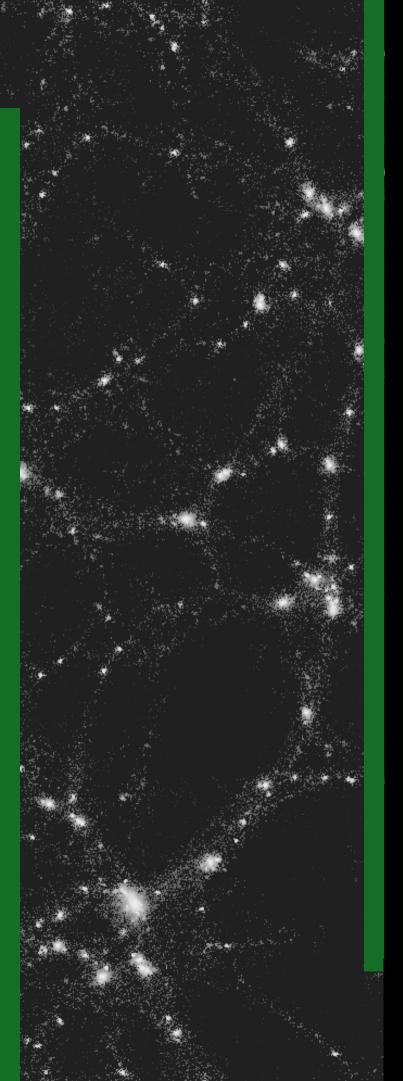
Fiona Alder DMUK 16/11/21

Direct Dark Matter Detection and Novel Directional Searches



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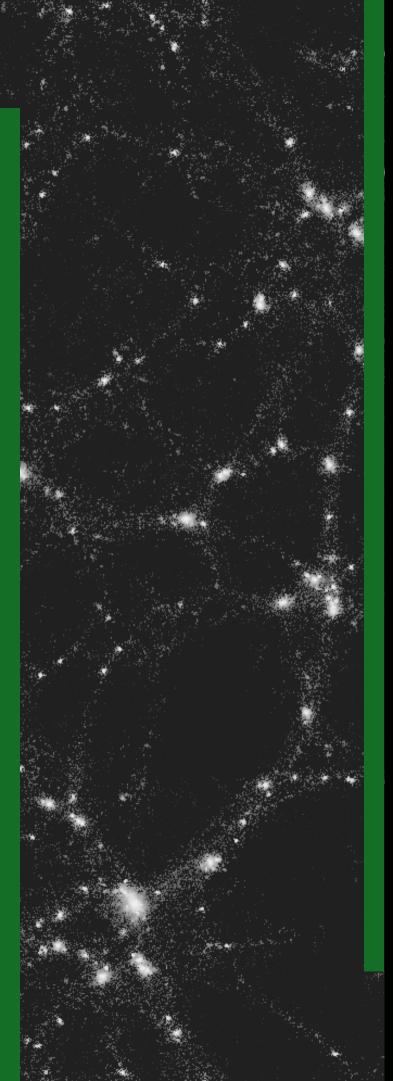
COLLABORATIVE WORK

HEP

Professor Chamkaur Ghag Robert James (PhD) Fiona Alder (PhD)

AMOPP

Professor Peter Barker Jonathan Gosling (PhD)



OUTLINE

Software package creation

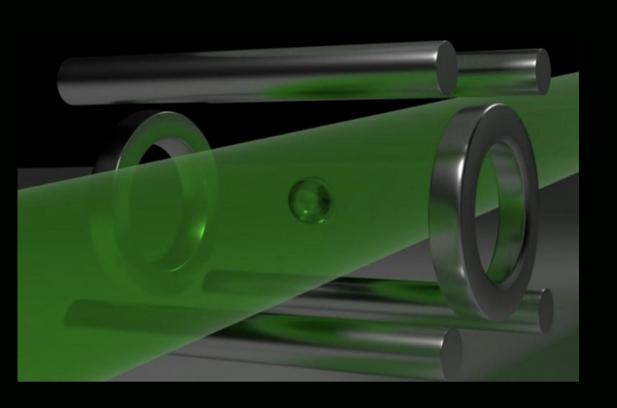
Optical tweezer experimental design and construction

