



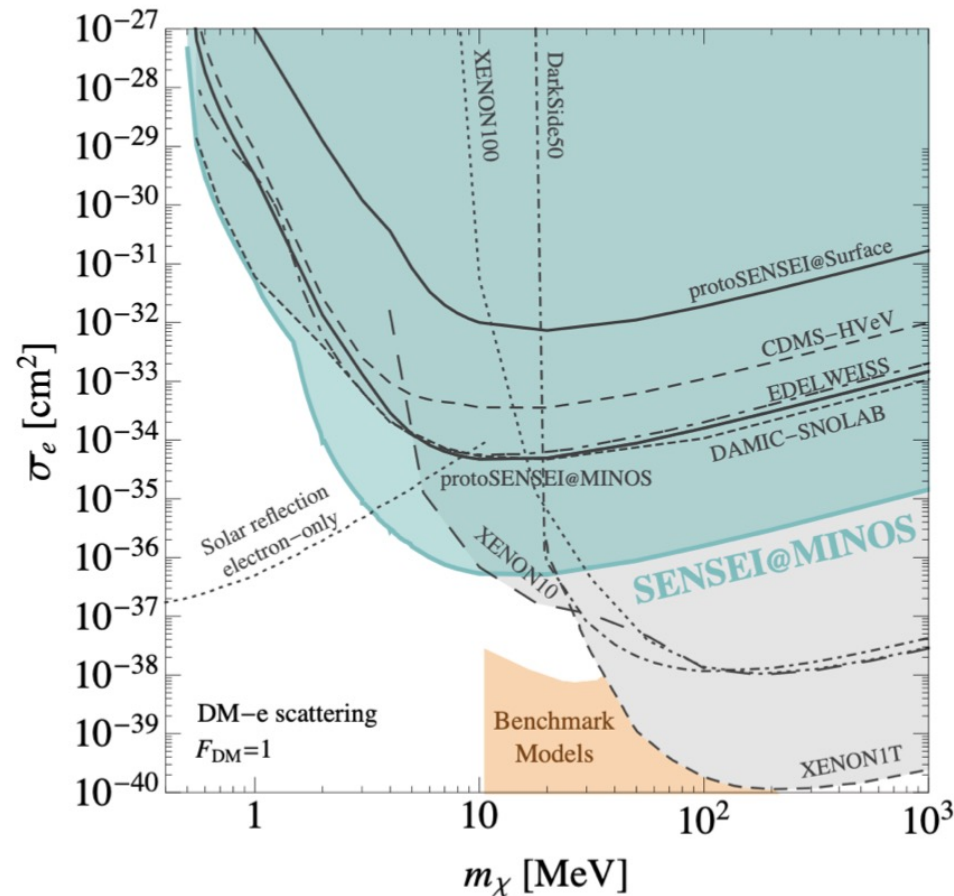
Fuelling the search for light DM-electron scattering

