

**LZ & XLZD**

Liverpool Group

Ewan Fraser

May 2025



# LZ + XLZD Liverpool Group

## MPhys Students:

- ❑ Brandon Crowley
  - ❑ CNN for veto position reconstruction ft. Tea
  - ❑ PhD @ Liverpool

## PhD Students:

- ❑ Bethan Twigg
  - ❑ Incoming, starting next academic year
- ❑ Tea Hall
- ❑ Megan Carter
- ❑ Sam Woodford
  - ❑ Now a post-doc @ Edinburgh
  - ❑ XLZD!

## Research Associates:

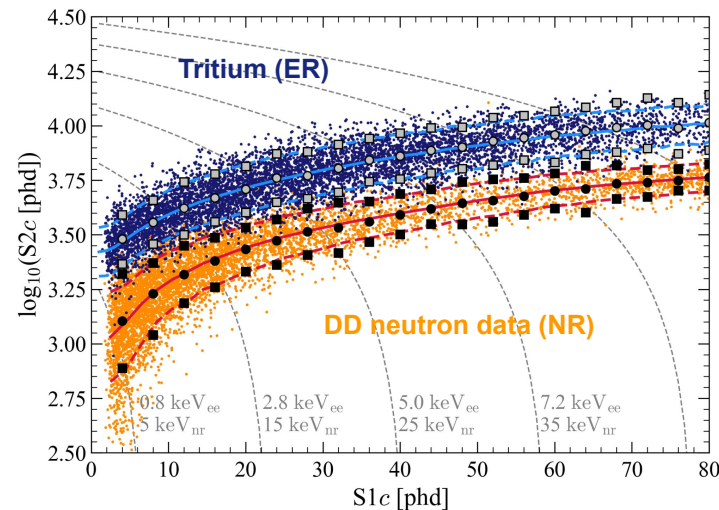
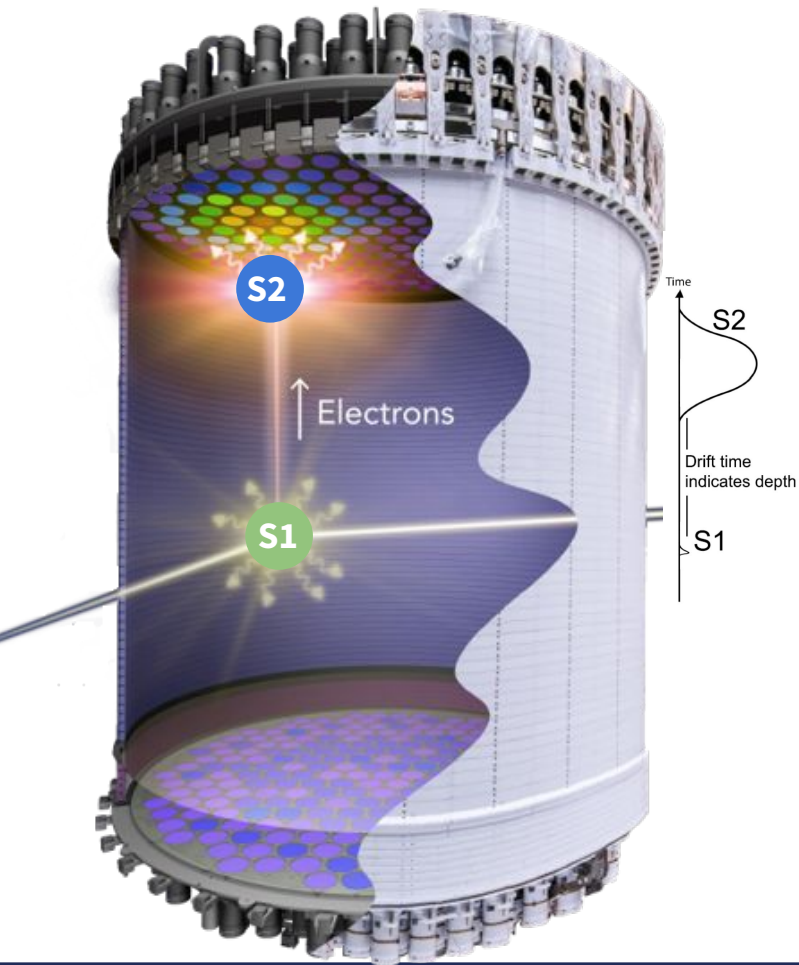
- ❑ Sean Hughes - XLZD
- ❑ Ewan Fraser - LZ

