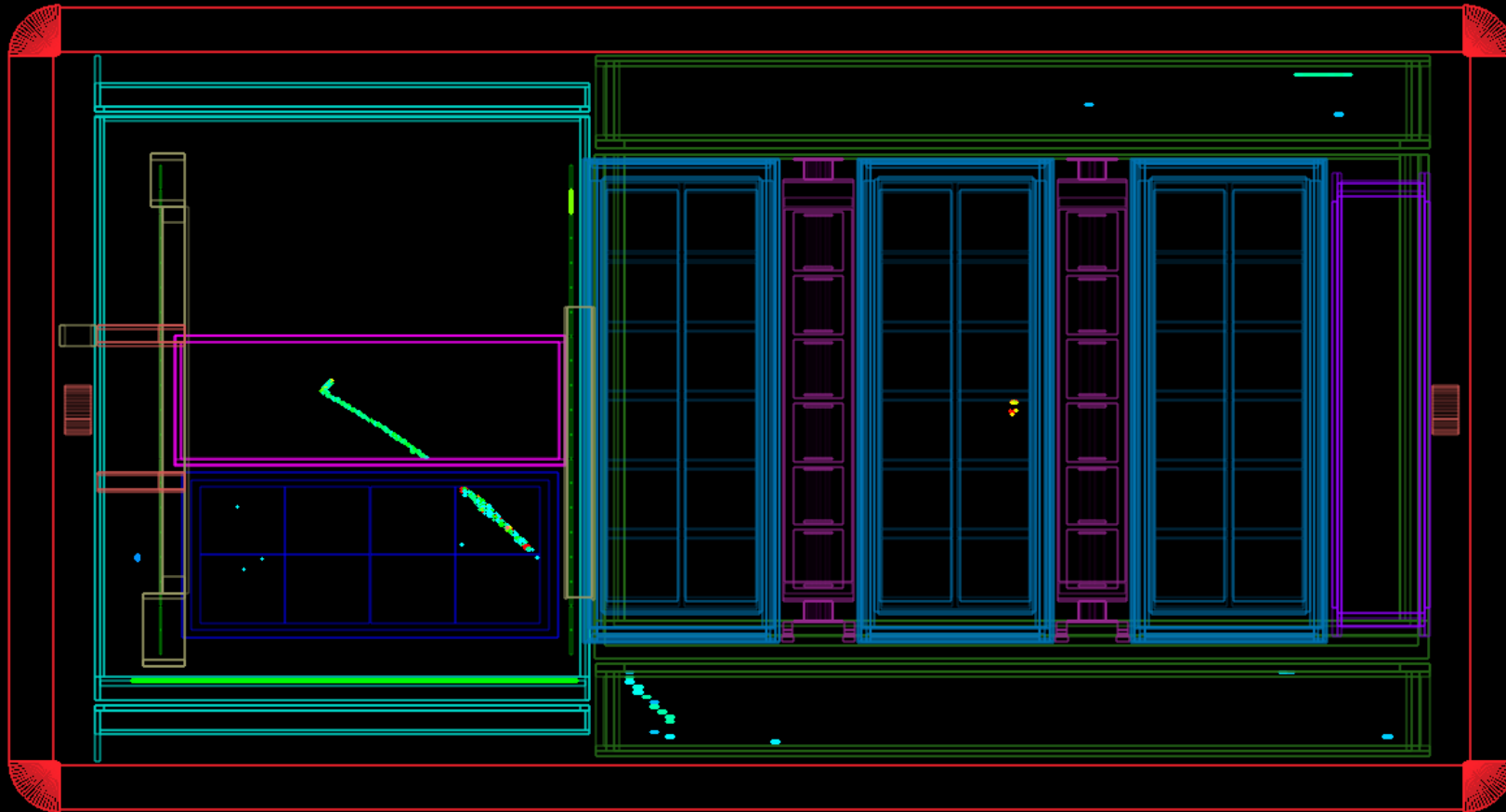
First data taking



- Technical runs taken in Nov/Dec 23 and Feb 24 with SFGD, bottom HAT and first TOF panels
- Installation of all sub-detectors completed with final two ToF panels in May 2024
- **First physics run with completed upgrade in June 2024**

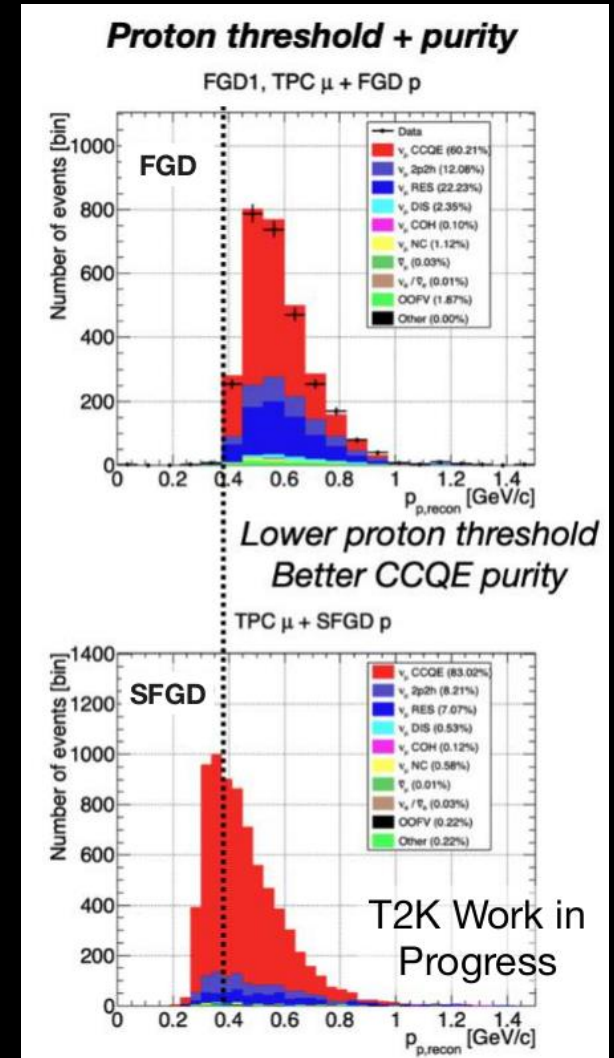
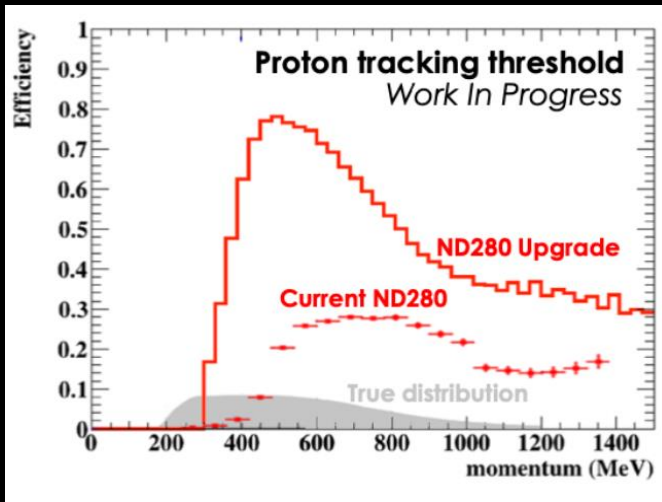


