

### Status of the NEWS-G experiment

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#### Dark Matter UK meeting November 16<sup>th</sup>, 2021, Rutherford Appleton Laboratory, UK



This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme under grant agreement 714893-ExclusiveHiggs and under Marie Skłodowska-Curie agreement 841261-DarkSphere, 895168-neutronSPHERE

## **Spherical Proportional Counter**

Electric field scales as  $1/r^2$ , volume divided in: "drift" and "amplification" regions Capacitance independent of size: low electronic noise  $\rightarrow$  single electron threshold

















## New Experiments With Spheres - Gas





#### NEWS-G Collaboration

- 5 countries
- 10 institutes
- ~40 collaborators
- Three underground laboratories
- SNOLAB
- Laboratoire Souterrain de Modane
- Boulby Underground Laboratory



## New Experiments With Spheres - Gas



0.0

0

2

4

Energy [keV]

6

8

Favourable quenching factor

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10

## W-value



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### W-value

#### Estimation of the ionisation quenching factor in gases from W-value measurements

I. Katsioulas,<sup>1,\*</sup> P. Knights,<sup>1,2</sup> and K. Nikolopoulos<sup>1</sup>

<sup>1</sup>School of Physics and Astronomy, University of Birmingham, B15 2TT, United Kingdom <sup>2</sup>IRFU, CEA, Universite Paris-Saclay, F-91191 Gif-sur-Yvette, France (Dated: May 5, 2021)

The effect of ionisation quenching for ions is critical for experiments relying on the measurement of low energy recoils, such as direct Dark Matter searches. We present ionisation quenching factor estimates over a range of energies for protons,  $\alpha$ -particles, and heavier ions in H<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>, Ar, CO<sub>2</sub>, and C<sub>3</sub>H<sub>8</sub> gases, estimated from the respective reference W-value measurements. The resulting ionisation quenching factors are compared with predictions from SRIM.

 $q_f = \frac{E_e e}{E} = \frac{N_i^i \cdot W_e(E)}{E} = \frac{W_e(E)}{W_i(E)}$ 





## NEWS-G: Prototype at LSM



### Search for Solar Kaluza-Klein axions



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## NEWS-G at SNOLAB



Ø140 cm detector installed at SNOLAB

- Detector commissioning on-going
- Physics data-taking to follow



## Installation at SNOLAB





## Installation at SNOLAB







# Reducing Backgrounds



SNOLAB detector: 4N Aurubis AG Oxygen Free Cu (99.99% pure)
▶ Out-of-equilibrium <sup>210</sup>Pb contamination: 29±10 (stat)+9-3 mBq/kg



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Bremsstrahlung X-rays from <sup>210</sup>Pb and <sup>210</sup>Bi β-decays in Cu





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Internal shield

Ultra-pure Cu layer on detector inner surface

Suppresses <sup>210</sup>Pb and <sup>210</sup>Bi backgrounds by factor 2.6 under 1 keV





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### Electroformed Cuprum Manufacturing Experiment



EuME

A Ø140 cm sphere electroformed underground in SNOLAB

- Builds on achievements of NEWS-G electroplating
  - ≥ 36 µm/day  $\rightarrow$  ~1 mm/month
- No machining or welding grow sphere directly



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#### **Current Status**

- Ø30 cm scale prototype to be produced at PNNL
  - Bath designed
  - Procurement of parts underway
  - Electroformation to start soon
- ø140 cm detector to follow shortly after
  - Use existing shielding for physics exploitation







Single anode: Drift and Amplification fields are connected

$$E = \frac{V_a}{r^2} \frac{r_a r_c}{r_c - r_a} \approx \frac{V_a r_a}{r^2}$$



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- **ACHINOS: Multi-anode sensor** JINST 12 (2017) 12, P12031
  - Multiple anodes placed at equal radii
  - Decoupling drift and amplification fields
  - Opportunity: individual anode read-out



Single anode: Drift and Amplification fields are connected  $E = \frac{V_a}{r^2} \frac{r_a r_c}{r_c - r_a} \approx \frac{V_a r_a}{r^2}$ 

3D printed ACHINOS with DLC coating



JINST 15 (2020) 11, 11



Aχινός (greek. sea urchin)