Prof. Sergey Burdin

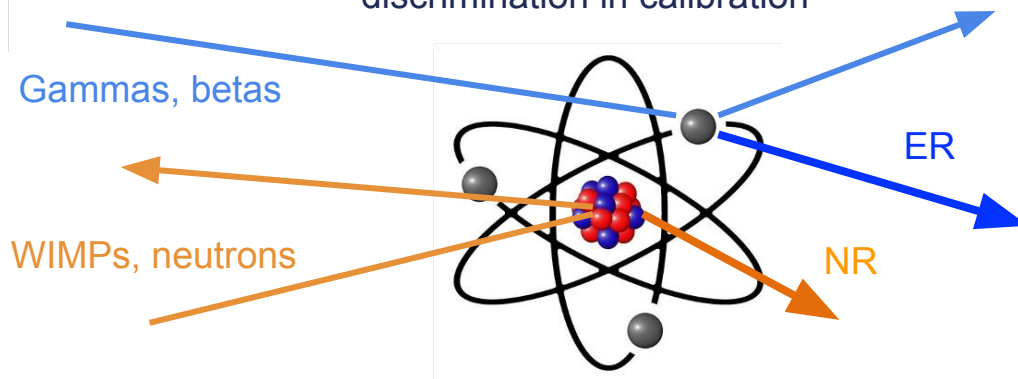


Sam Woodford's Viva  
Thesis: LZ's WS2024, Vetoes and Data Quality

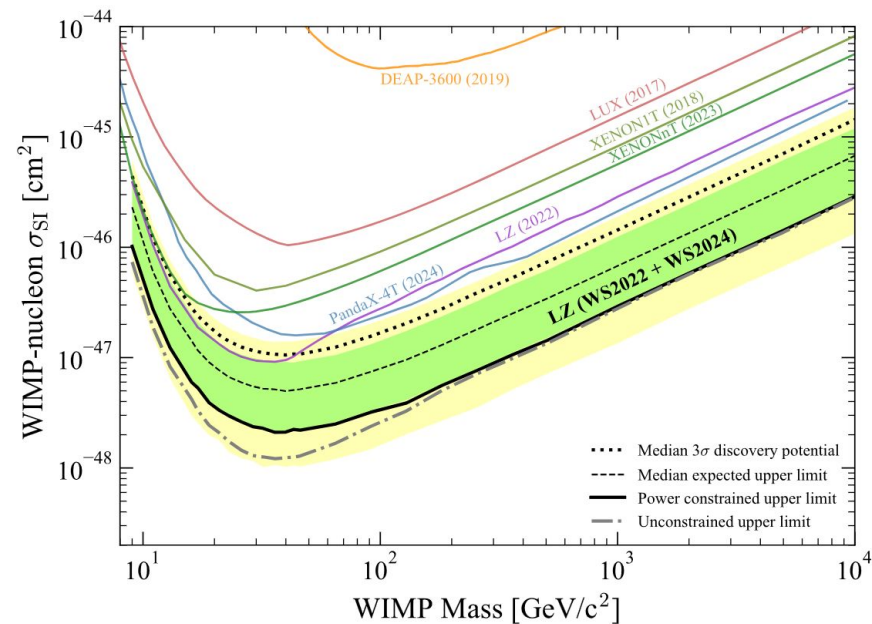
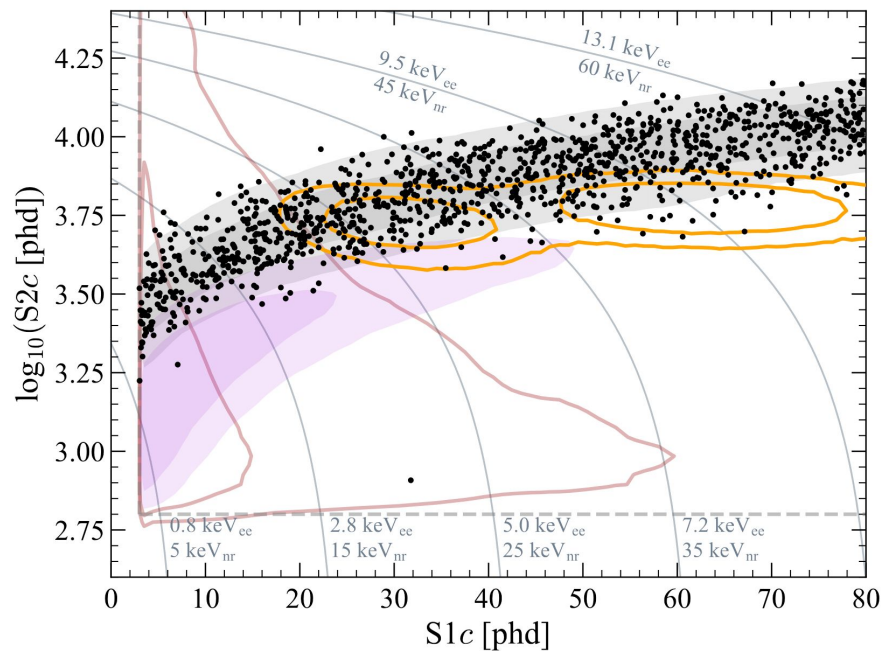
# Dark Matter Search with LXe TPC



