



UNIVERSITY OF
LIVERPOOL

Machine Learning and Crystal Xe R&D for Improving the Sensitivity of Existing Dark Matter Searches

Tea Hall

Supervised by Dr. Ewan Fraser & Prof. Sergey Burdin

with thanks to University of Texas at Austin's LZ/Crystalize group

HEP Annual Meeting 2026



Dark matter direct detection experiment

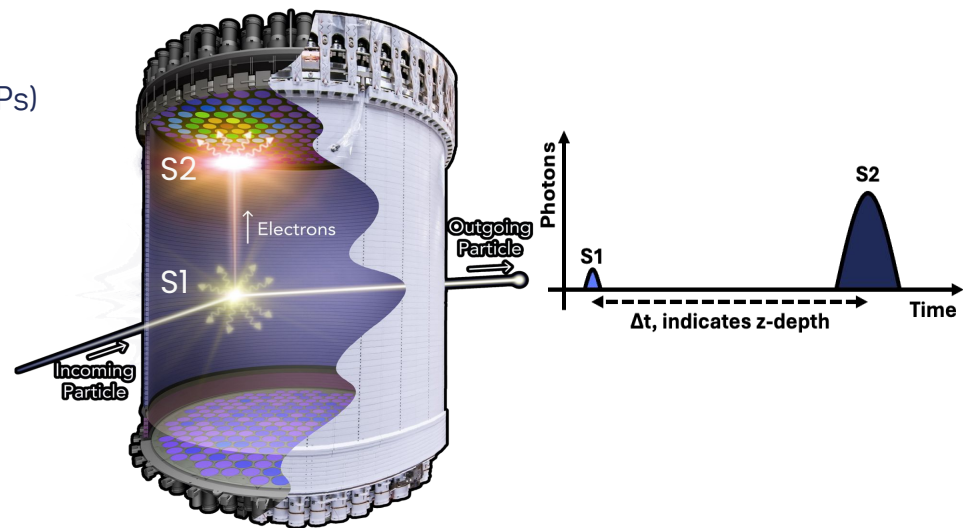
- Aim: detect Weakly Interacting Massive Particles (WIMPs)

Dual-phase time projection chamber (TPC)

- 7t of active liquid Xe (LXe)

Two signals from interactions:

- Prompt scintillation (S1)
- Delayed electroluminescence from drifted electrons (S2)



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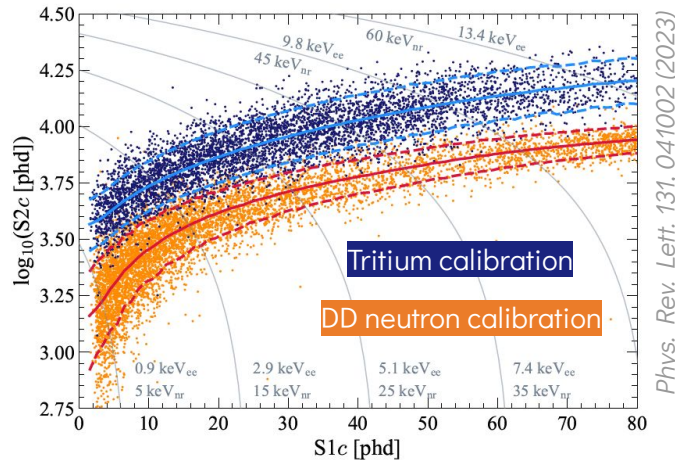
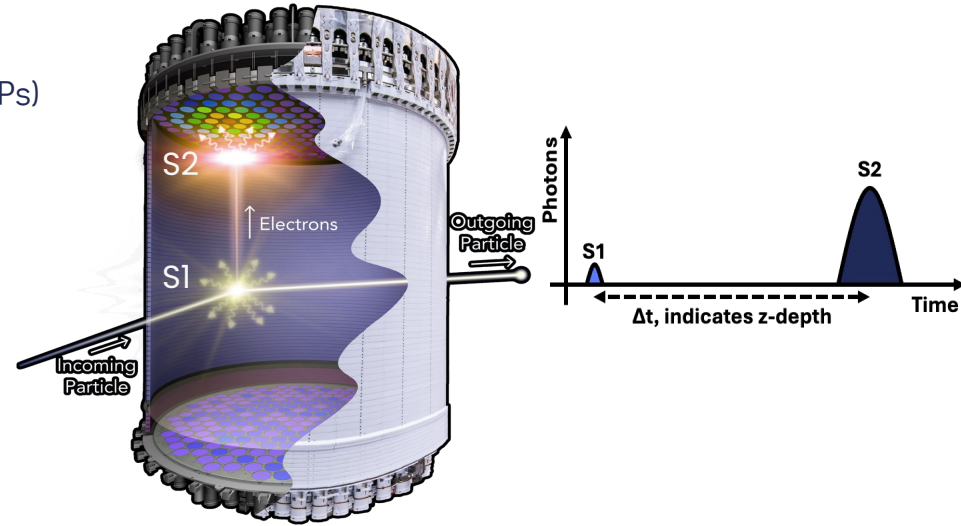
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S2/S1 → interaction type:

- Nuclear recoil (NR) – small ratio, e.g. WIMPs, neutrons
- Electron recoil (ER) – large ratio, e.g. β 's, γ 's

3D reconstruction:

- (X, Y) - S2 light pattern
- Z - Δ time between S1 and S2

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Number of ionisation electrons which reach the liquid surface after an interaction in LXe.