Event display

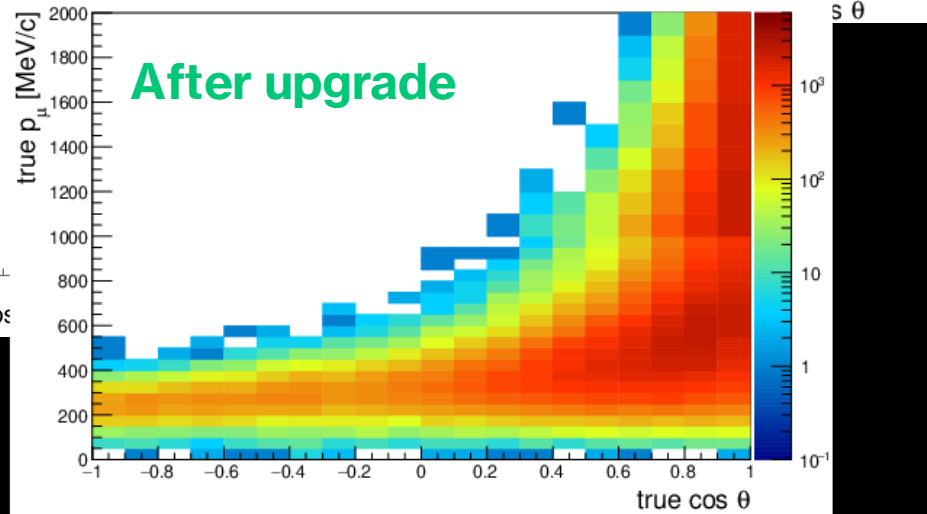
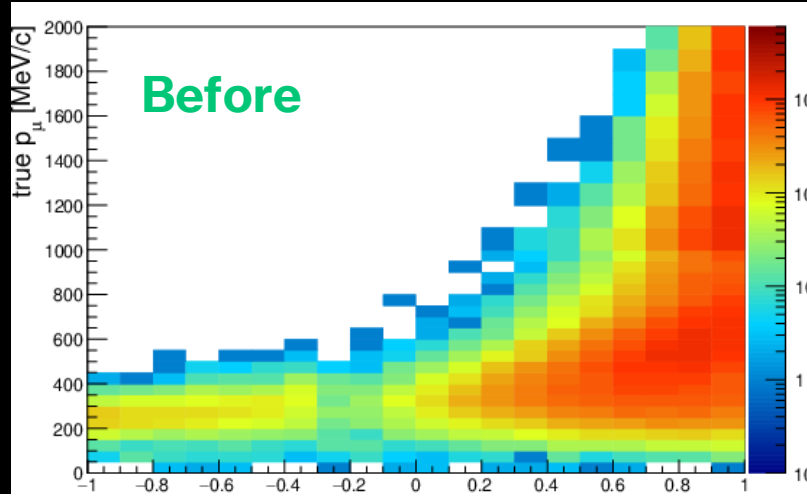
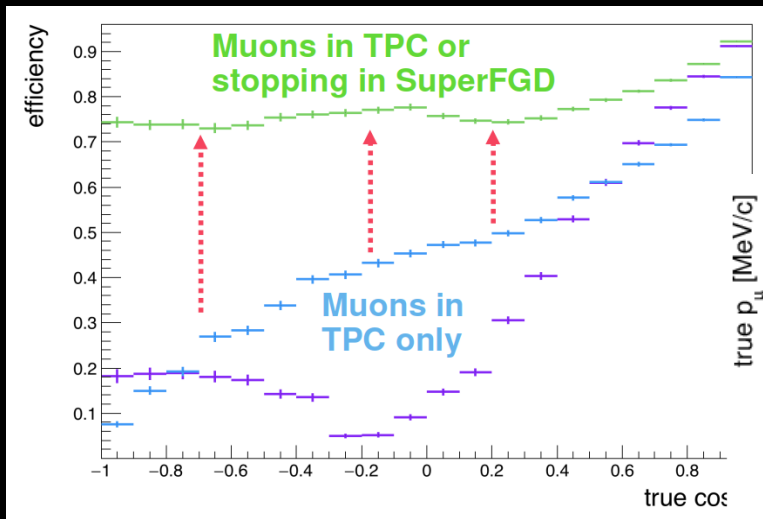


Lower momentum threshold

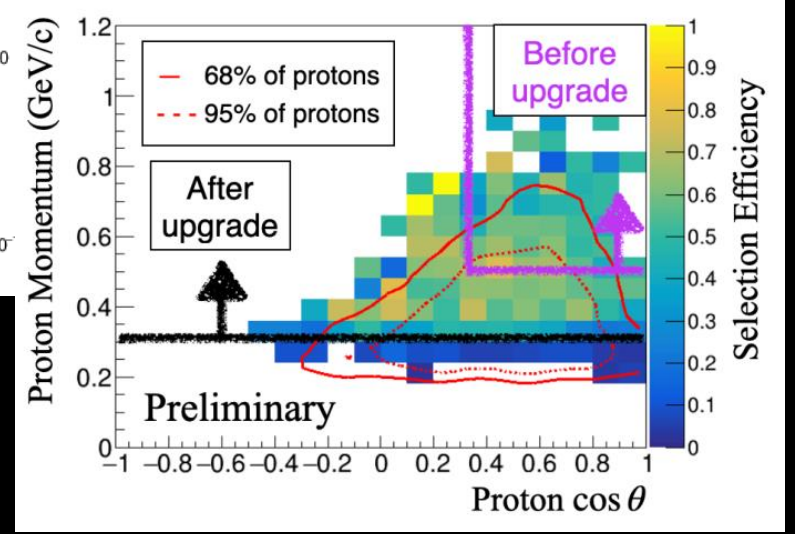
- SFGD has reduced momentum threshold for pions and protons
- Proton threshold reduced to approx 300 MeV
- Additionally, a higher efficiency over the whole momentum range