2024 Electronic Recoil (ER) and Nuclear Recoil (NR)  
discrimination in calibration



# WIMP Search 2024 Results

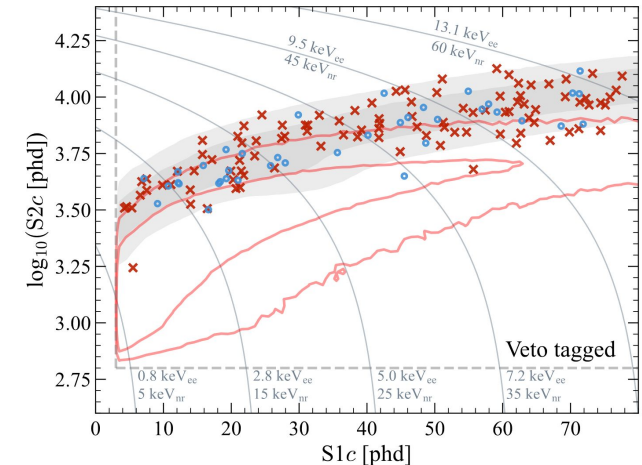
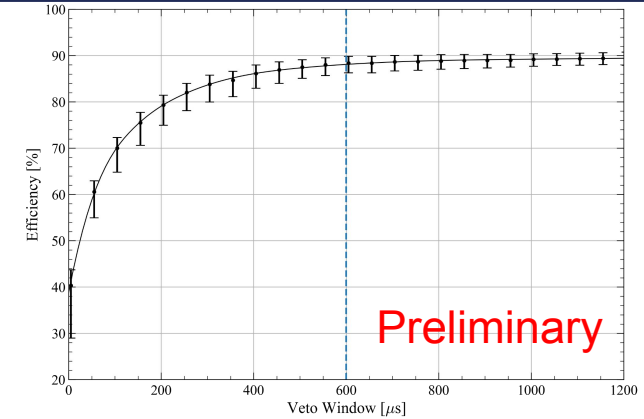


- ❑ Final exposure: 4.2 tonne years LXe from 280 live days
- ❑ Results are consistent with a background only hypothesis
- ❑ Best limit from combined analysis of  $\sigma_{SI}=2.1 \times 10^{-48} \text{ cm}^2$  for  $36 \text{ GeV/c}^2$



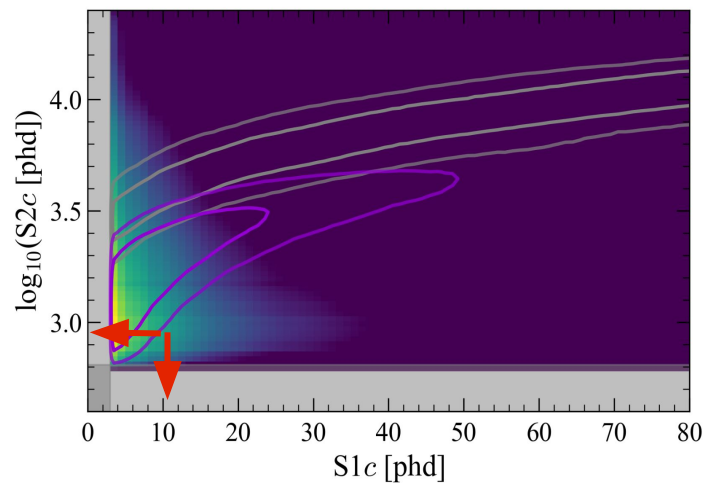
# Neutrons & Vetoes

- ❑ Coordination of Vetoes & Neutron background group
- ❑ Neutrons are the only WIMP Search (WS2024) NR background
  - ❑ If they single scatter
- ❑ Outer Detector + Skin veto:
  - ❑  $89 \pm 3\%$  neutron veto efficiency from AmLi calibration
  - ❑ 3% deadtime, (false veto rate \* veto window length)
- ❑ Constrain the likely number of neutron events in WS2024
- ❑ Veto tagged sideband fit
  - ❑ Expectation  $0^{+0.2}$  single scatter neutrons for WS2024
- ❑ Infer from the number of multiple scatter neutrons
  - ❑  $0.3 \pm 0.2$  single scatter neutrons for WS2024