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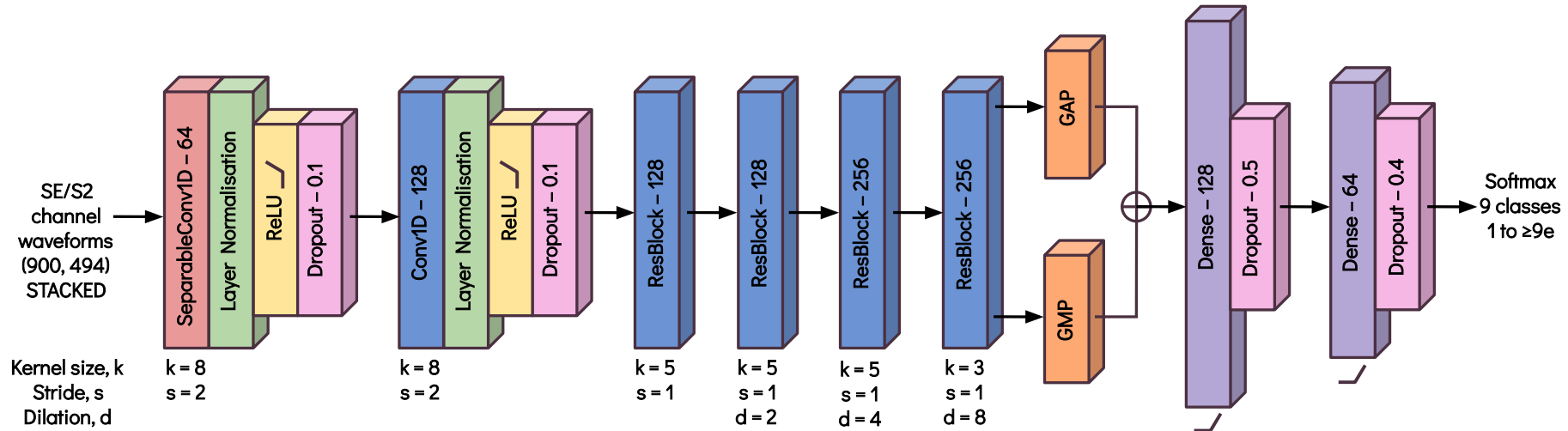
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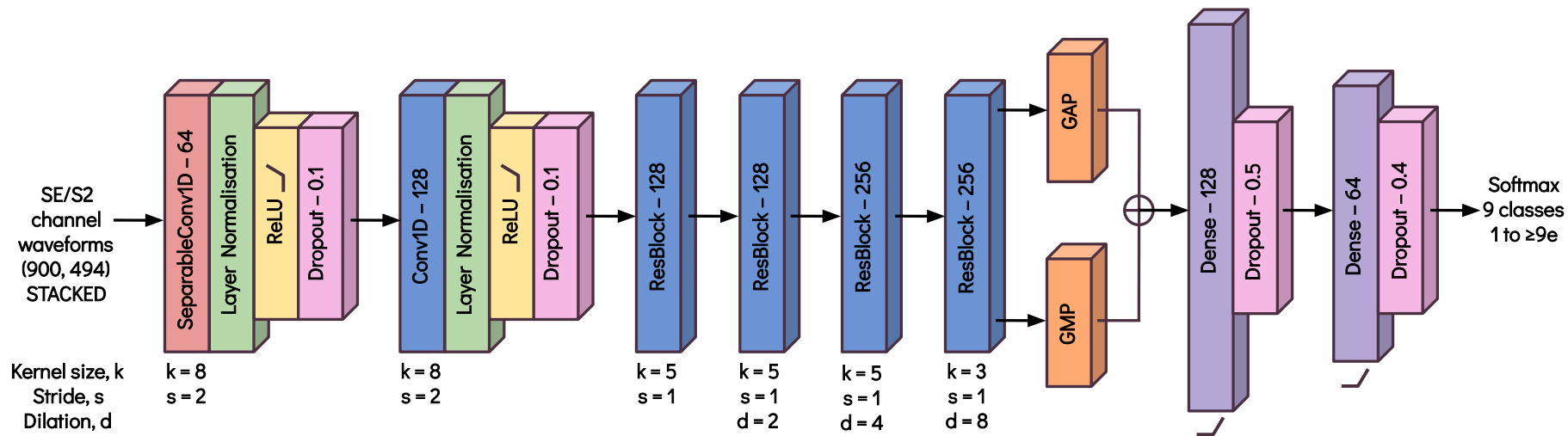
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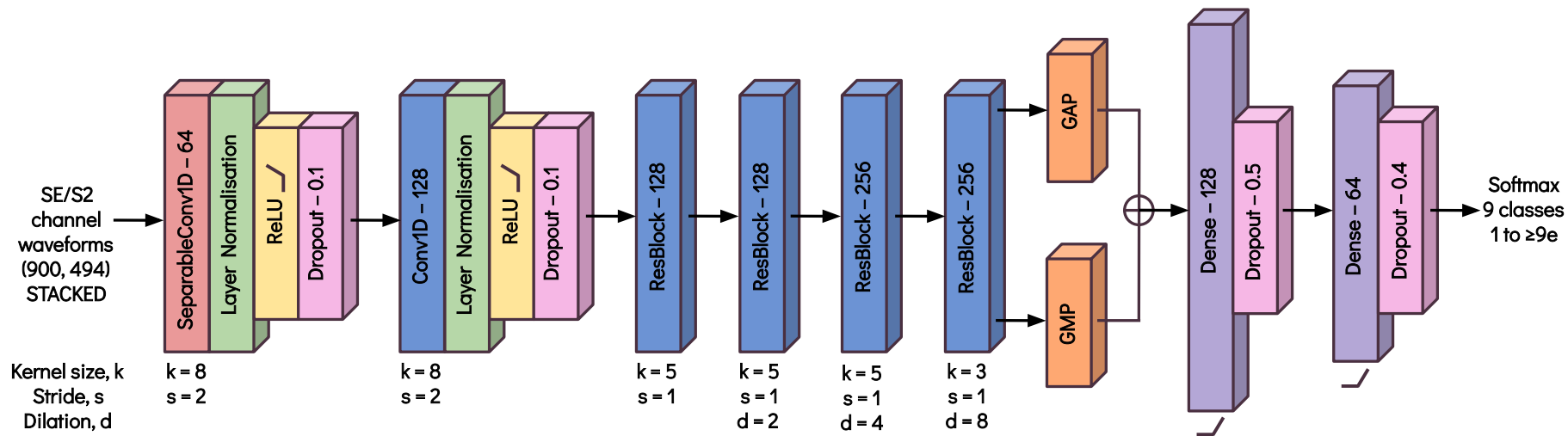


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Model trained on 1-10e waveforms from electron bomb simulations.

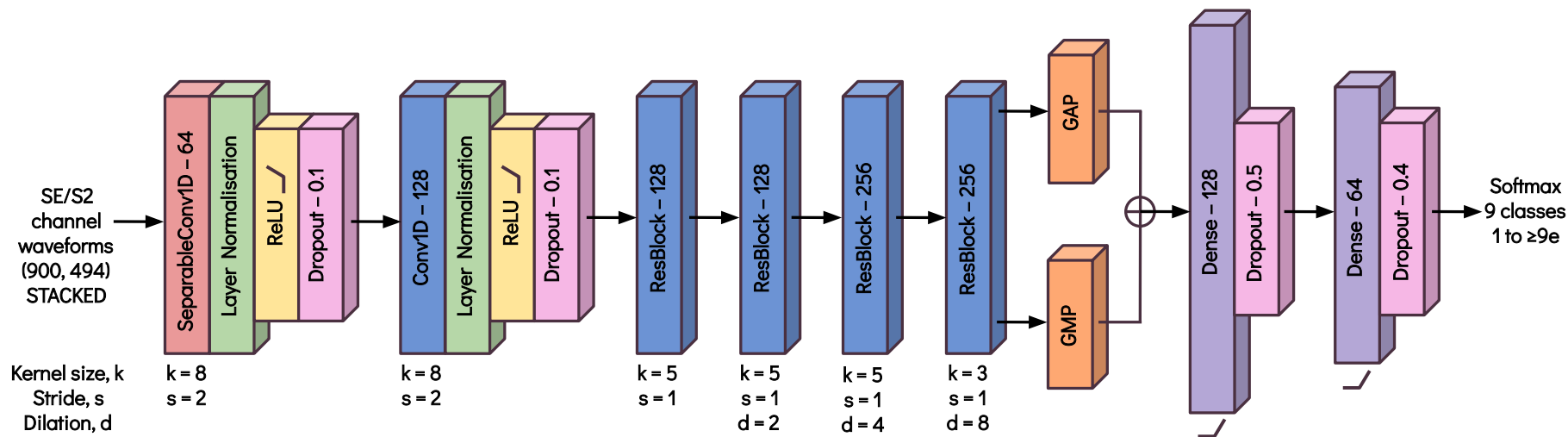
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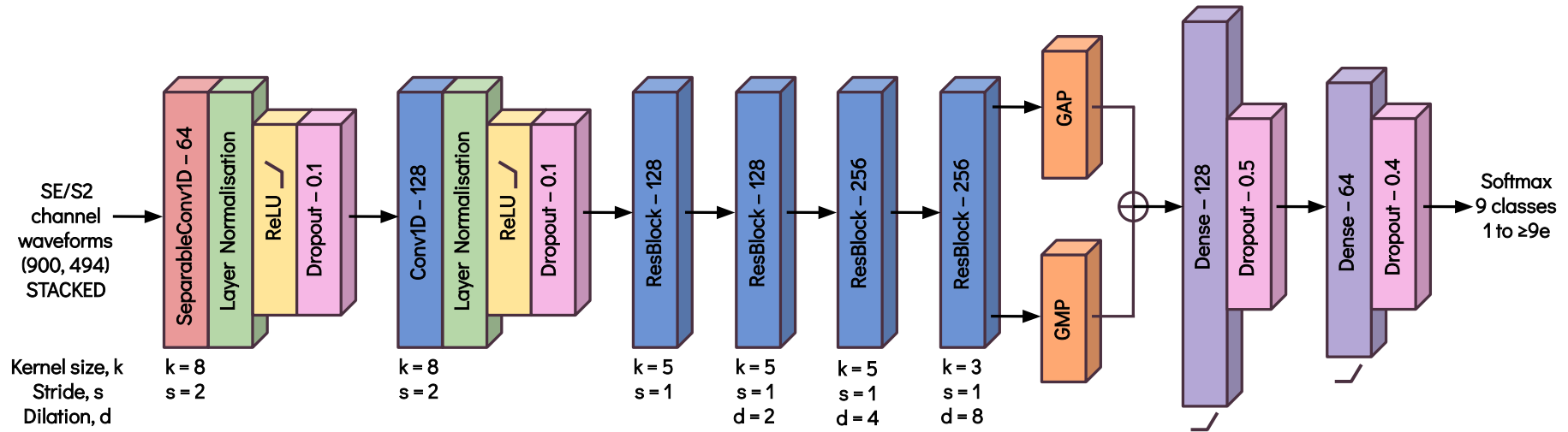