Better angular acceptance



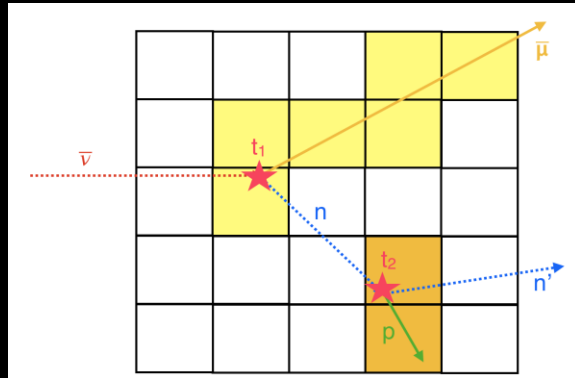
Proton distribution



4 π acceptance for tracks due to HATPC detectors

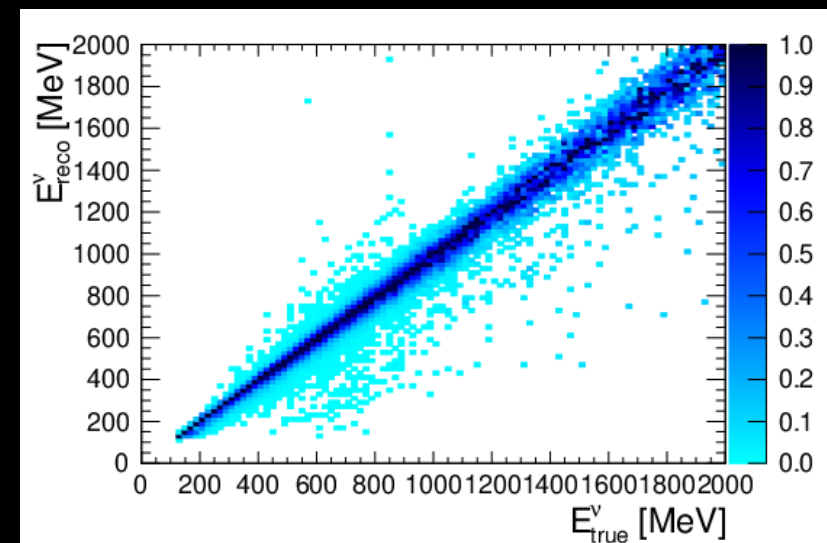
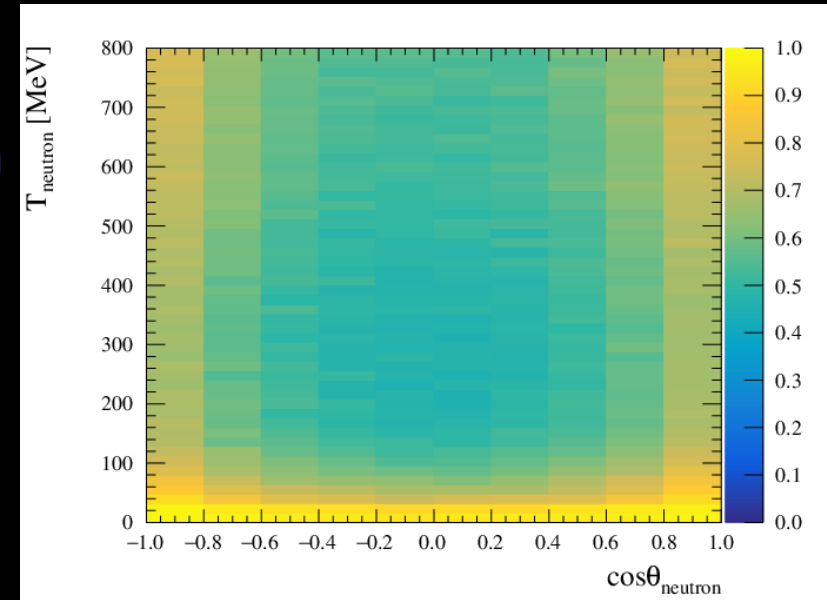
Neutron detection

Efficiency
of neutron
detection



Time of flight between neutrino vertex and secondary vertex can be used to reconstruct neutrons in the SFGD

- This is especially important for anti-neutrino channels, where the reconstruction of the neutron will massively increase resolution of reconstructing the neutrino energy
- Approx 50% tagging efficiency expected

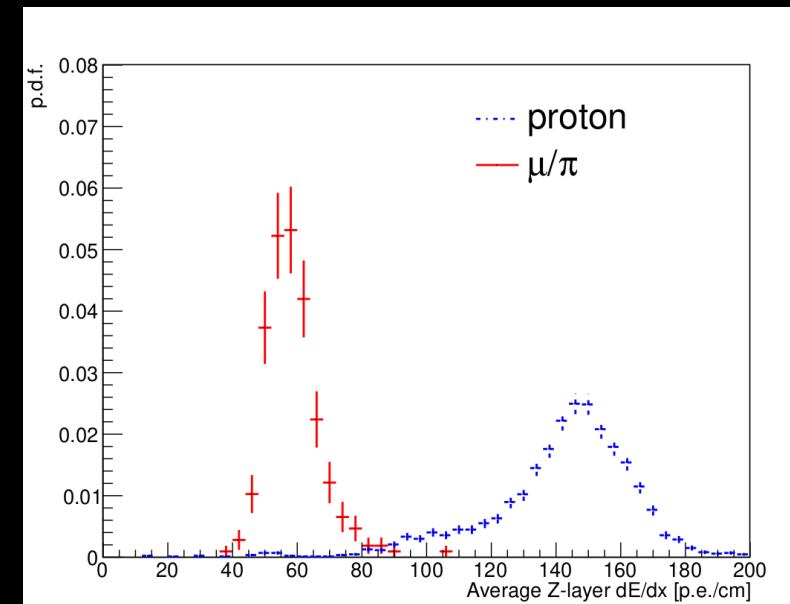
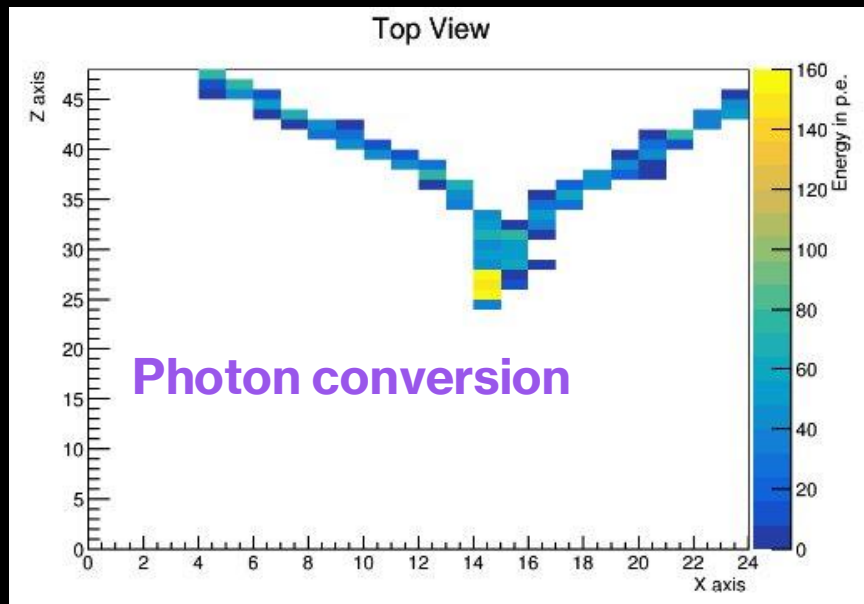