Insulated wires
Support rod

ACHINOS: Multi-anode sensor

Multiple anodes placed at equal radii

Decoupling drift and amplification fields

Opportunity: individual anode read-out



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## ACHINOS performance with DLC coating



- Good energy resolution
- High pressure operation
- High gain
- Stability
- 2 channel read-out



Measurement of the 5.9 keV <sup>55</sup>Fe X-ray line



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## Fiducialisation

Birmingham simulation framework, combining strengths of Geant4 and Garfield++



Reading out individual ACHINOS anodes: position of interaction can be reconstructed

- First tests: Separate the anodes in two electrodes "Near" and "Far" (from the rod)
  - Asymmetry of pulse amplitudes: zenith angle
  - Pulse rise-time: radius



### Event reconstruction

Individual anode read-out: track reconstruction



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## R2D2: R&D towards 0vββ with SPC

- Rare Decays with Radial Detector (R2D2)
  - R&D effort for 0vββ searches with SPC
- Aim towards tonne-scale <sup>136</sup>Xe with 1% FWHM at Q<sub>ββ</sub>=2.458 MeV
- Recent and future work:
- Demonstrated energy resolution in Ar-prototype
- Light and charge read-out
- Demonstrate zero background
- Demonstrate large mass operation

#### R2D2 spherical TPC: first energy resolution results

#### R. Bouet,<sup>a</sup> J. Busto,<sup>b</sup> V. Cecchini,<sup>a,f</sup> C. Cerna,<sup>a</sup> A. Dastgheibi-Fard,<sup>c</sup> F. Druillole,<sup>a</sup> C. Jollet,<sup>c</sup> P. Hellmuth,<sup>a</sup> I. Katsioulas,<sup>d</sup> P. Knights,<sup>d,e</sup> I. Giomataris,<sup>e</sup> M. Gros,<sup>e</sup> P. Lautridou,<sup>f</sup> A. Meregaglia,<sup>a,\*</sup> X.F. Navick,<sup>e</sup> T. Neep,<sup>d</sup> K. Nikolopoulos,<sup>d</sup> F. Perrot,<sup>a</sup> F. Piquemal,<sup>a</sup> M. Roche, $^a$ B. Thomas, $^a$ R. Ward $^d$ and M. Zampaolo $^c$ <sup>a</sup>CENBG, Université de Bordeaux, CNRS/IN2P3, F-33175 Gradignan, France <sup>b</sup>CPPM, Université d'Aix-Marseille, CNRS/IN2P3, F-13288 Marseille, France <sup>c</sup>LPSC-LSM, CNRS/IN2P3, Université Grenoble-Alpes, F-73500 Modane, France <sup>d</sup>School of Physics and Astronomy, University of Birmingham, Birmingham, B15 2TT, U.K. e IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France <sup>f</sup> SUBATECH, IMT-Atlantique, Université de Nantes, CNRS-IN2P3, F-44307 Nantes, France E-mail: anselmo.meregaglia@cern.ch



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JINST 16 (2021) 03, P03012

## In-situ neutron measurements



#### Neutrons: critical background in DM searches

- Underground measurements scarce
- Nitrogen gas



### In-situ neutron measurements



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# Reaching the neutrino floor



## DarkSPHERE

Volume ×10: Ø300cm intact underground electroformed spherical proportional counter Shielding: Full water shielding option and water/lead option considered





5 bar He:C<sub>4</sub>H<sub>10</sub> (90%:10%) (27 kg target mass) Possibility to host DarkSPHERE at Boulby's Large Experimental Cavern



Science and Technology Facilities Council

- Possibility for a 7×7×7m<sup>3</sup> detector without further excavations
- Scoping study on possible electroformation facility on-going



## DarkSPHERE: Physics Potential



## Summary

NEWS-G has a rich R&D programme towards exploring new DM territory with Spherical Proportional Counters

- Significant instrumentation advances
  - ▶ Electroformation, ACHINOS, ...
- Several detectors scheduled/planned for the coming years
- Sensitivity down to the neutrino floor
- Many physics opportunities: DM Nuclear (spin dependent and independent) and electron recoils, CEvNS, axions, ...

11 spherical metallic anodes

Insulated wires

Support rod

15°

0°

-15°