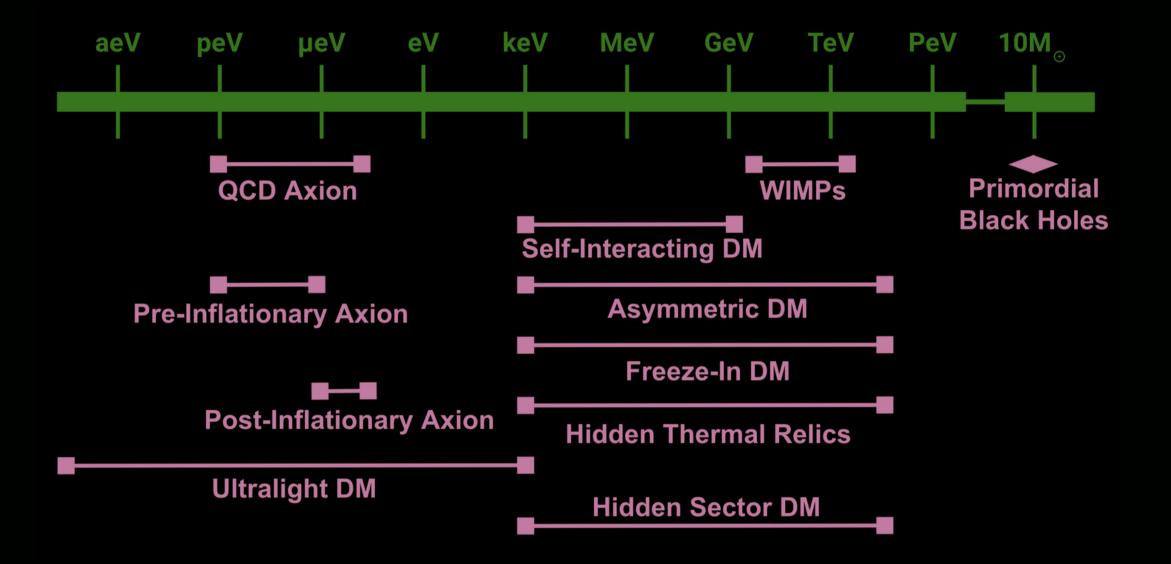
Paul trap experimental design

Background characterisation and reduction

Setting initial limits

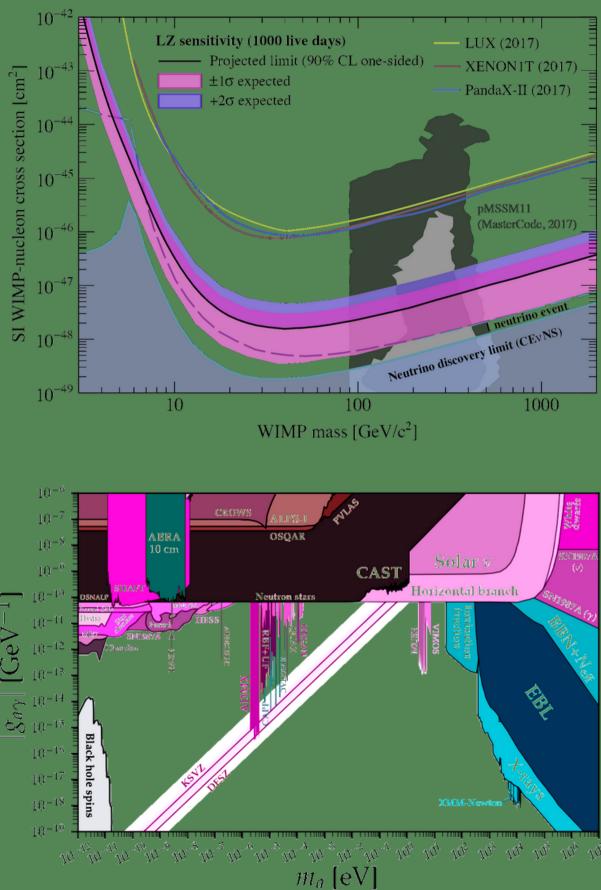
Future searches



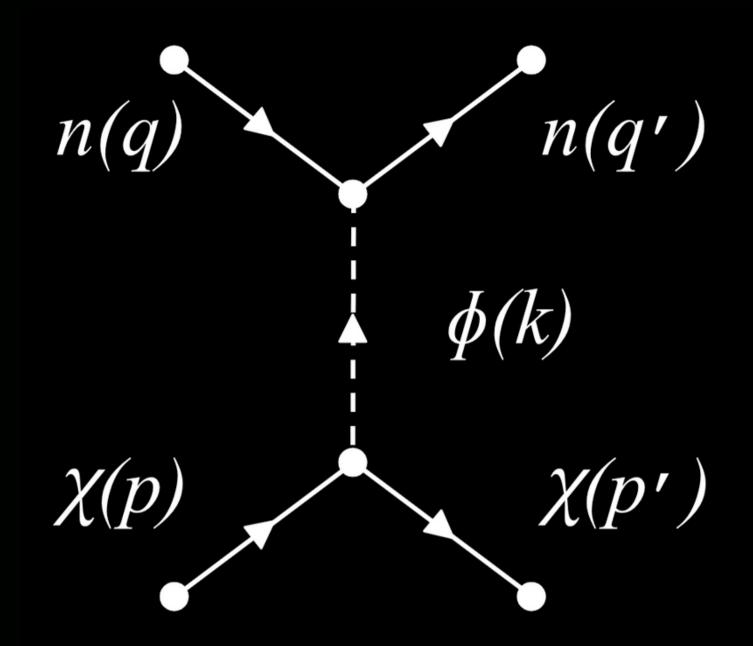


DARK MATTER CANDIDATES

DIRECT DETECTION LIMITS



₩₩ [♥ ♥]



SELF INTERACTING DM

$\mathscr{L} \supset -g_{\chi}\phi\chi^*\chi - g_n\phi\overline{n}n$

- One theoretically motivated candidate for self-interacting DM is a class of model known as DM 'nuggets'
- Contain a number of constituents, N_{y}
- Interact with standard model particles via a light mediator, ϕ (mass m_{ϕ} ≤ eV)