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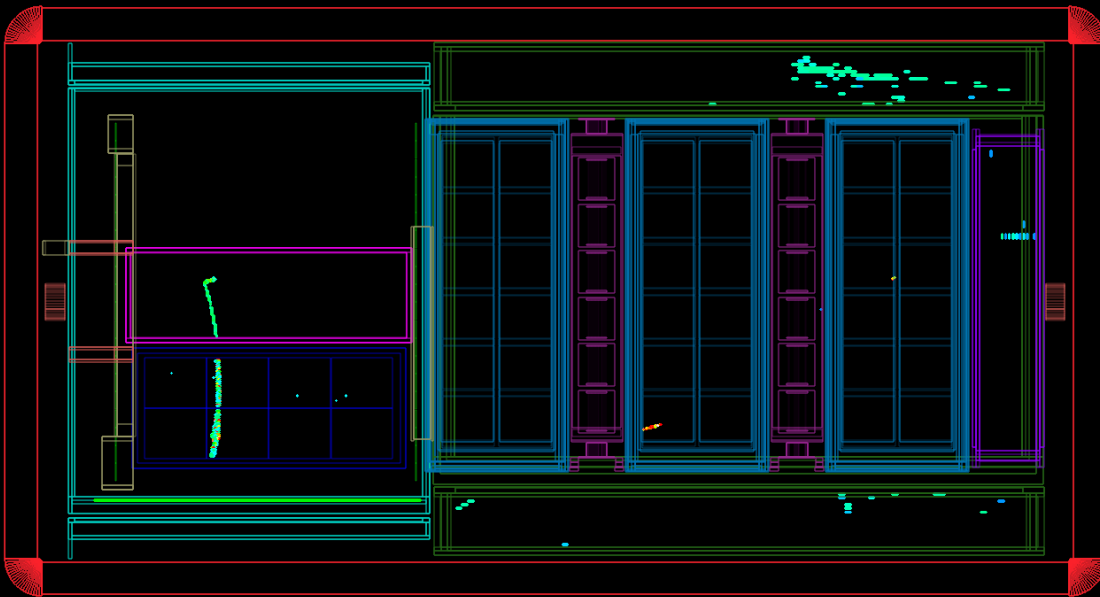
Better PID

- Photons from π^0 production are a dominant background to electron neutrino selections in ND280
- SFGD high granularity has much improved electron/gamma identification