Louis Hamaide – DMUK Conference

Based on arxiv 2110.02985

Going To Lower DM Masses

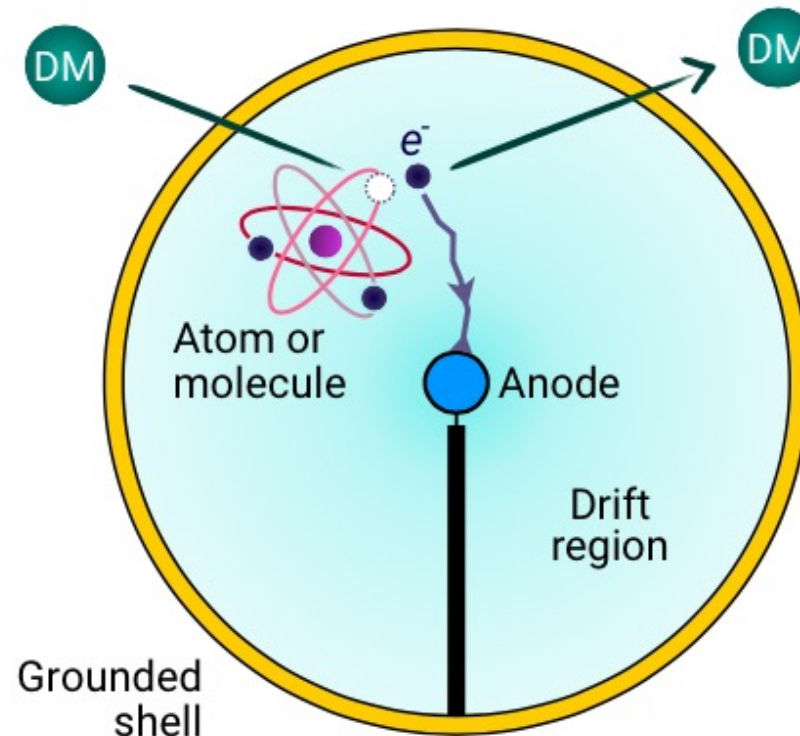
- DM-electron scattering opens searches for lighter DM (< 1 GeV), but requires sensitivity to small number of electrons ($< 4e^-$)



arxiv 2004.11378

What Are Spherical Proportional Counters

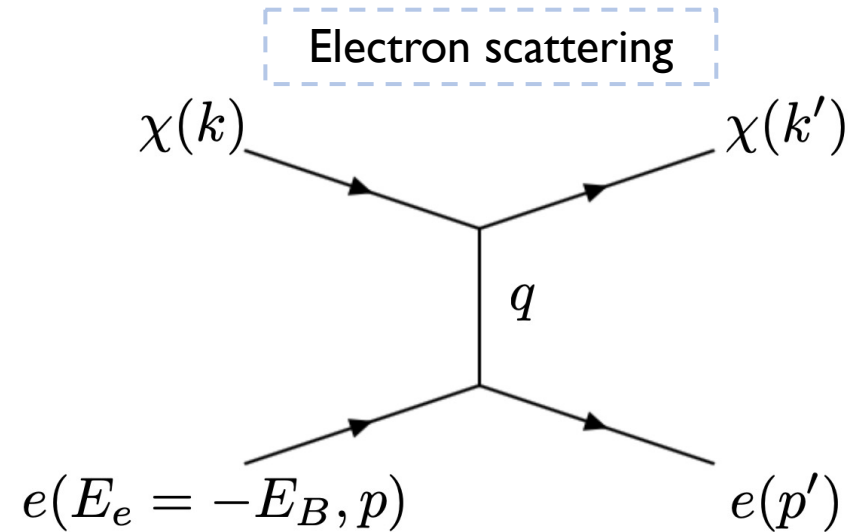
- SPCs consist of gas in a sphere sensitive to $1e^-$ events and exhibit low noise
- We study He and Ne (lighter), Xe (for comparison), CH_4 and C_4H_{10} (quenchers)



Ionization
electrons drift to
the anode where
the initial energy
can be
reconstructed

Dark Matter Electron Scattering

- Dark matter transmits some energy to electron, ionizing atom/molecule
 - Scattering rate depends on electronic bound and unbound wavefunction
- Need to solve N-body atomic/molecular Hamiltonian



$$\mathcal{M}(n, l \rightarrow \text{free } e^-) = g_x g_y \frac{1}{q^2 - m^2} \int d^3x \tilde{\psi}_{p'l'm'}^*(\mathbf{x}) e^{i\mathbf{q}\cdot\mathbf{x}} \psi_{nlm}(\mathbf{x})$$

where

$$dR \propto |\mathcal{M}|^2$$

$F_{\text{DM}}(q)$

$f_{\text{ion}}^{nl}(q)$

Bound Electron Wavefunctions

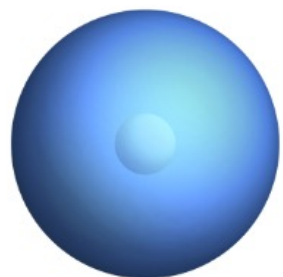
- Hartree-Fock approximation: mean field self consistent bound states:

$$\begin{aligned} -\frac{1}{2} \frac{d^2 P_{n_a l_a}}{dr^2} + \frac{l_a(l_a + 1)}{2r^2} P_{n_a l_a}(r) - \frac{Z}{r} P_{n_a l_a}(r) + \sum_{n_b l_b} (4l_b + 2) \left(v_0(n_b l_b, r) P_{n_a l_a}(r) - \sum_l \Lambda_{l_a l_b} v_l(n_b l_b, n_a l_a, r) P_{n_b l_b}(r) \right) \\ = \epsilon_{n_a l_a} P_{n_a l_a}(r) + \sum_{n_b \neq n_a} \epsilon_{n_a l_a, n_b l_a} P_{n_b l_a}(r) \end{aligned}$$

- **PySCF** : Quantum chemistry package that solves HF eqs. for atoms and molecules:
 - HF equations solved self-consistently using gaussian basis
 - Includes relativistic treatments, molecular dipoles, and more
 - Expect accuracy of O(30%) in event rates/bounds



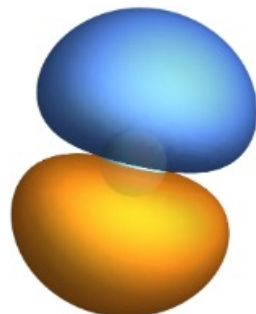
Bound Electron Wavefunctions - Results



1s orbital

HELIUM

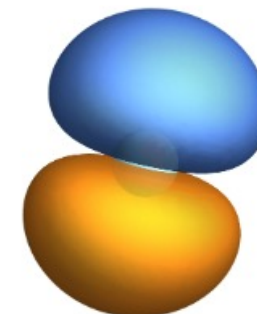
$E_{\text{ion}} = -24.6 \text{ eV}$
→ events from 2
electrons



2p orbital

NEON

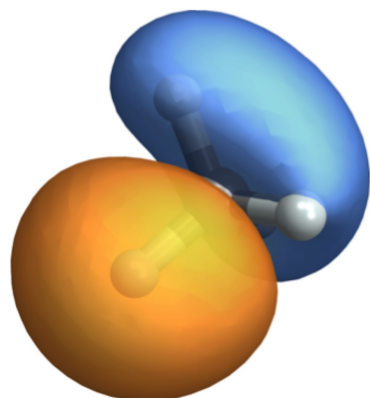
$E_{\text{ion}} = -21.7 \text{ eV}$
→ events from 8
electrons



5p orbital

XENON

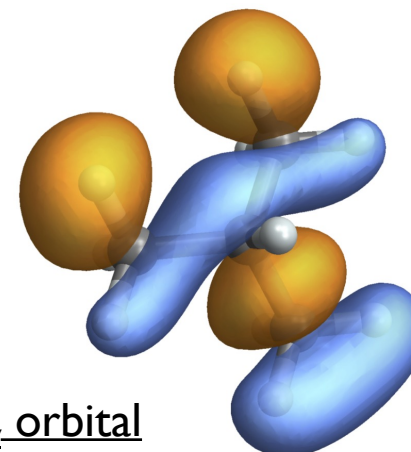
$E_{\text{ion}} = -12.7 \text{ eV}$
→ events from 18
electrons



2t_{2x} orbital

METHANE (CH₄)

$E_{\text{ion}} = -13.6 \text{ eV}$
→ events from 8
electrons



5e_y orbital

ISOBUTHANE
(C₄H₁₀)

$E_{\text{ion}} = -11.1 \text{ eV}$
→ events from 26
electrons

Unbound Electron Wavefunctions

- **Atoms:** Continuum limit: Hartree-Fock integrated, approximate the self-consistent piece of the potential + use frozen core:

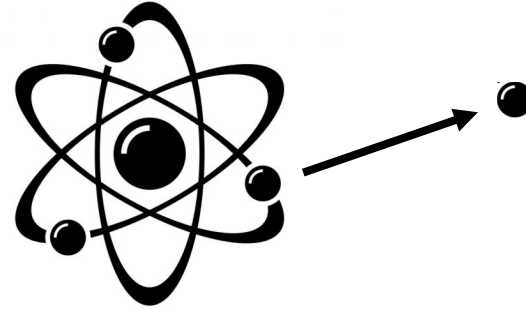
$$-\sum_{i \neq j} \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|} = V_{dir}(r) + V_{exch}(r)$$

$$V_{dir}(r) = \sum_{n_b l_b} (4l_b + 2) \int_0^\infty \left(\frac{P_{n_b l_b}^2(r_1)}{\max(r_1, r)} \right) dr_1$$

$$V_{exch} = k_x \left(\frac{24\rho(r)}{\pi} \right)^{1/3}$$

- **Molecules:** Coulomb potential/wavefunction:

$$P_{kl}(r) = \frac{4\pi}{2k} \frac{|\Gamma(\ell + 1 - \frac{iZ}{k})| e^{\frac{\pi Z}{2k}}}{(2\ell + 1)!} (2kr)^{\ell+1} \times e^{-ikr} M\left(\ell + 1 + \frac{iZ}{k}, 2\ell + 2; 2ikr\right)$$

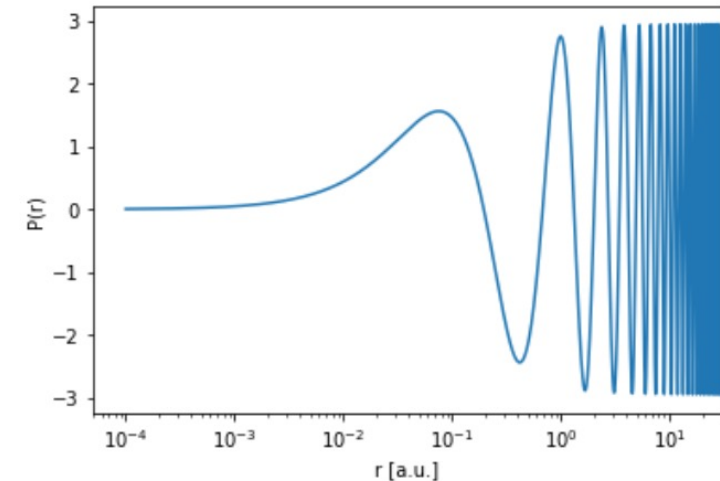


Potential:

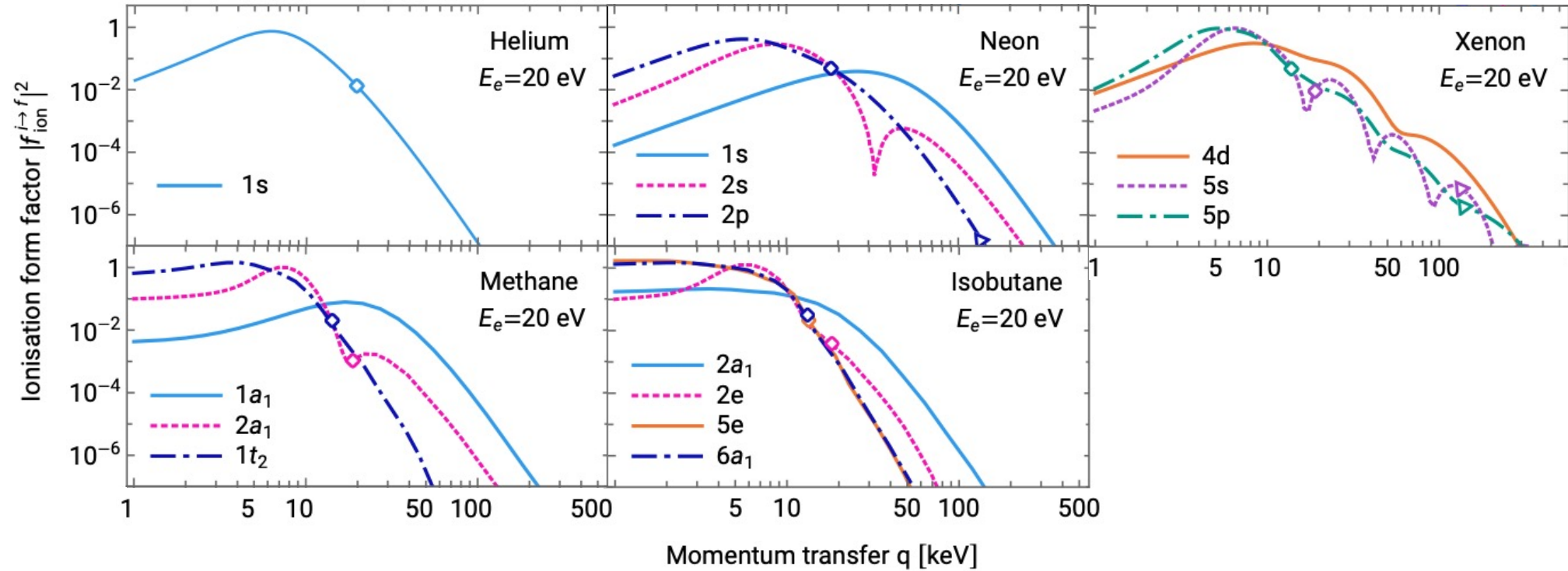
- Point nucleus : $-Z/r$
- Electron "cloud": $V_{dir}(r)$
- Electron "exchange": V_{exch}