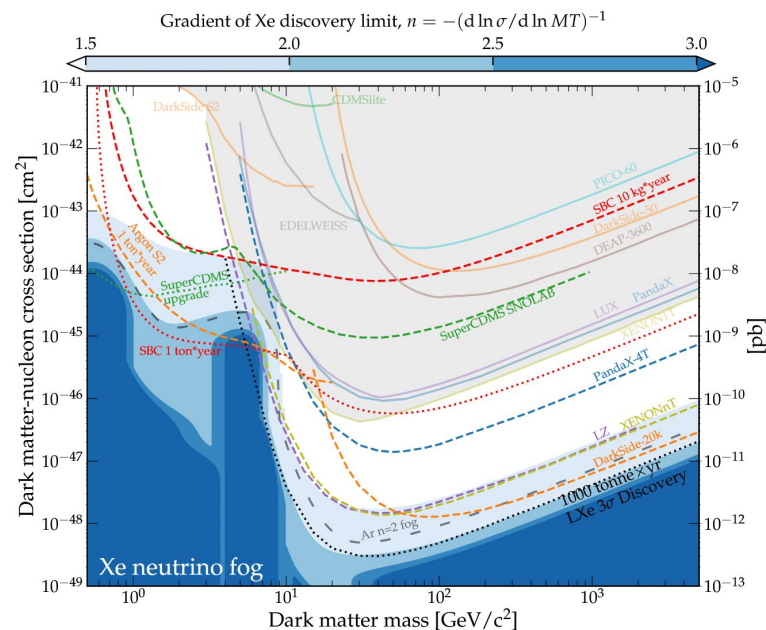


# Low Mass WIMP Search

- ❑ WS2024 lowest WIMP mass tested 9 GeV/c<sup>2</sup>
- ❑ Extend search to lower masses O(GeV)
- ❑ New background expected!
  - ❑ CE $\nu$ NS from <sup>8</sup>B solar neutrinos



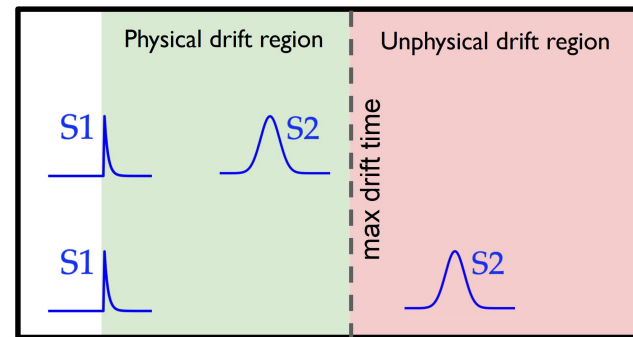
LZ accidental background PDF



Snowmass 2021 - 2203.08084

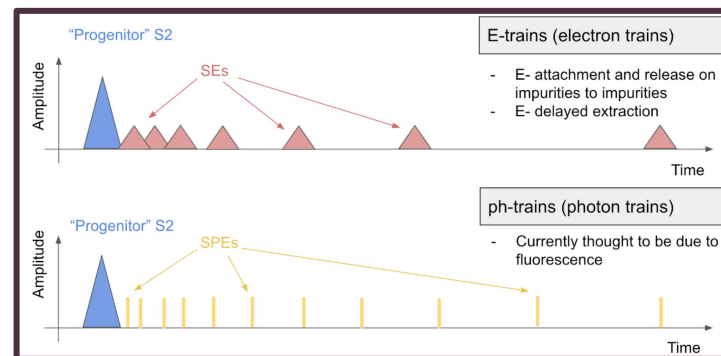
# Data Quality

- ❑ Coordination of LZ's Data Quality group
- ❑ Accidental backgrounds most prevalent at low energies
  - ❑ Coincidence of isolated S1 and S2 pulses
- ❑ Pile-up photons or electrons can appear to be isolated S1s and isolated S2s respectively
- ❑ Exclusion of time periods with elevated background rates



Accidental background events

- ❑ Contemporaneous and retrospective data quality monitoring
  - ❑ Online: Underground Performance Monitor (UPM)
    - ❑ Primarily monitors PMT and Trigger health
  - ❑ Offline: Physics Readiness Monitor (PREM)
    - ❑ Detector health, which data should be included in WS



Photon and electron trains

# XLZD & the Xenon Futures program



- MoU signed in 2021 by **XENONnT**, **LUX-ZEPLIN**, **DARWIN**
  - Collaboration formed end of 2024 => **XLZD**
- Frequent meetings with collaborators, yearly collaboration meetings
  - Next is June 2025 @ LNGS
- Xenon Futures program concluded late March 2025
- Pivoted fully towards XLZD project
  - Sergey leading WP2 (Outer Detector)
- Working towards bringing the definitive direct detection experiment to **Boulby**
  - => **XLZD@Boulby**
  - Working towards CDR -> late 2025