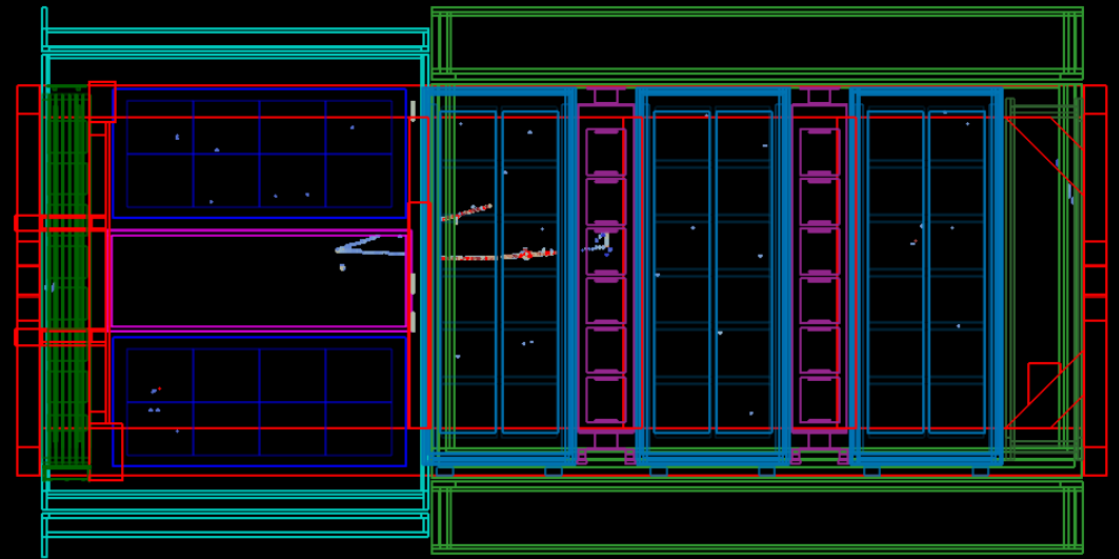
- Excellent proton /MIP separation



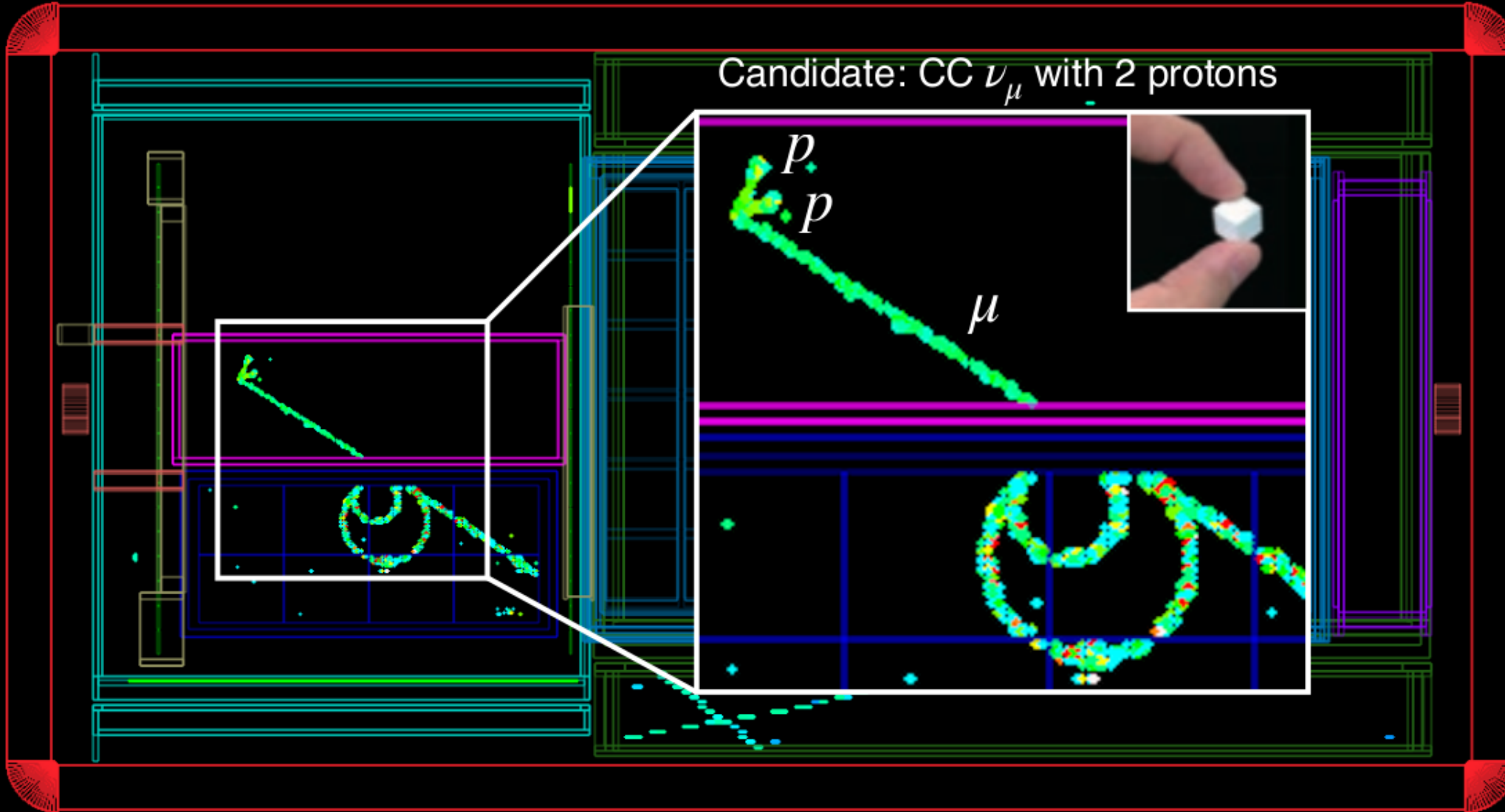
Interesting events



Candidate event with very **high angle muon** and forward going proton (would be hard to measure pre-upgrade)



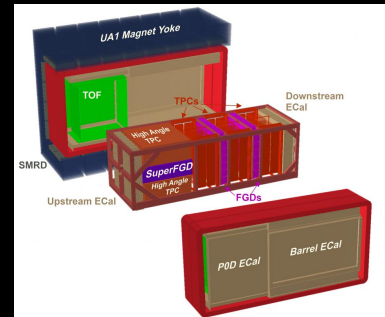
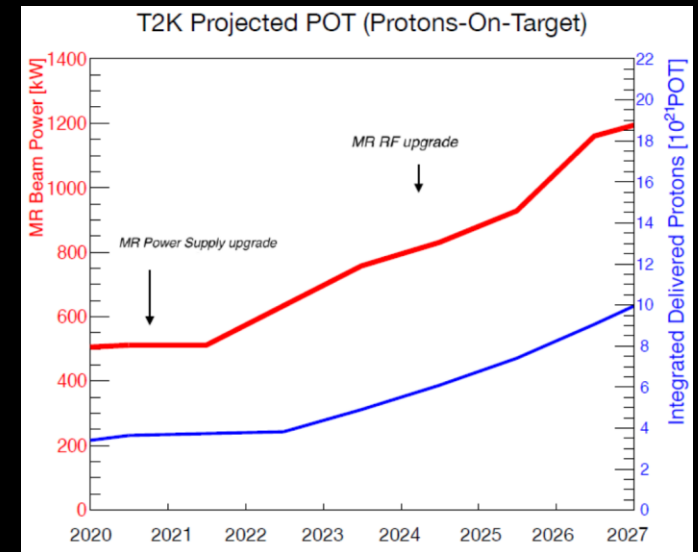
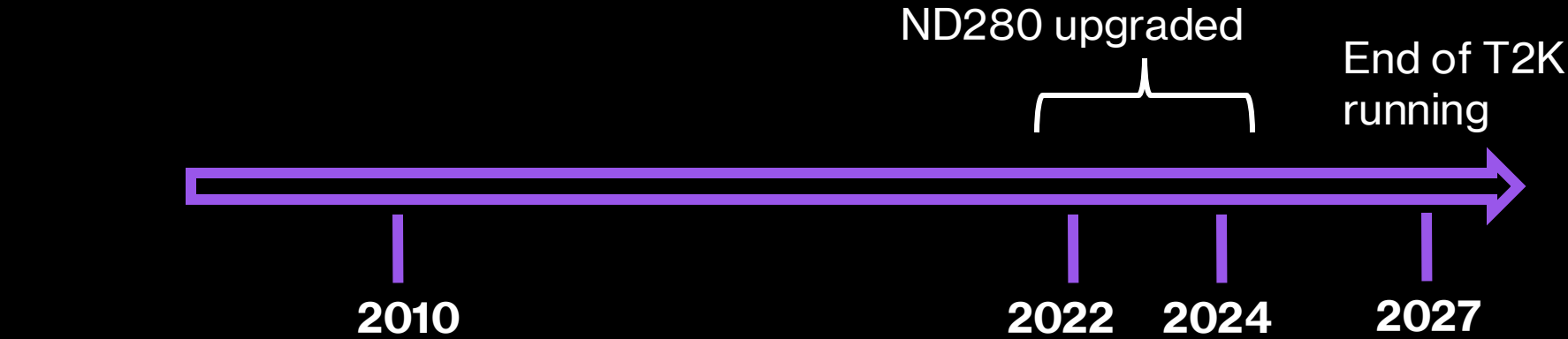
Candidate event with **neutron** emitted from neutrino vertex