→ $1/r$ at large r

Neon 2p → $l_R=0, E = 250\text{eV}$



Form Factors

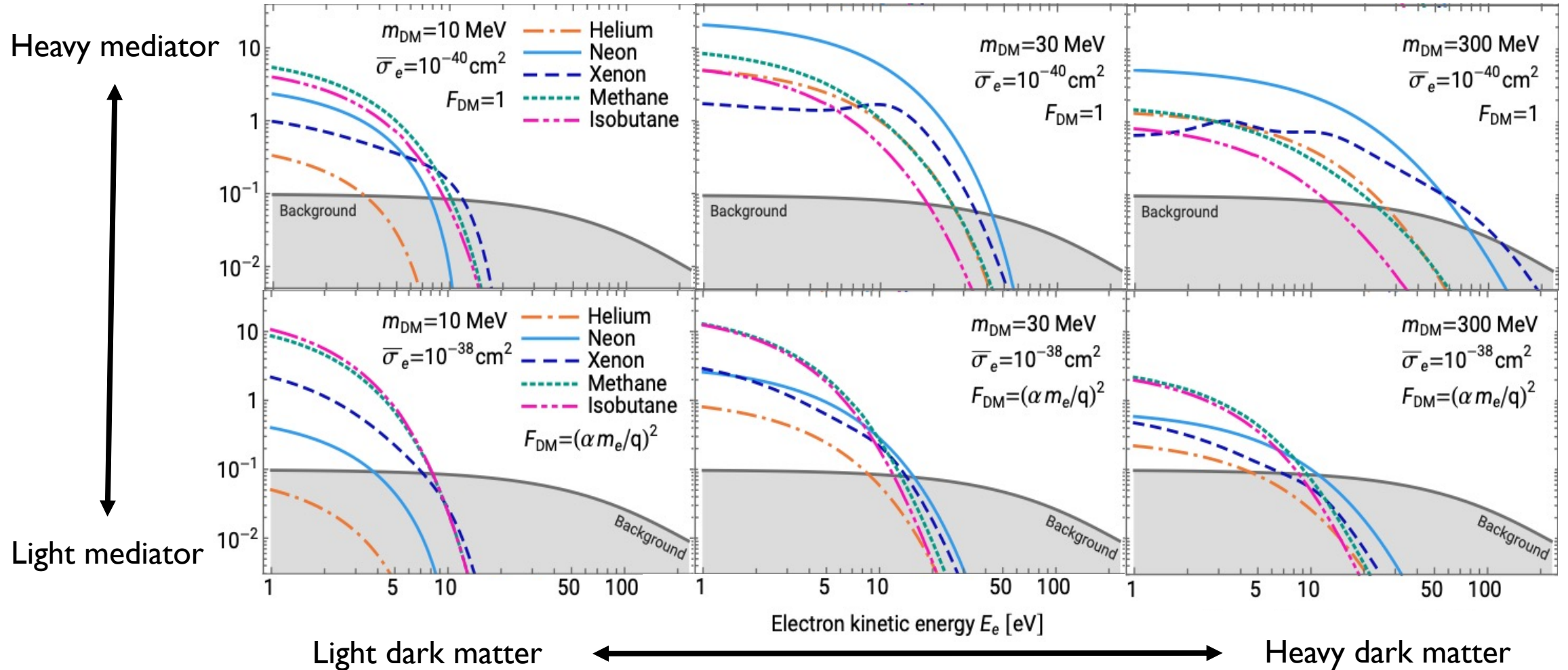


Form factor ($\alpha |\mathcal{M}|^2$):

$$|f_{\text{ion}}^{nl}(q)|^2 = \frac{(2m_e E_R)^{3/2}}{\pi^2} \sum_{\text{states:i,f}} \left| \int d^3x \tilde{\psi}_{p'l'm'}^*(\mathbf{x}) e^{i\mathbf{q}\cdot\mathbf{x}} \psi_{nlm}(\mathbf{x}) \right|^2$$