## [XLZD White Paper](#) & [Design book](#)

The XLZD Design Book: Towards the Next-Generation Liquid Xenon Observatory for Dark Matter and Neutrino Physics

J. Aalbers<sup>1</sup>, K. Abe<sup>2</sup>, M. Adrover<sup>3</sup>, S. Ahmed Maouloud<sup>4</sup>, D. S. Akerib<sup>5,6</sup>, A. K. Al Musalhi<sup>7</sup>, F. Alder<sup>8</sup>,

A Next-Generation Liquid Xenon Observatory for Dark Matter and Neutrino Physics

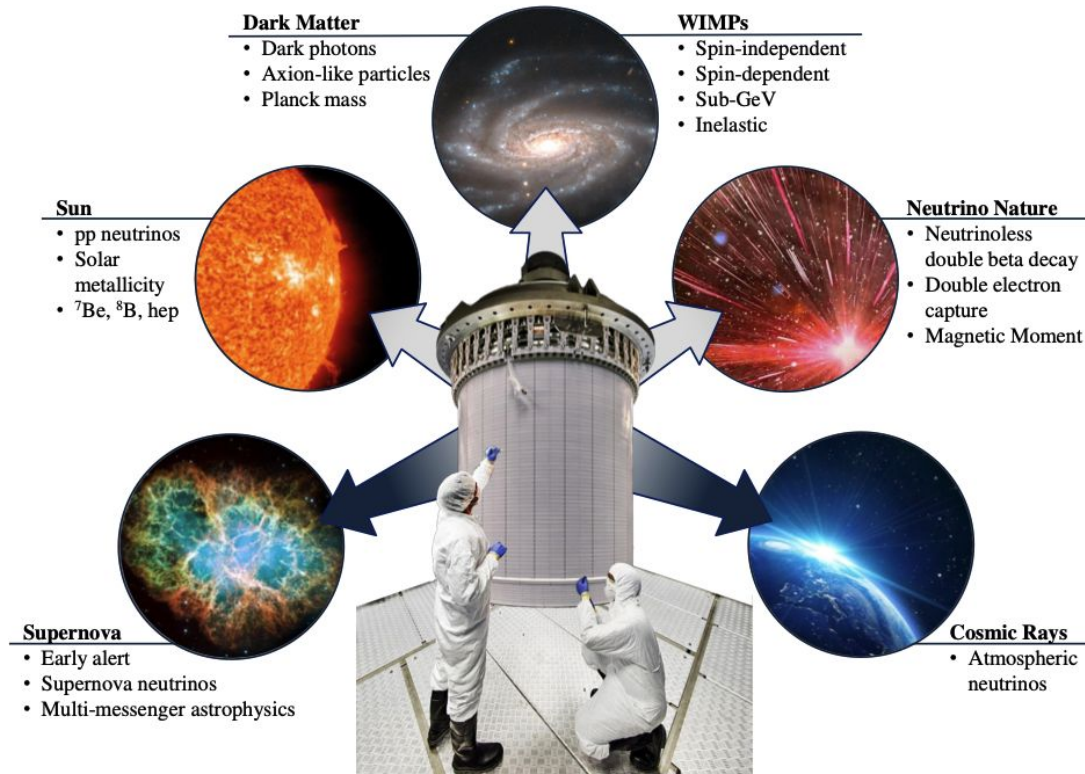
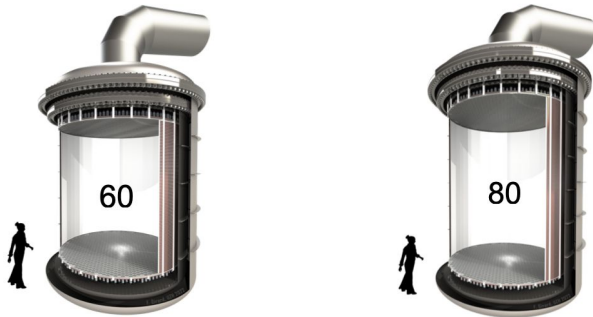
J. Aalbers,<sup>1,2</sup> K. Abe,<sup>3,4</sup> V. Aerne,<sup>5</sup> F. Agostini,<sup>6</sup> S. Ahmed Maouloud,<sup>7</sup> D.S. Akerib,<sup>1,2</sup> D.Yu. Akimov,<sup>8</sup> J. Akshat,<sup>9</sup> A.K. Al Musalhi,<sup>10</sup> F. Alder,<sup>11</sup> S.K. Alsum,<sup>12</sup> T. Althueser,<sup>13</sup> C.S. Amarasinghe,<sup>14</sup> F.D. Amaro,<sup>15</sup> A. Ames,<sup>1,2</sup>