 $V(r) = (-) \underbrace{g_{\chi} g_n e^{-rm_{\phi}}}_{\chi}$

1 2 3 SIMULATIONS CALIBRATION AQ AND DATA STREAMING

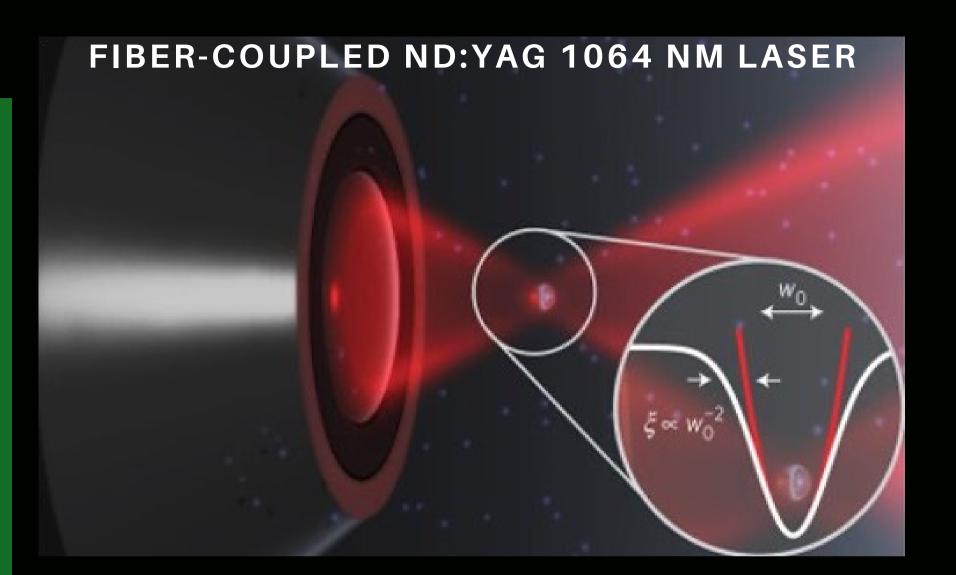
QS PACKAGE

4

PUSLE FINDING AND CATEGORISATION



STATISTICAL ANALYSIS



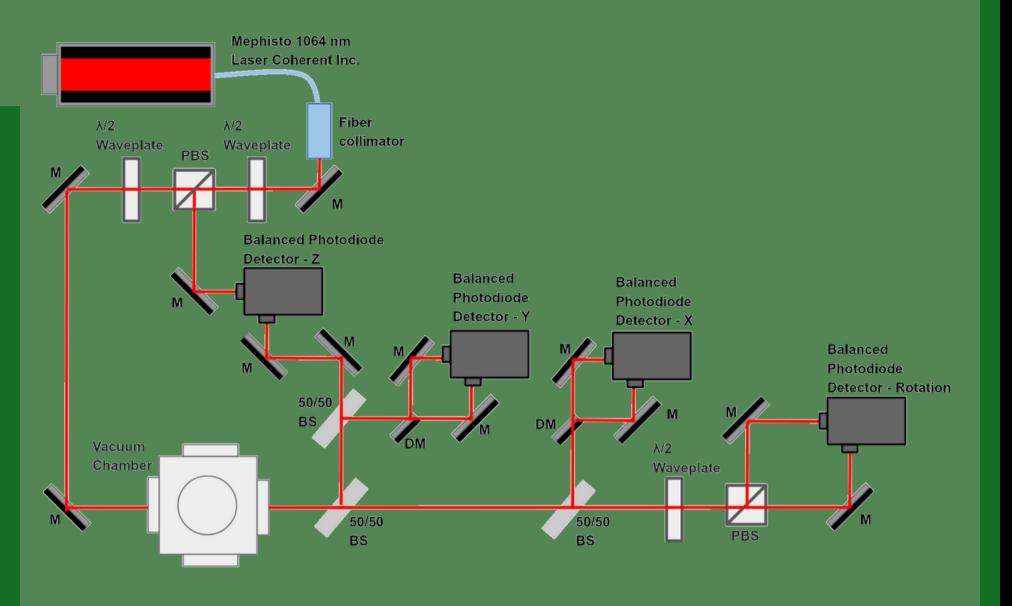
 $m\ddot{q} = -m\gamma\dot{q} + F_{Gr} + F_{Dr} + F_{fluct}$ Gradient Force Fluctuating Force Driving Force

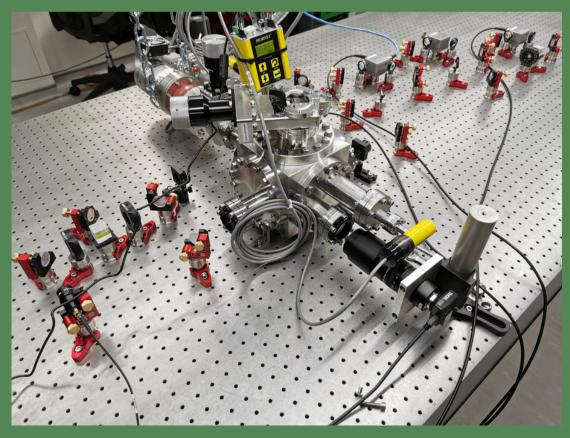
Trapping potential formed through focusing of EM fields to the focal point of an optical lens

Interference of the incident light rays occurs at this focal point

 $\langle \vec{\mathsf{F}} \rangle = \frac{\alpha'}{2} \langle \nabla | \vec{\mathsf{E}} |^2 \rangle$