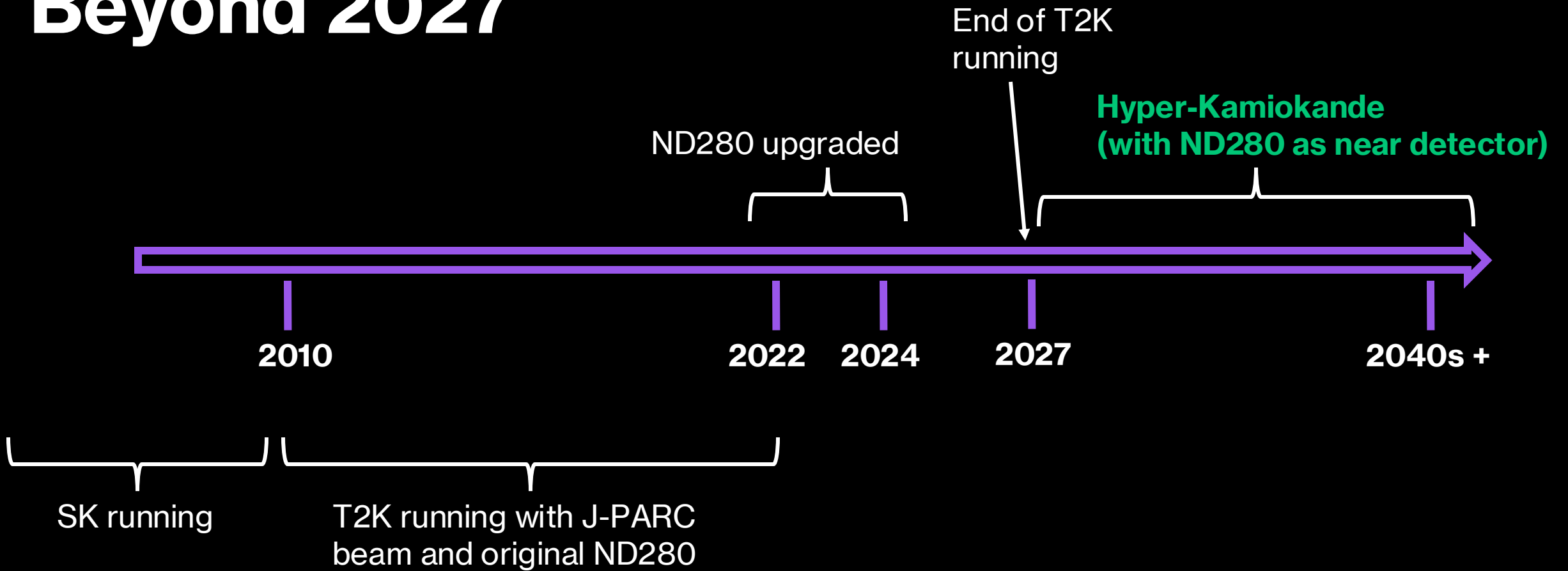
Part Four: Future plans for ND280?

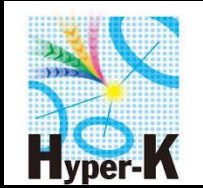
Timeline

- T2K will continue to run until 2027
- Beam power to continue to increase up to 1.3 MW
- Plan to collect 10×10^{21} POT, for a 3 sigma measurement of δ_{CP}



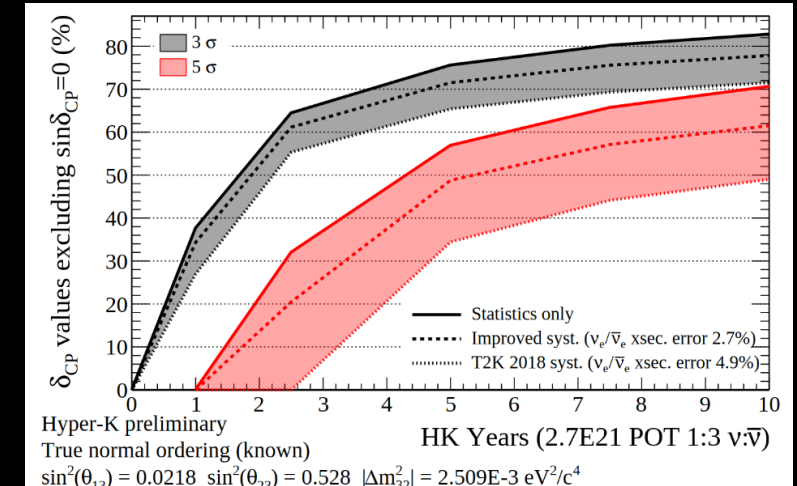
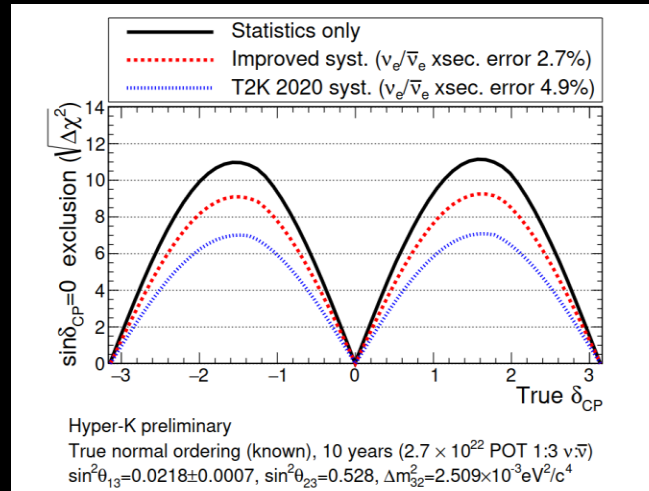
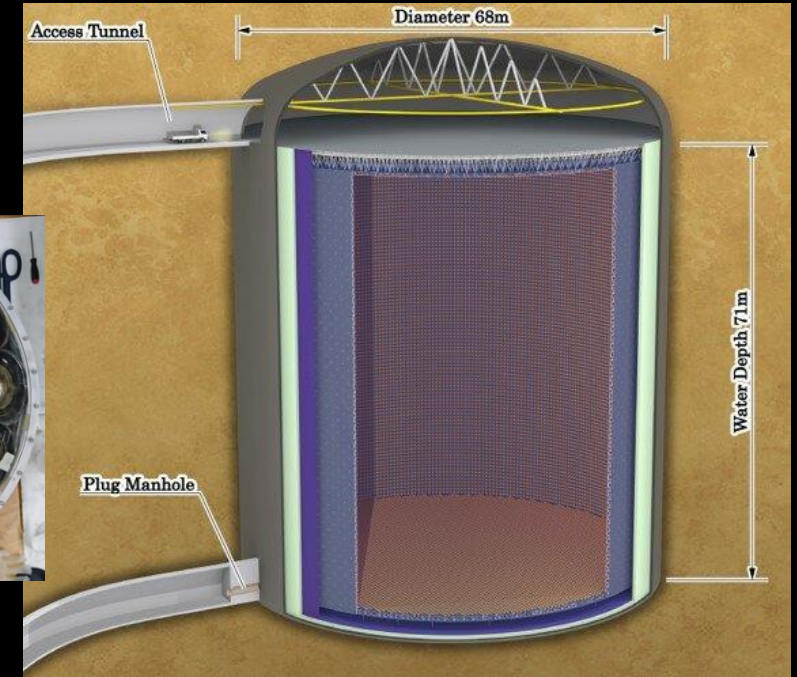
Beyond 2027