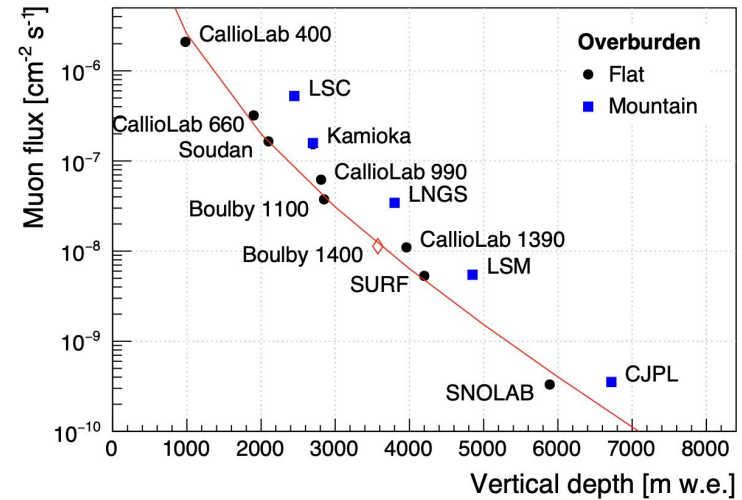
# XLZD – Overview

- Next generation xenon-based detector
  - Dual phase TPC
  - Multi-purpose observatory
- 60t -> 80t of (active) LXe target mass
- Will reach neutrino floor and **definitively** rule out or discover WIMPS
- Can be used also to study the neutrino sector – rich physics program
  - Neutrinoless double beta decay
  - Astrophysical sources

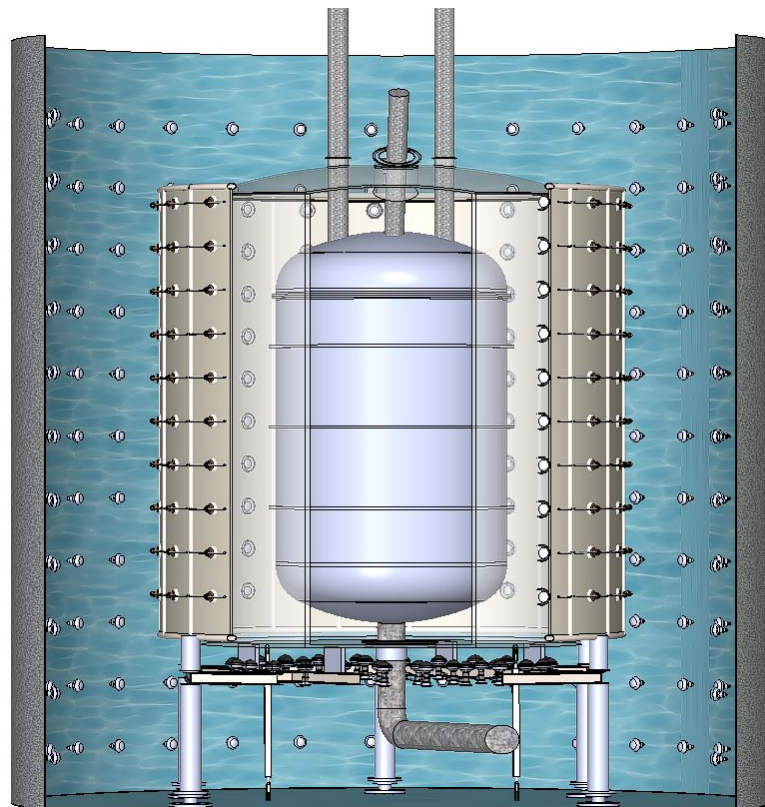


# XLZD – Possible locations

- Five sites have submitted their intention to host XLZD:
  - Kamioka, Japan
  - SNOLAB, Sudbury
  - SURF, South Dakota
  - Boulby, North Yorkshire
  - LNGS, Gran Sasso
- Each site has advantages and disadvantages
- Determines which technologies can be used in the experiment
- Boulby has rich Dark Matter history:
  - Hosted the ZEPLIN experiment
  - Boulby Underground Germanium Suite (BUGS)