OPTICAL TWEEZER THEORY





EXPERIMENTAL REQUIREMENTS

Trapping Optics

Vacuum and Optical Trap

Detection System

TRAPPING AND DETECTION

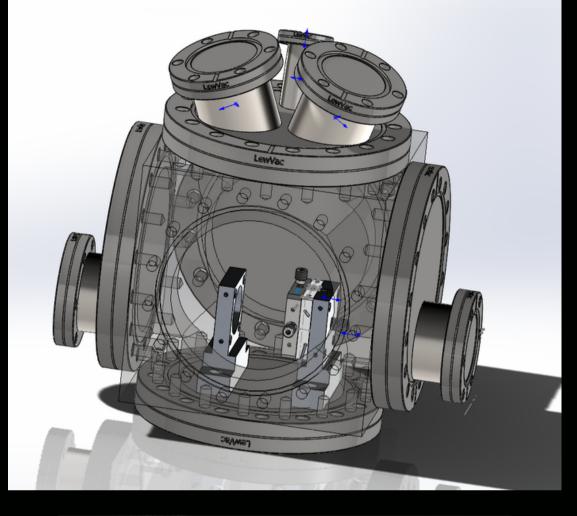
Reduction of cantilever movement

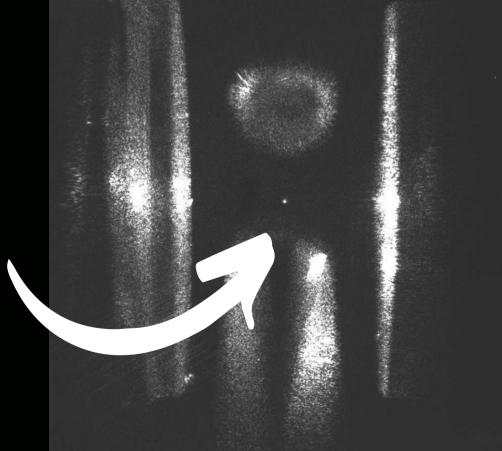
Laser power controlled

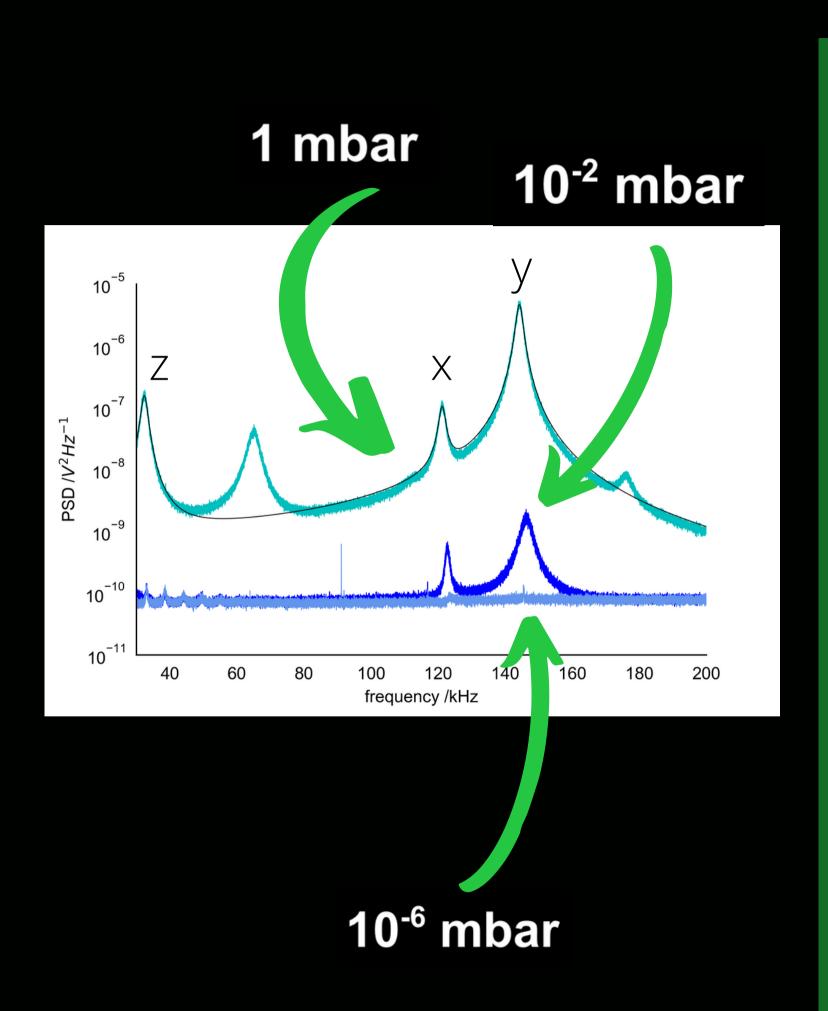
Feedback cooling and calibration

3D directional reconstruction

Rotational detection







- "velocity damping"