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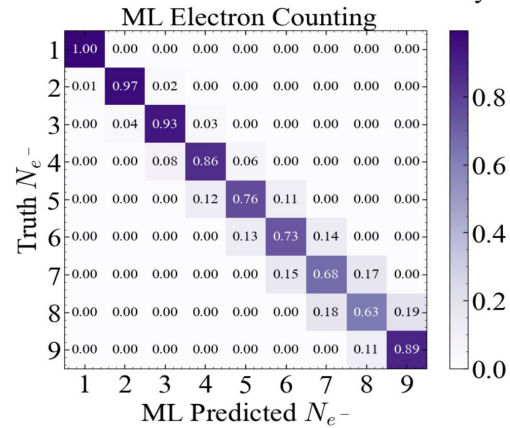


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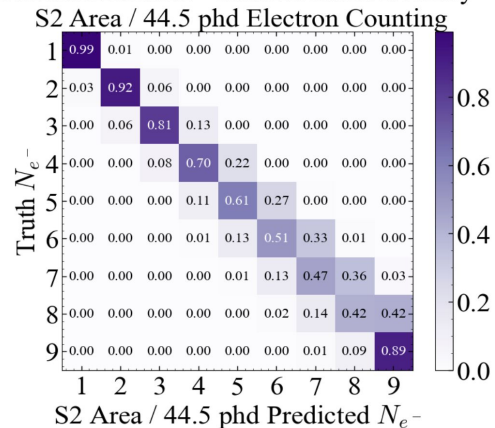
- Channel waveforms from all TPC channels – 494 PMTs across two arrays
- $N_{e, \text{true}}$ labels from MCTruth simulation z-vertices
- Modified class-distance weighted cross-entropy loss [[arXiv:2412.01246](https://arxiv.org/abs/2412.01246)]
 - Reduced training time vs standard categorical cross-entropy loss

ML e^- counting > standard approach

Confusion Matrix - Classification Accuracy

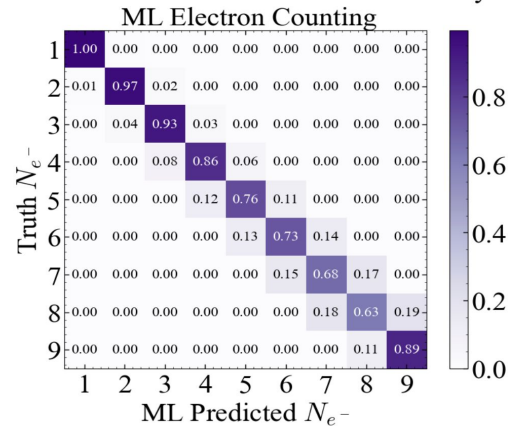


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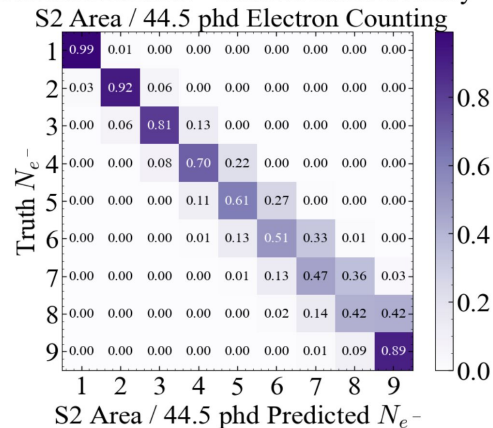


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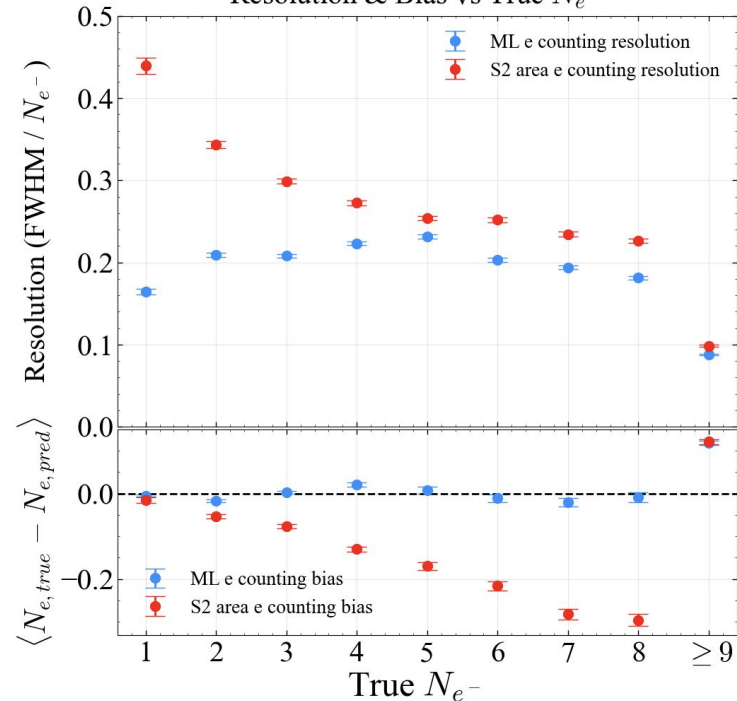
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Resolution & Bias vs True N_{e^-}



Full waveform vs integral only → better discriminating power

- Less misclassification & leakage → bias centred around zero
- More information → better resolution

S2 approach overpredicts, increasing bias with increasing S2 size

- Area-only susceptible to gain fluctuations, PMT afterpulsing, etc.