Hyper-Kamiokande

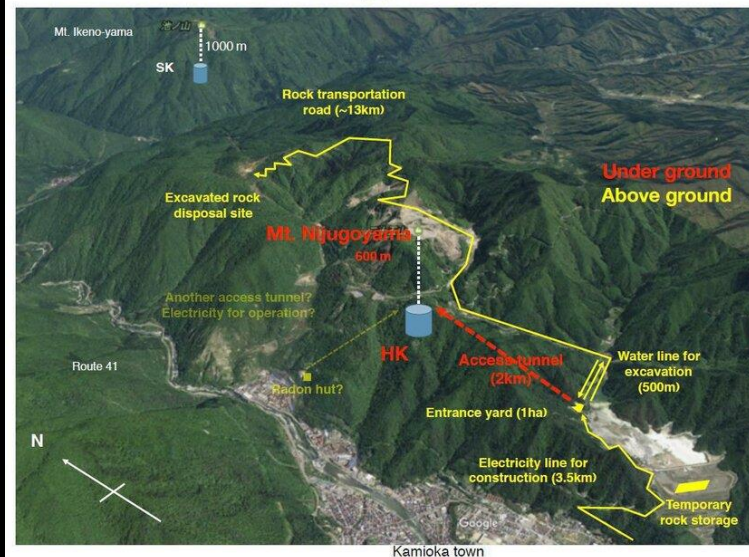
- Next generation water Cherenkov detector currently under construction, with fiducial volume approximately 8 time larger than Super-K
- Liverpool strongly involved in the calibration system for HK and in particular the light injection system
- Expected to take data from 2027 onwards
- Will use ND280 as the near detector



Hyper-Kamiokande is currently under construction

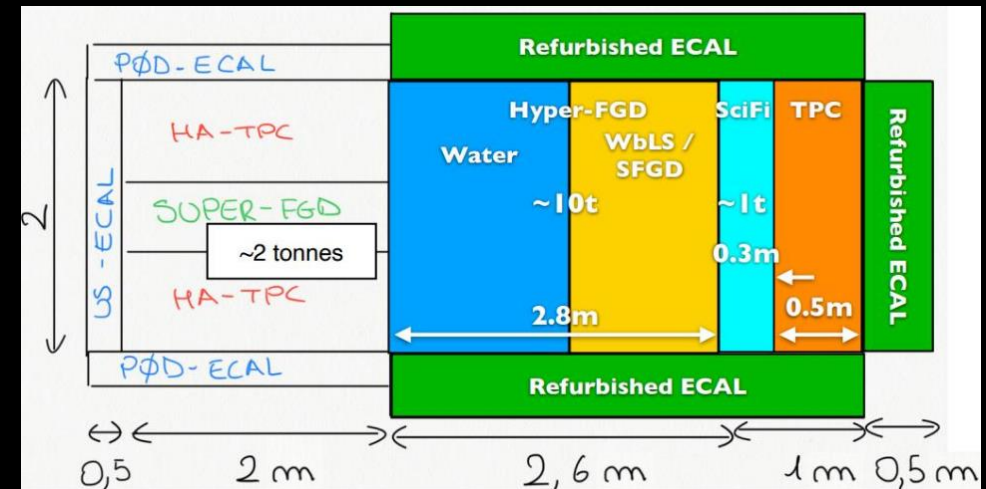
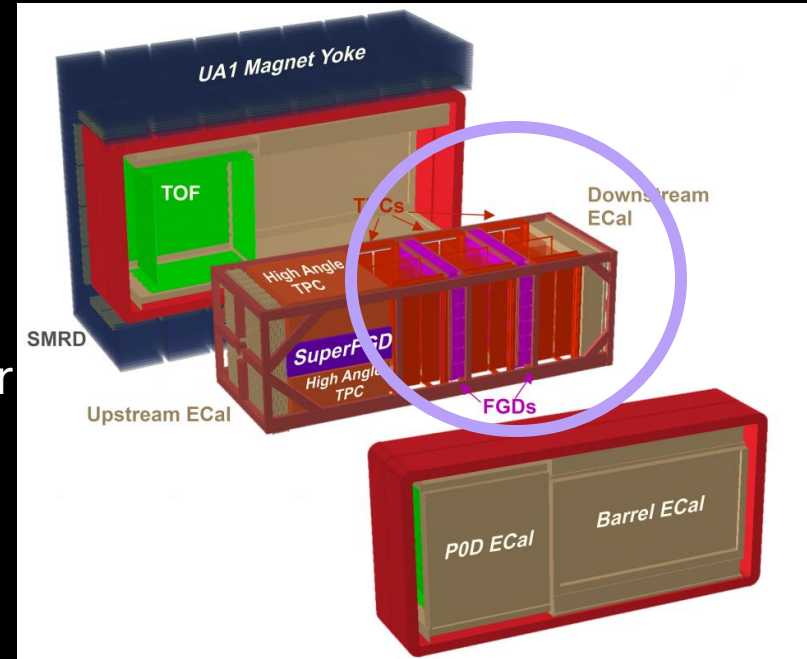