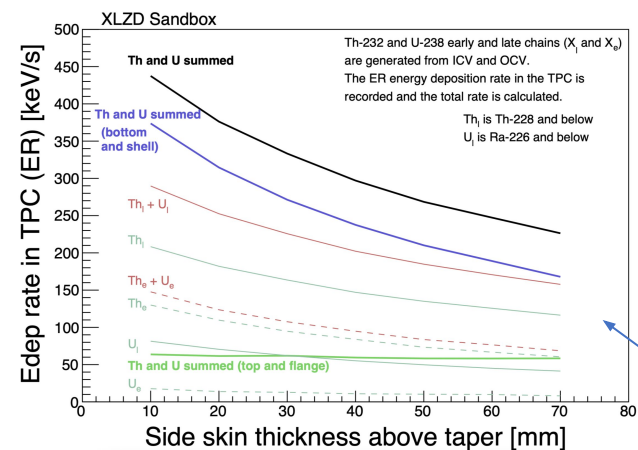


- Geant4-based software developed by Liverpool
  - XLZD Sandbox
- Enables for optimisation of the various subsystems
- Benchmarking possible with RAT-PAC2
- Focus with Liverpool: The design of the Outer Detector
  - Transfer of expertise from LZ -> XLZD
    - Neutron Veto Efficiency
    - Position reconstruction
- Maximise for neutron veto efficiency, test various new technologies
  - Selection of OD media: **Gd-Water, Gd-WbLS, Gd-LS?**
  - Synergy with collaborators across the world
    - BUTTON and 30t BNL demonstrator

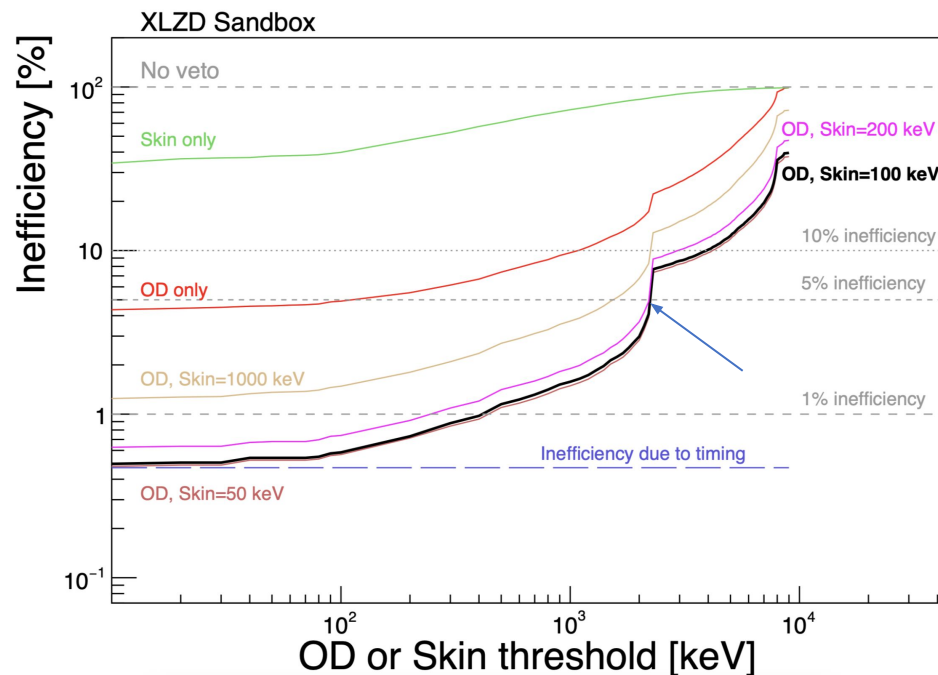
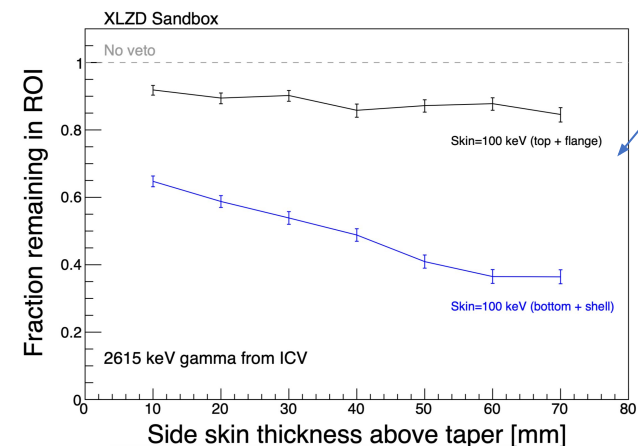