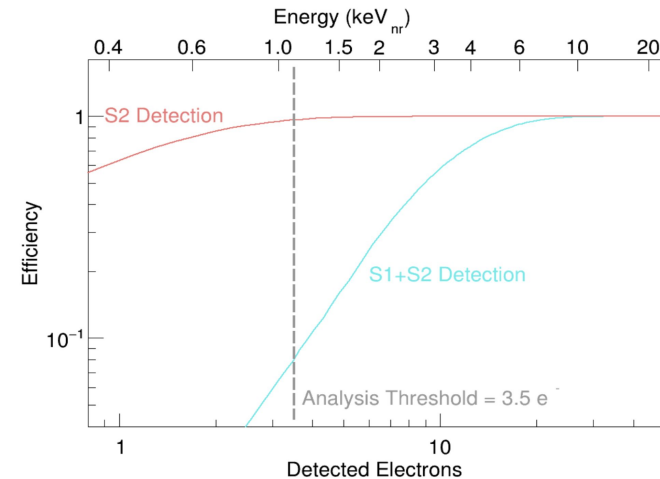
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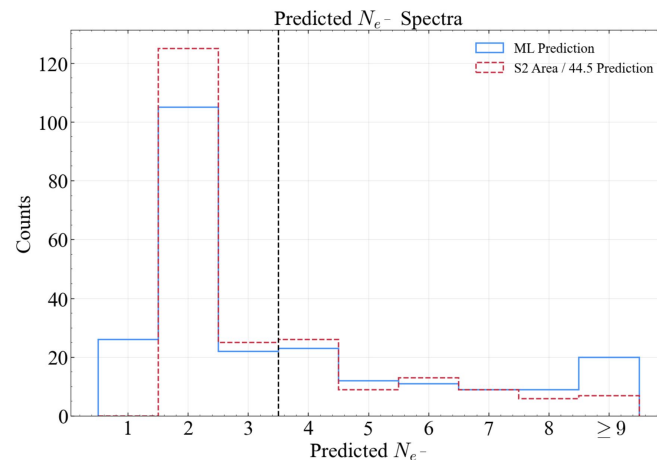
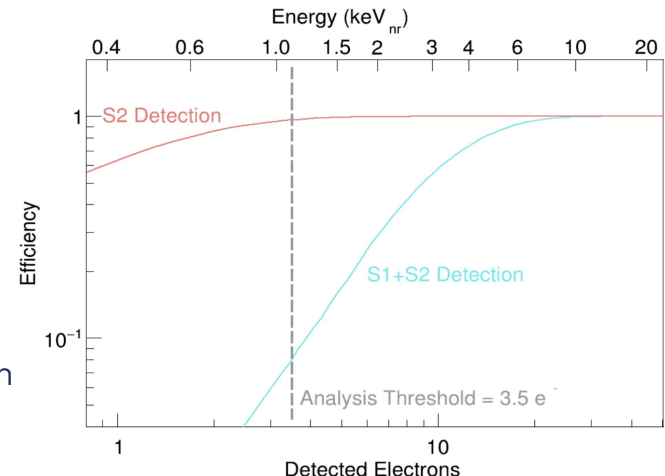
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Discrete $N_e \rightarrow$ improved S2 energy resolution \rightarrow better background discrimination

Small e^- S2s can leak into this continuous N_e space

- New cut at $4e^-$ in discrete N_e space \rightarrow 25% reduction of S2s at threshold



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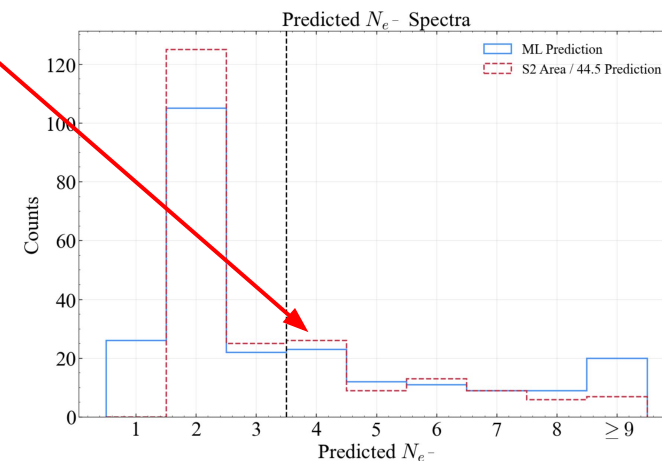
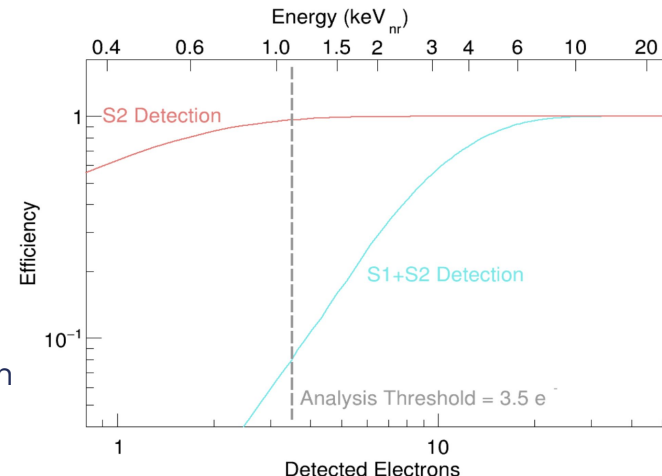
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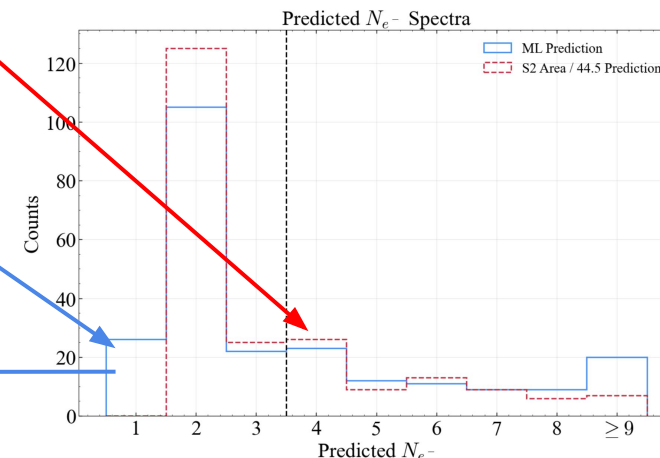
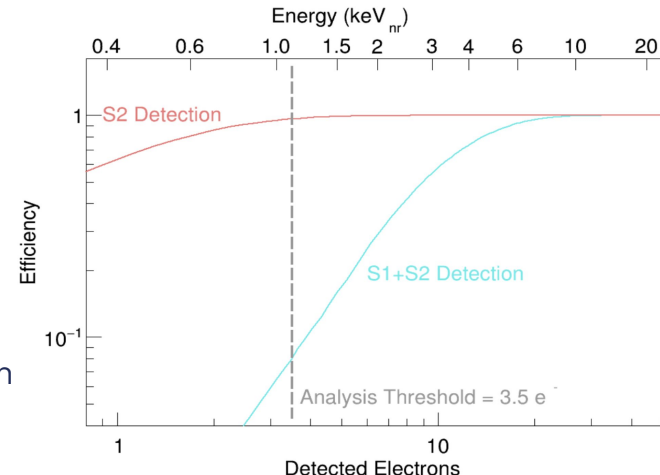
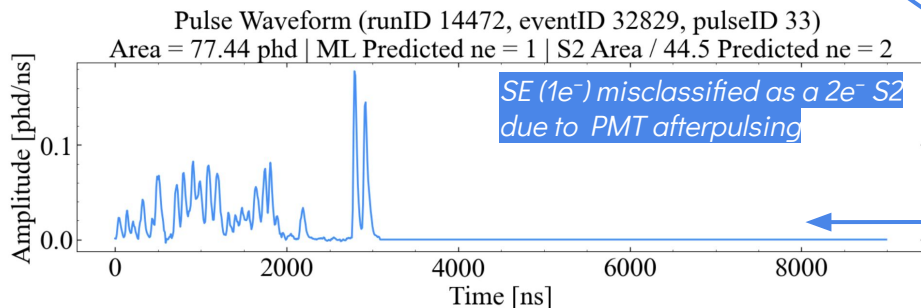
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Identification of $1e^-$ events misclassified as S2s by pulse classifier