Overview of the Hyper-K constructions



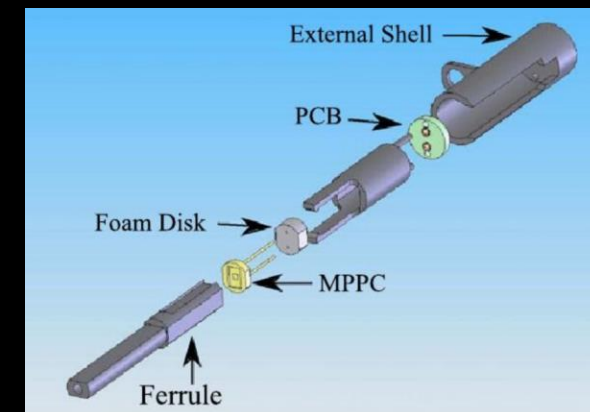
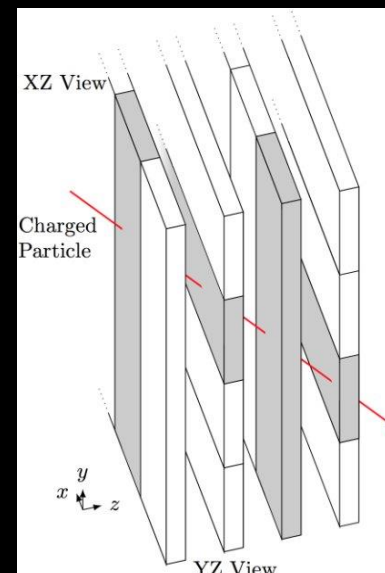
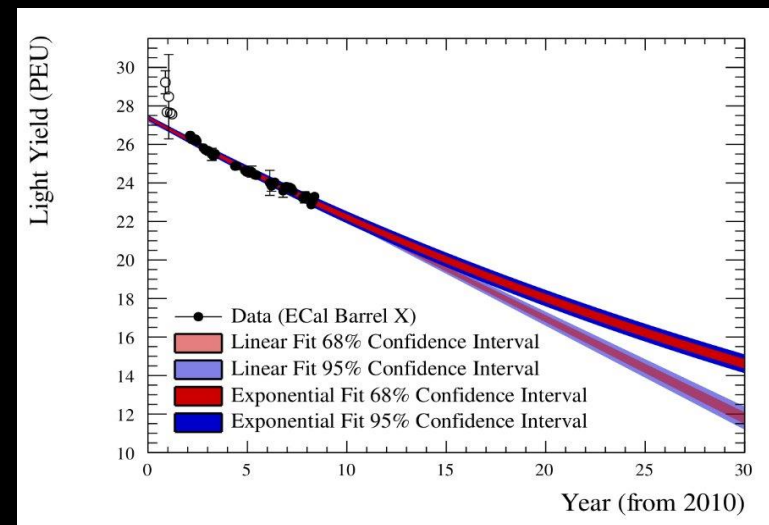
ND280++?

- Proposal to carry out additional upgrades to ND280 for the HK era
- This would primarily involve the "tracker" part of the detector which was not upgraded in the last few years
- Goal of high granularity and large mass of hydrogen/water
- R&D ongoing to develop new technology
 - "hyper"-FGD (larger volume of SFGD)
 - Water based liquid scintillator in segmented cubes
 - Scintillating fibres



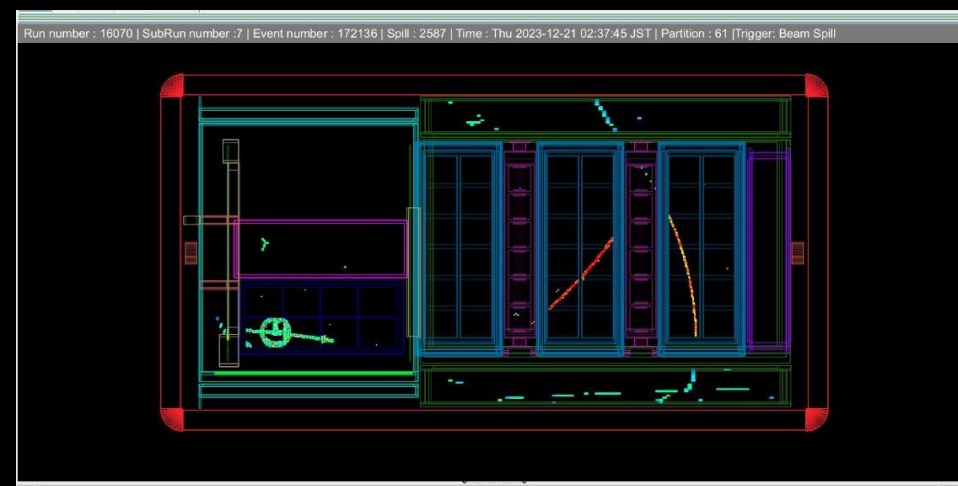
ND280 ECal refurb

- The Electromagnetic Calorimeter part of ND280 is made up of scintillator bars with wavelength shifting fiber running through them – the light collected travels along the fiber to MPPC sensors
- Over time, the scintillator bars have aged, reducing the light yield
- Not currently a problem for physics, however ageing will continue during HK era
- Large increase in light yield could be achieved by replacing the MPPCs to newer sensors, due to improvements in tech since 2009
- This refurb could massively lower the threshold for detecting eg. low energy gammas exiting the detector
- Studies ongoing at Liverpool to understand the requirements and possible benefits of this refurb



Conclusions

- The near detector suite, including ND280, are a crucial part of the T2K long baseline neutrino oscillation experiment
- An upgrade of ND280 was completed in 2024
 - Three new subdetectors
 - Improvements including neutron detection, better angular acceptance and low energy thresholds
 - Huge amount of work into designing, building, installing and commissioning these detectors
- First data has been taken with the upgrade and analyses are underway
- Targeting summer conferences to show first event selections with data
- Starting to look ahead to further upgrades for the HK era



Thank you for listening