CAD from James Coleman-Mills

# Physics Studies



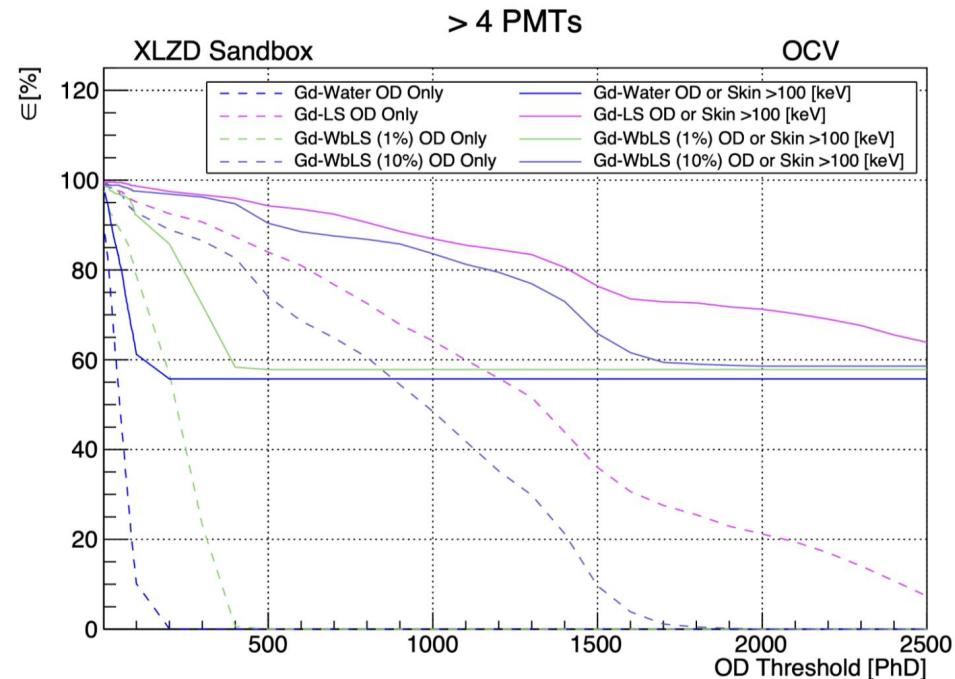
Optimising  
Skin  
thickness



Using only threshold in energy:

**Semi-medium agnostic view (0.1% Gd): Skin + OD can together surpass 95 % tagging efficiency**

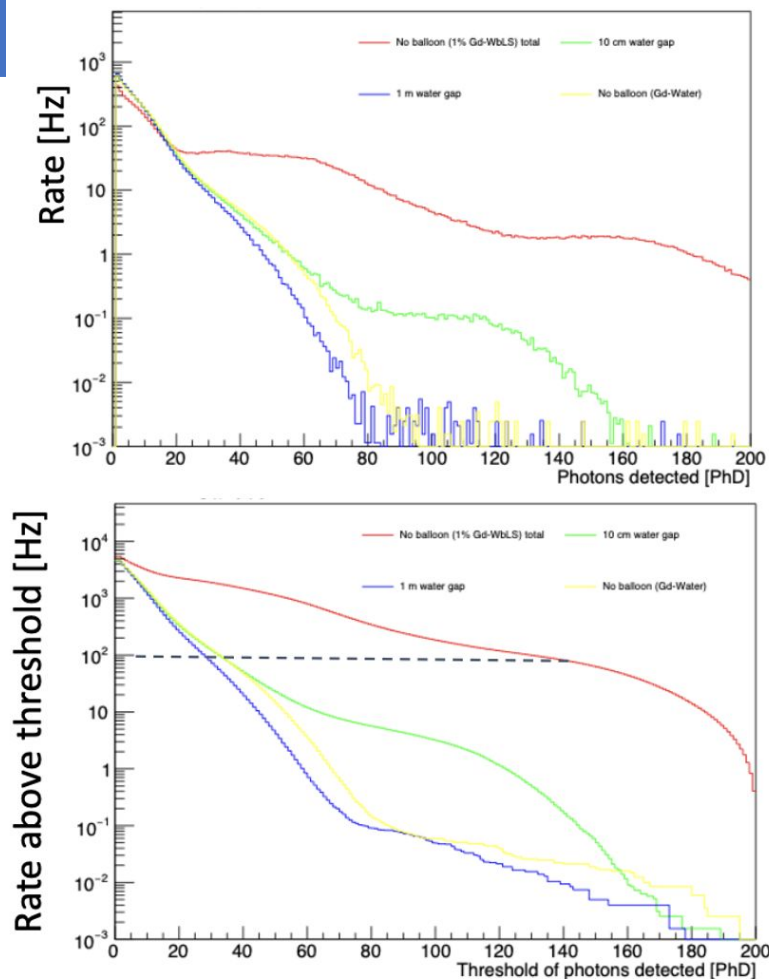
# Physics Studies



When looking at photons detected, story becomes a little more complicated:

If PMTs are set inside the WbLS => rate from PMTs is too high!

Studies ongoing to minimise the rate...





# Conclusion

- Conclusions of XF3 -> Pivoted to XLZD
- Significant contributions from Liverpool towards the XLZD@Boulby initiative
  - Development of XLZD Sandbox
  - Prototyping ongoing
- Working towards CDR late this year!