Reduce S2-only analysis threshold & conduct search in $1-5e^-$ regime

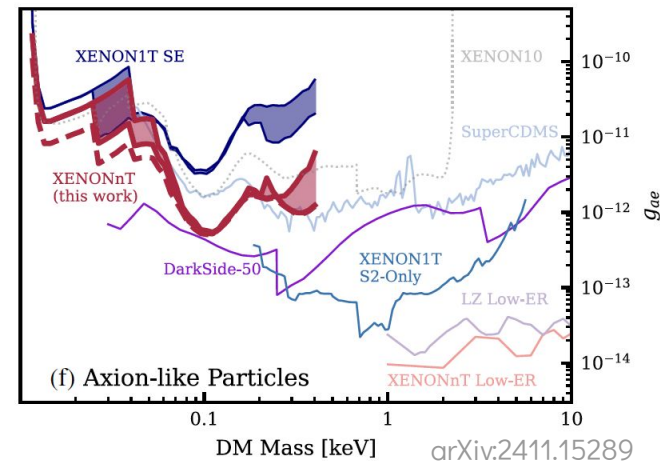
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XENONnT showed it was possible in this regime

- World-leading limit in S2-only regime to ALPs
- Limited by spurious e⁻ emission from grids → ~2.9x loss in fiducial mass



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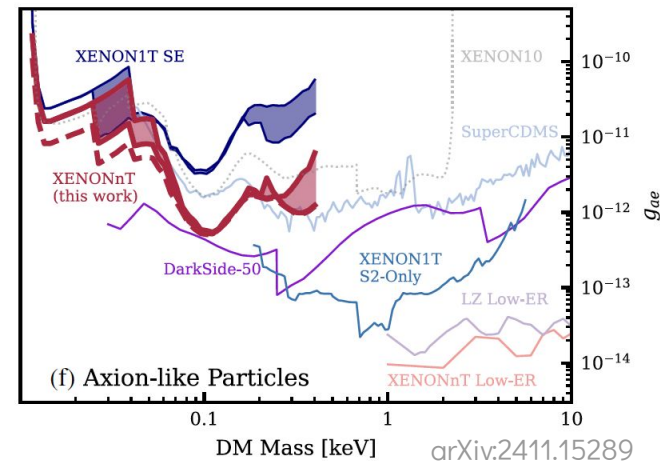
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LZ can go further

- Improved reconstructed N_e with less bias
- Larger active volume \rightarrow more fiducial mass
- Alternative electrode grid design



Toward a Few-Electron Search in LZ

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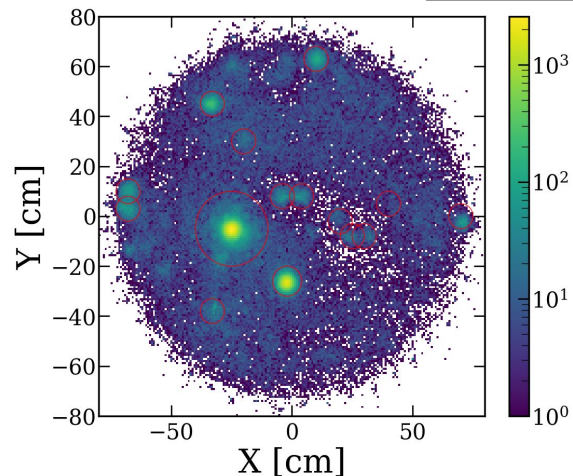
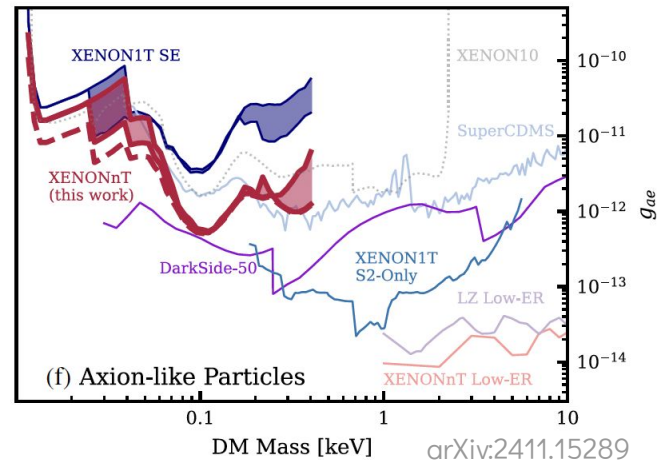
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Analysis requirements for a light DM search with LZ

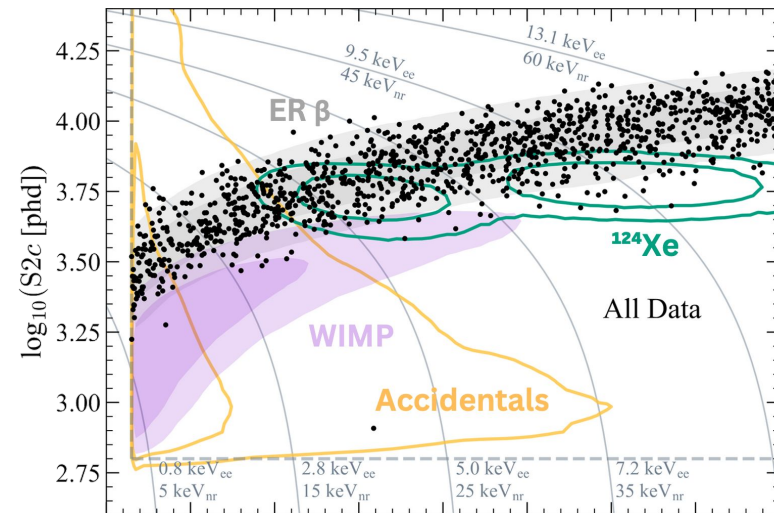
- Aggressive exclusion cuts e.g. hotspots
- Robust background modelling
- Fiducial volume tuned in (x,y,z)
 - Z-reconstruction from e⁻ arrival times → ML model in development

Discrete N_e counting can help LZ place competitive limits in these regime!