- driving force

• Cooling through active feedback mechanism

Shown to cool to the quantum ground state*

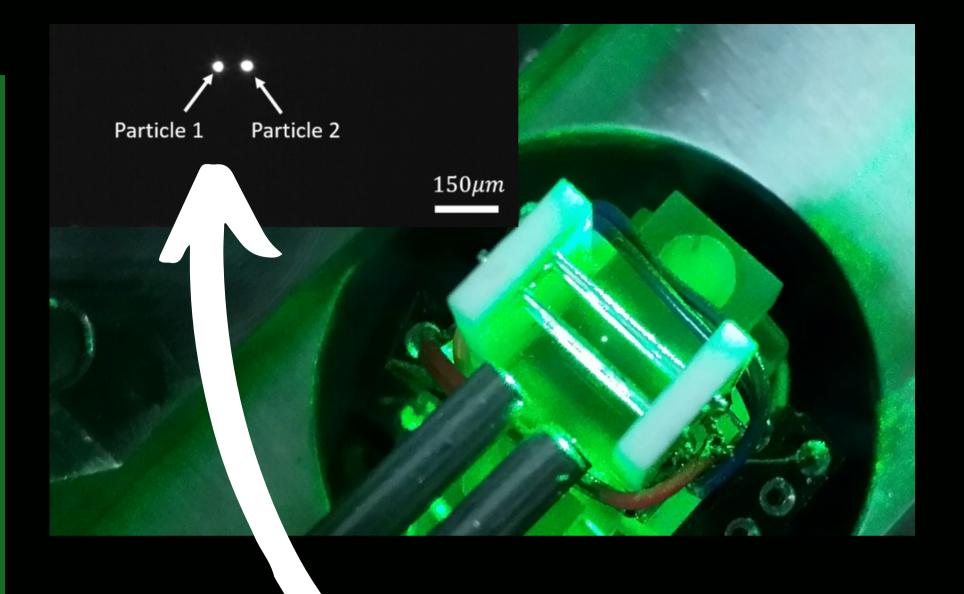
More effective than other methods of cooling**

• Calibration through application of harmonic

• Charging and discharging of particle through application of xenon UV flash lamp

COOLING AND CALIBRATION

*F. Tebbenjohanns et al., Phys. Rev. Lett., 2019 **A. Pontin, T. W. Penny and P.F. Barker, Phys. Rev. A, 2021



Co-levitated nanoparticles

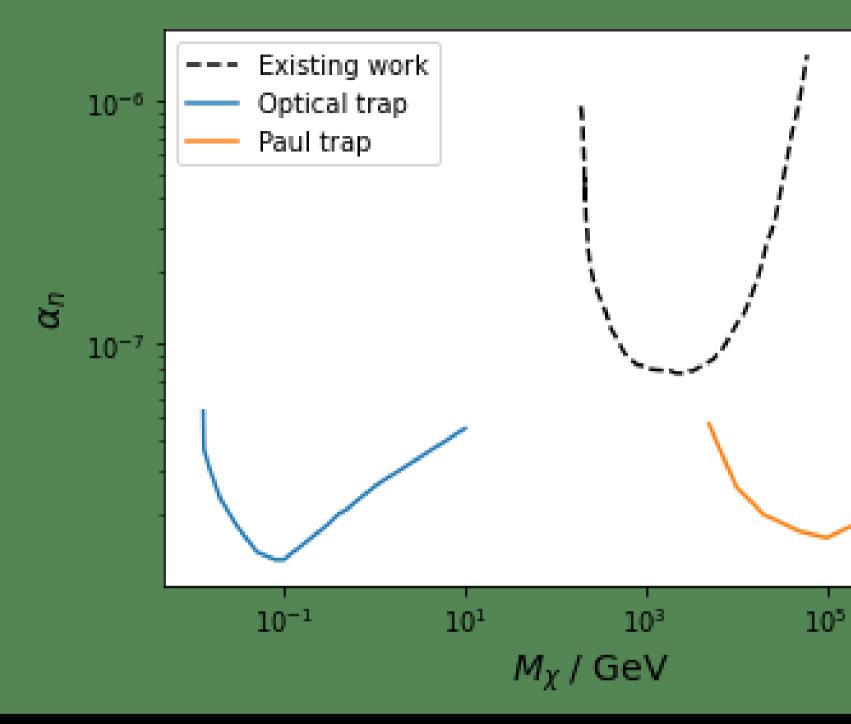
*T. W. Penny et al, Sympathetic cooling and squeezing of two co-levitated nanoparticles, 2021

Particles O(100µm) in diameter have been trapped in vacuum using this method

Trapping of metallic particles of much higher density is possible, allowing probing of heavier dark nugget parameter space

Levitated particle arrays cooled to mK temperatures using sympathetic feedback cooling*