Existing noble liquid dark matter searches are dominated by Radon

- Rn chain events make up ~67% of events in LZ's WS2024

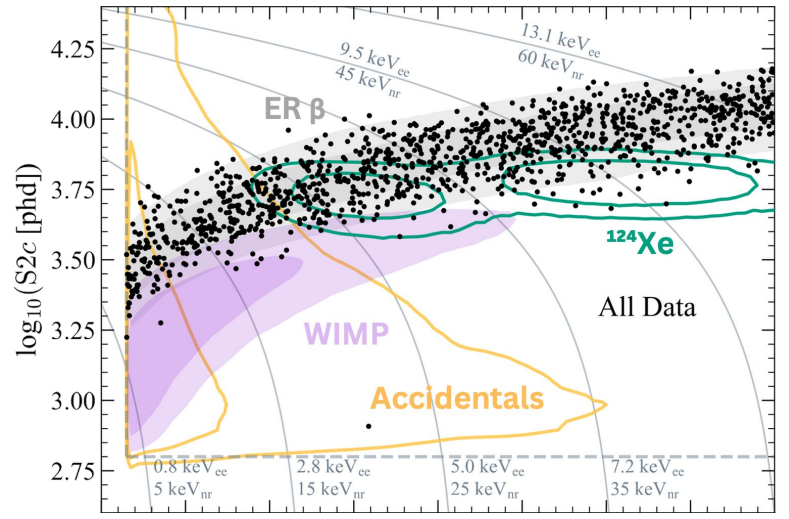
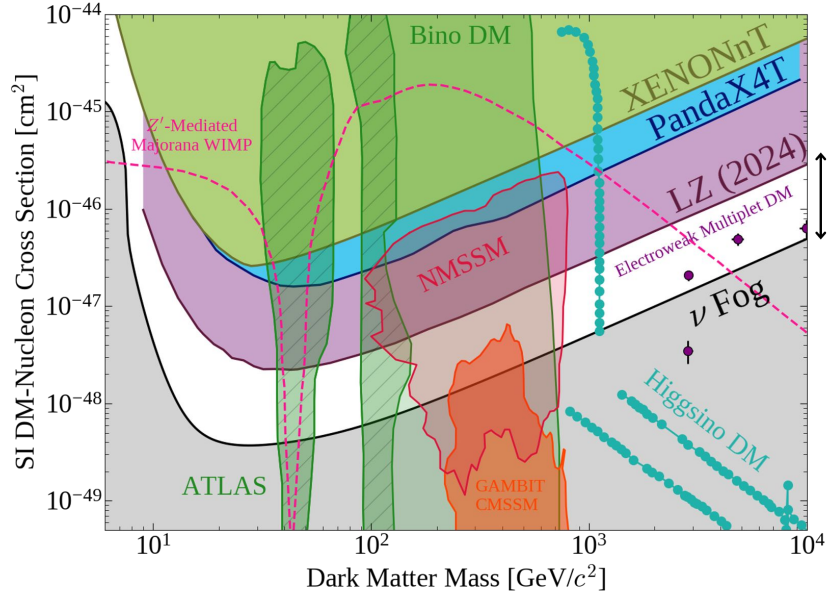


Phys. Rev. Lett. **135**, 011802

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Increasing sensitivity such that experiments become limited by the neutrino fog requires that the Rn background be minimised by 20x

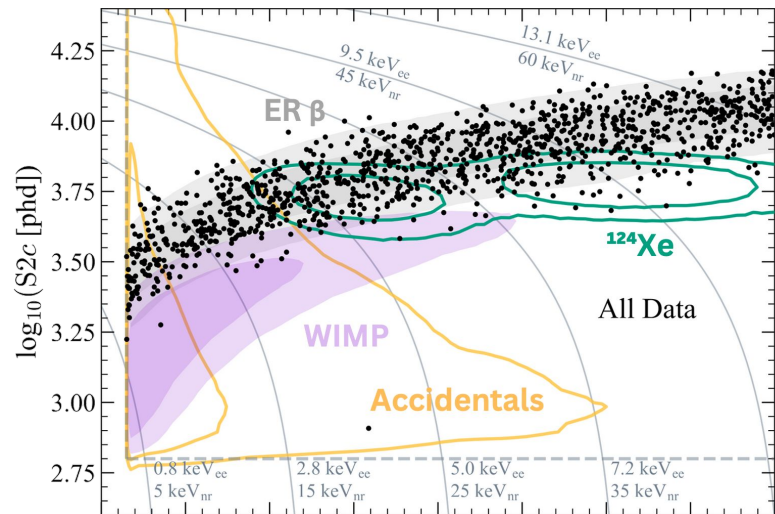
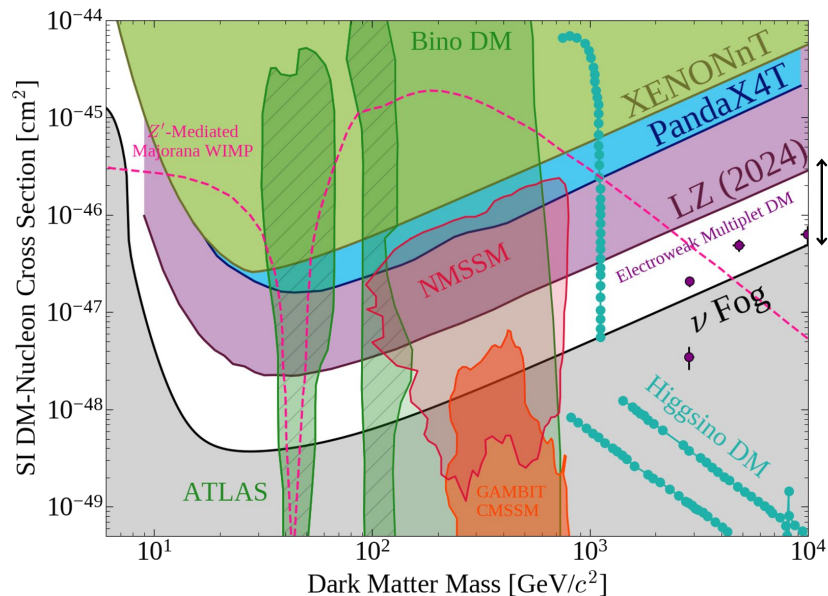


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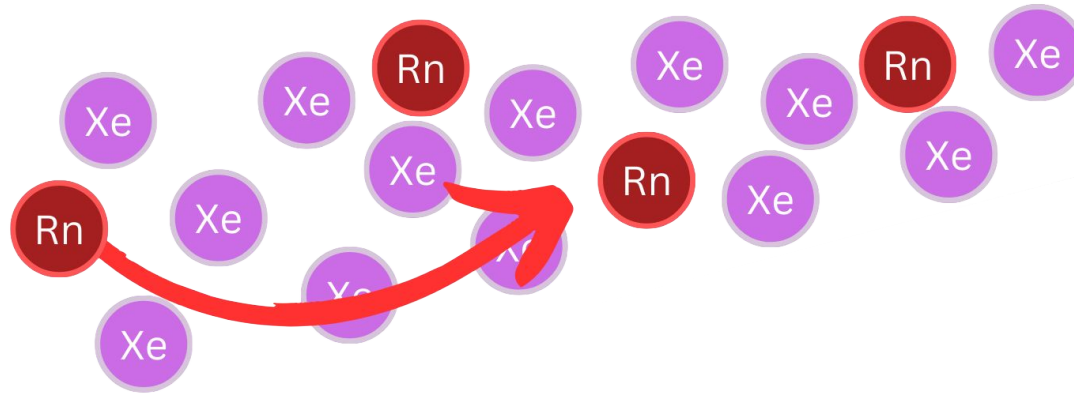


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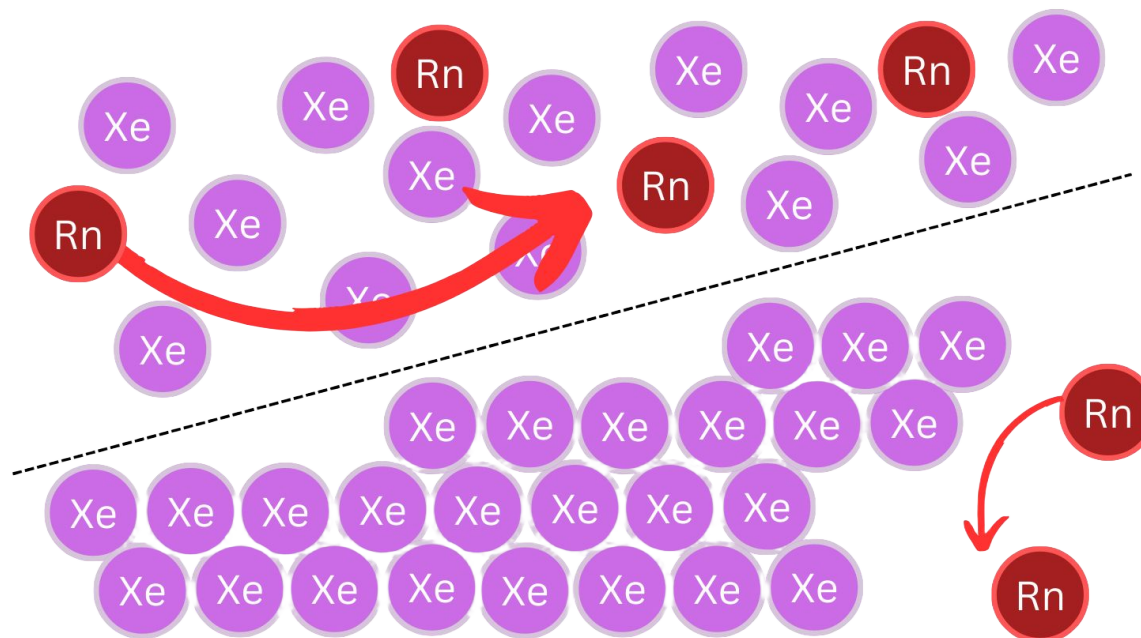
How can Rn be mitigated?

- Purify the Xe more effectively
- Tag Rn events during detector operation
- ★ Stop Rn from entering

A liquid noble detector allows Rn to be circulated during operation



A liquid noble detector allows Rn to be circulated during operation



While a frozen noble detector allows Rn to be removed from the bulk of the detector

CrystalLiZe is a test-bed experiment for crystal Xe (CXe) R&D based at the University of Texas at Austin.

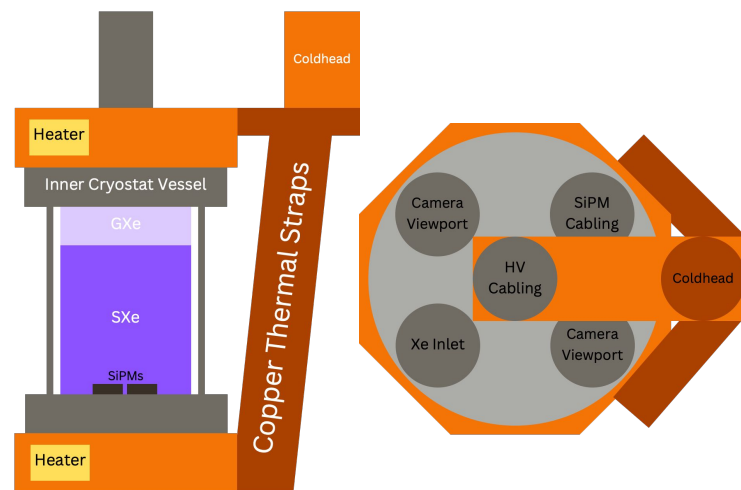
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CrystaLiZe system

- 2–9kg Xe in smallest vessel
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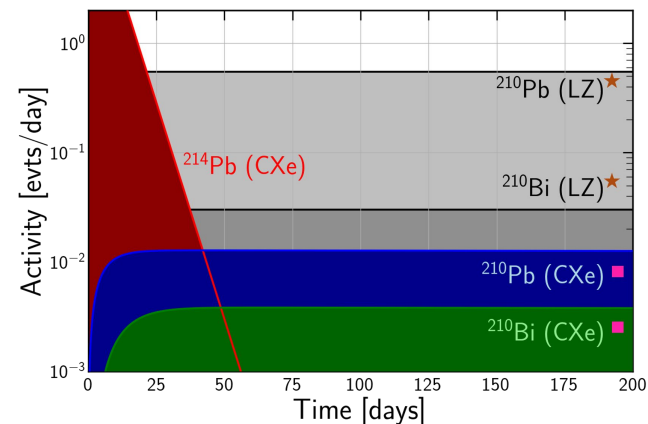
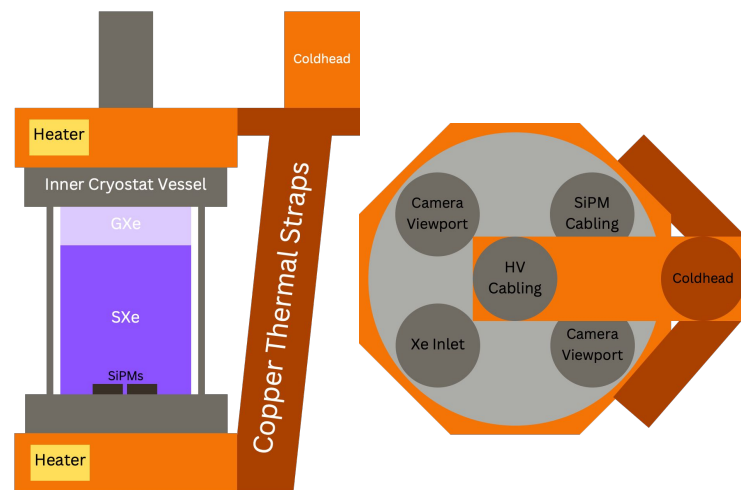
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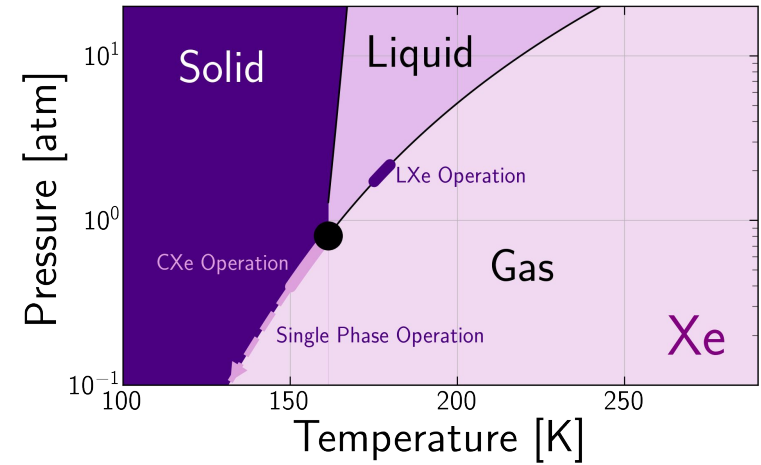
CXe has major benefits vs LXe

- Rn excluded by > 600x → increased fiducial volume
- Improved electron mobility
- 2x reduction in accidental coincidences



Bridgeman method for CXe growth [[arXiv:2312.15082](https://arxiv.org/abs/2312.15082)]

- ~1.2 K/cm temperature gradient across system
- Slow descent along the phase boundary from triple point

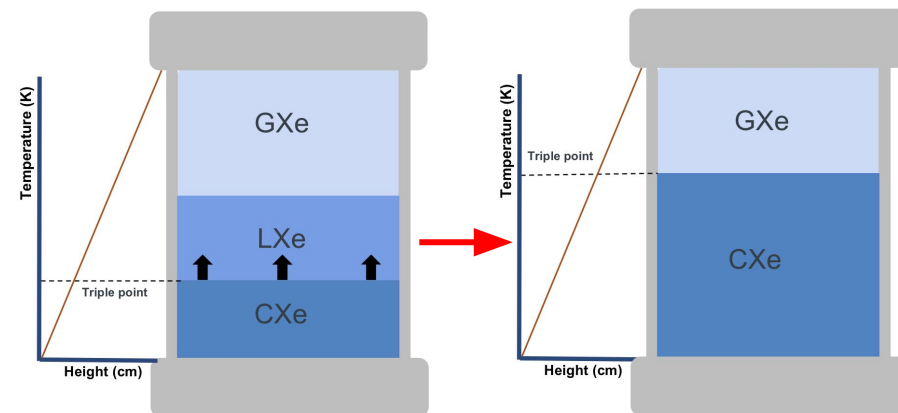
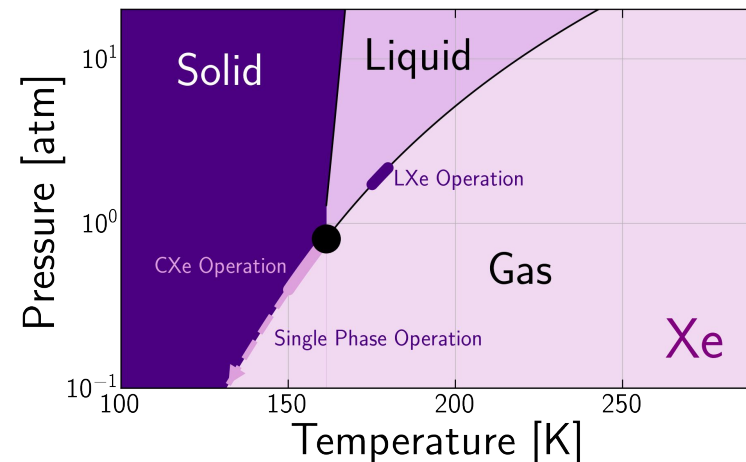