LINEAR PAUL TRAP

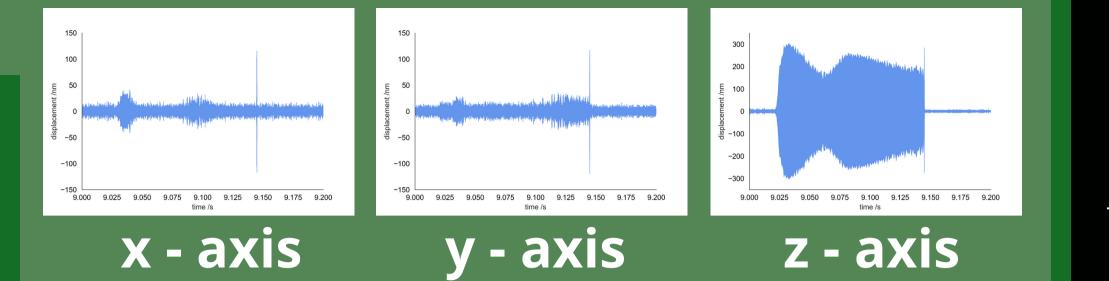


PROJECTED 90% DARK NUGGET SCATTERING SENSITIVITY

 10^{7}

Background-free experiment with the observation of 0 events and a mediator mass of 0.1 eV

ROBERT JAMES



IMPULSE LIKE INTERACTIONS

THERMAL NOISE

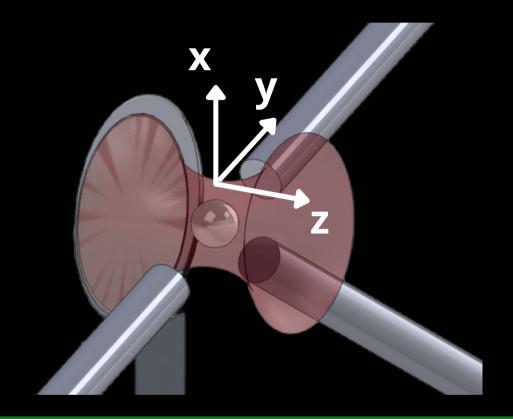
ELECTRONIC NOISE

SEISMIC NOISE

LASER NOISE

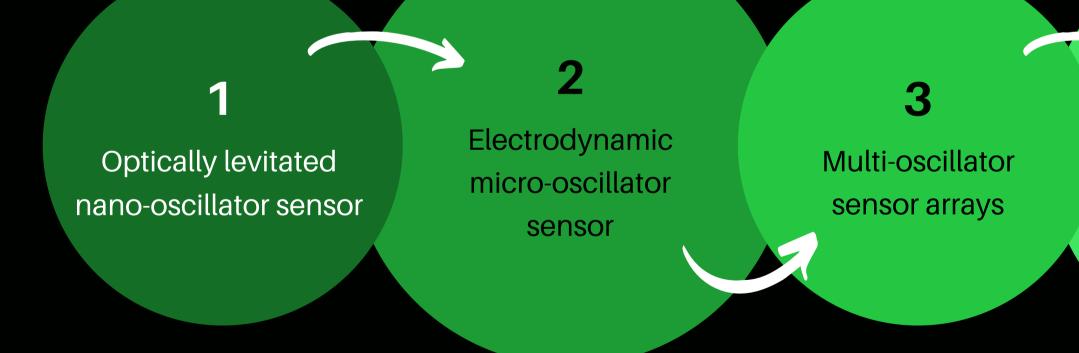
This will involve simulation code and pulse finding and categorisation