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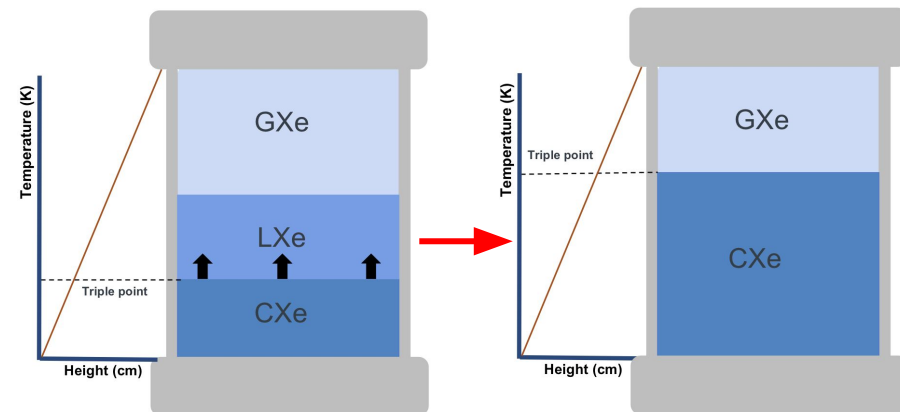
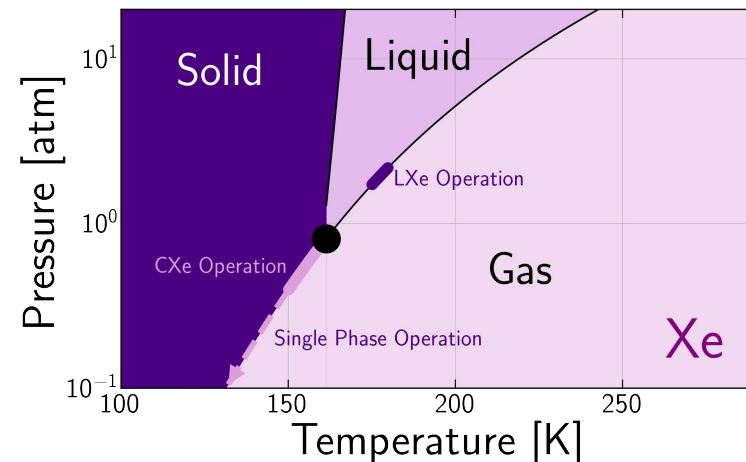
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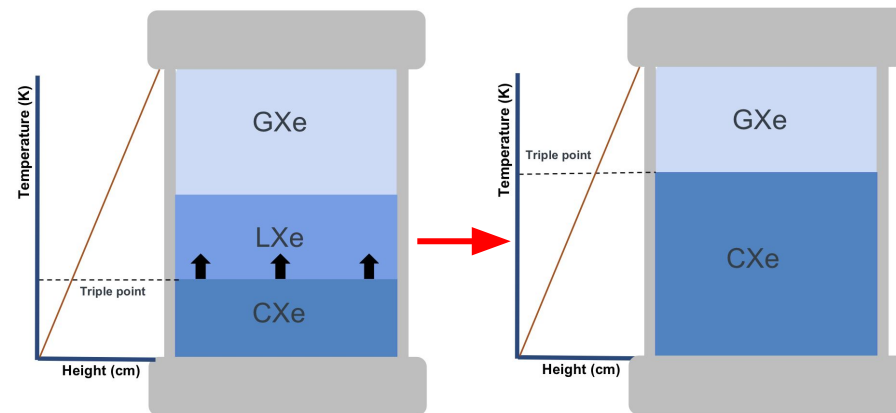
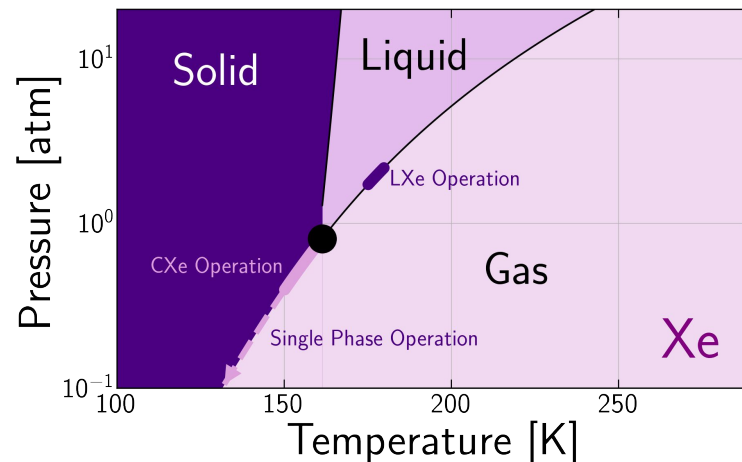
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Time to freeze (TTF) reduced from > 1 week → 18 hours

- Ongoing investigations with various Xe masses on TTF
- Current projection to freeze LZ in O(month)

[Watch Xe freeze in Crystalize!](#)



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Me & the smallest vessel :)

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 - Unbiased & improved resolution
- Cut in reconstructed N_e space reduces S2-only analysis threshold excess by 25%
- Enables a future LZ S2-only search for light DM
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Discrete electron counting and CXe R&D contribute to the improvement in direct dark matter experiments sensitivities, through extracting more information from the data we have, to developing new detector techniques for the future!

Electron counting

- 1D ResNet CNN successfully discretely counts e^- 's
 - Unbiased & improved resolution
- Cut in reconstructed N_e space reduces S2-only analysis threshold excess by 25%
- Enables a future LZ S2-only search for light DM
 - Aim to set new limits for DM interactions for dark photons, ALPs, solar reflected and halo DM

CrystalLiZe

- CXe reduces R_n by 600x
 - 20x reduction required to hit neutrino fog
- TTF taken from 1 week → 18 hours
 - Pushing for speed and high VUV transparency
- Scaling studies and TPC development ongoing
- Paving the way toward existing and future dark matter searches being neutrino-limited

Discrete electron counting and CXe R&D contribute to the improvement in direct dark matter experiments sensitivities, through extracting more information from the data we have, to developing new detector techniques for the future!

Post LTA

- Returning to Liverpool within the next month
- Finalise Outer Detector studies (spatial neutron veto → implementation to analysis framework)
- Lead the S2-only analysis in the 1 to $5e^-$ regime ($< \text{GeV}/c^2$) for dark photons, ALPs, solar reflected and halo DM

THANK YOU