A background study is integral to the commissioning of the experimental setup



BACKGROUND CHARACTERISATION AND REDUCTION

PROJECT PLAN



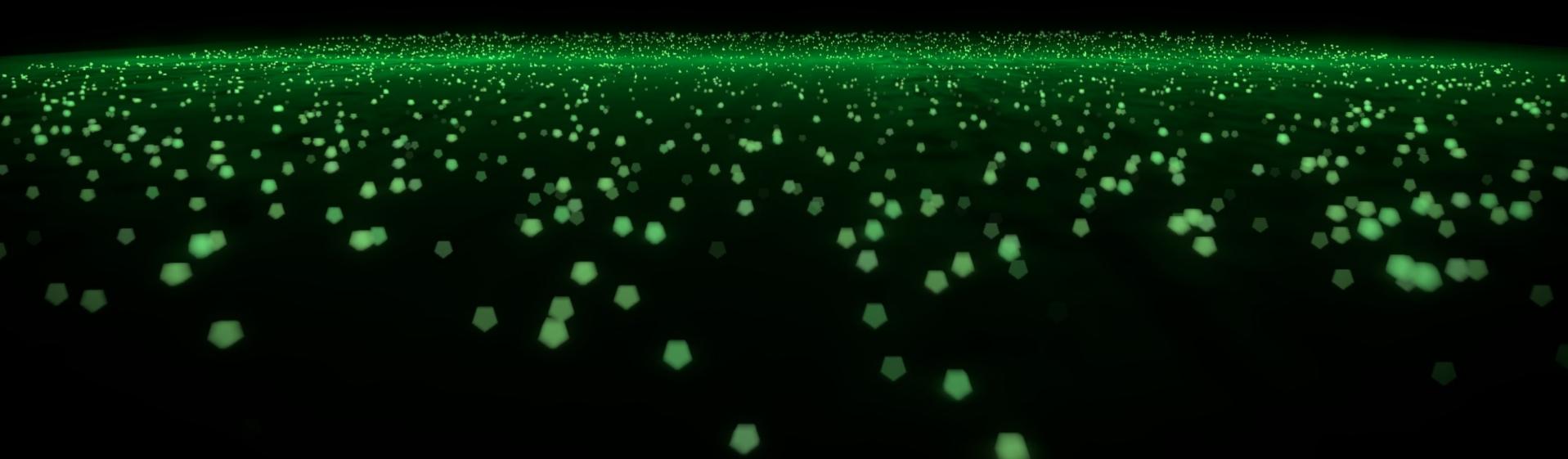


Ultralight dark matter detection with levitated oscillators

Further theoretical investigation

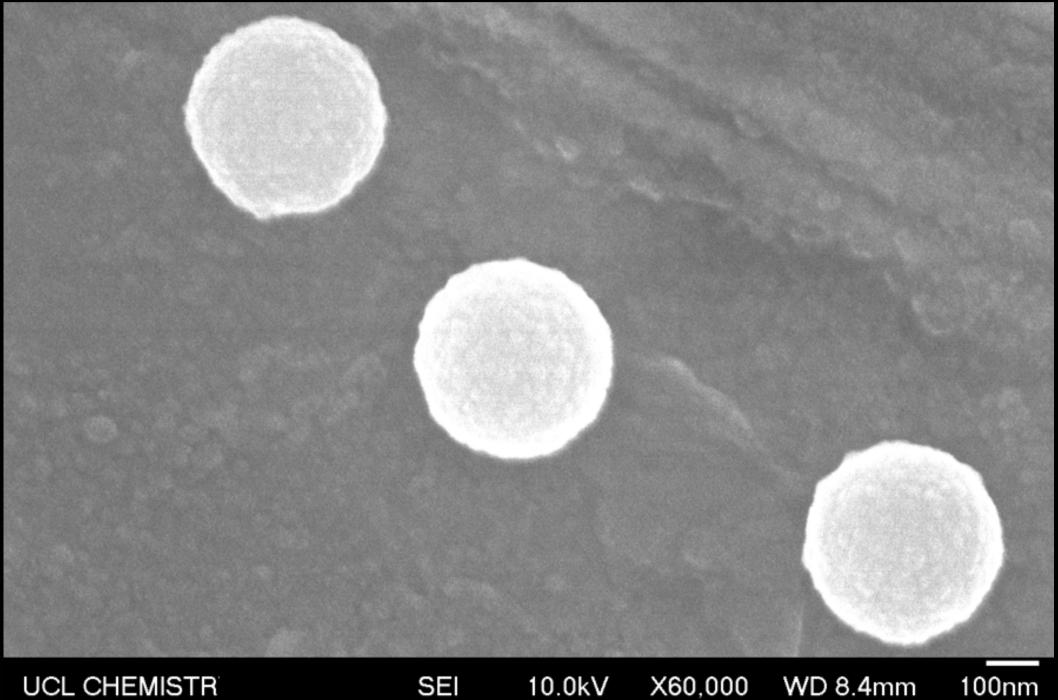
5

ADDITIONAL SLIDES





SILICA NANOPARTICLES



UCL CHEMISTR

SEI 10.0kV X60,000