Event Rates

arxiv 2110.02985

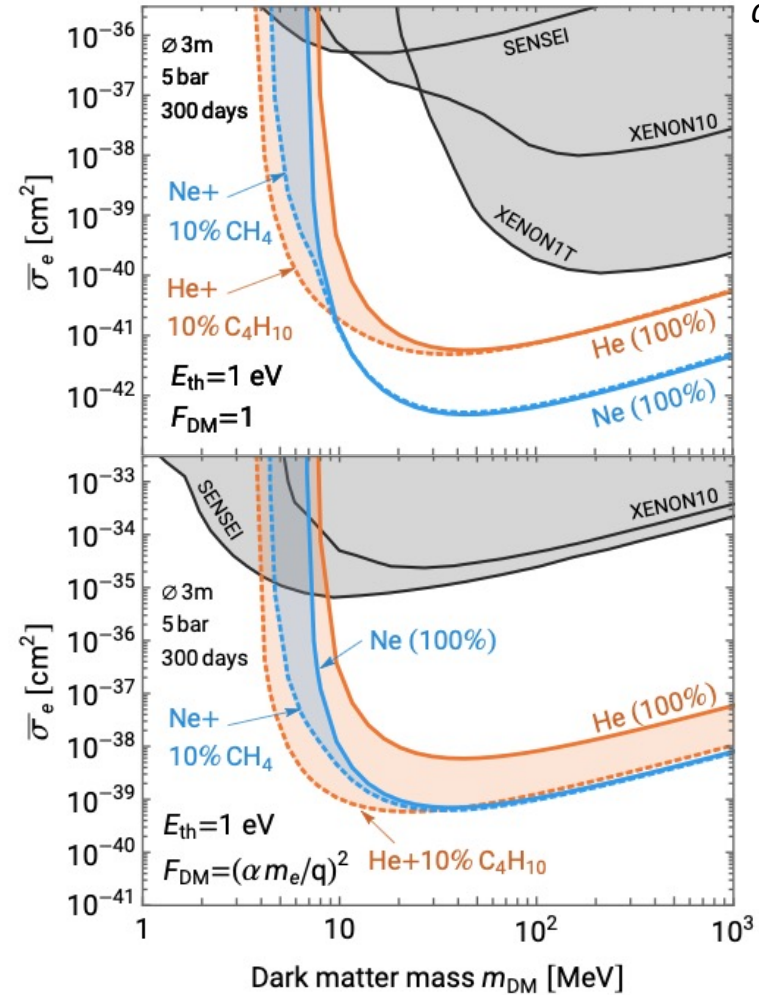
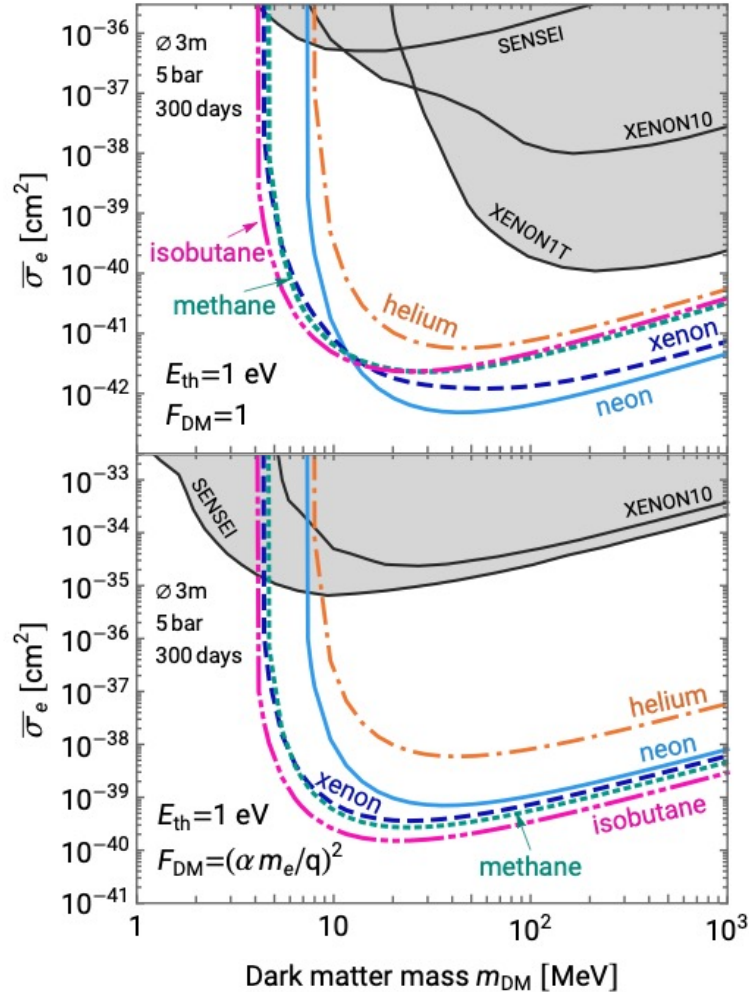


Sensitivities

Heavy mediator



Light mediator



arxiv 2110.02985

Summary

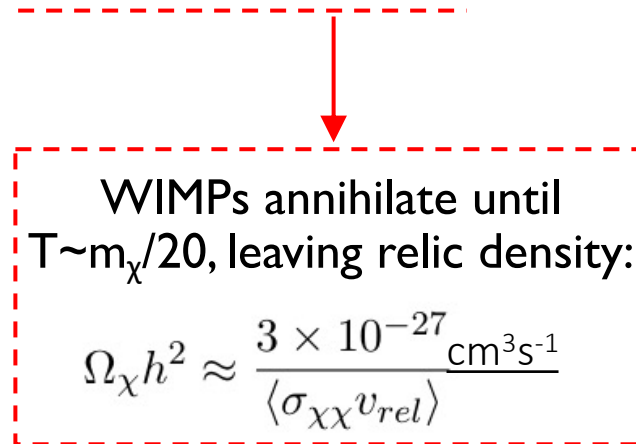
- Dedicated direct detection of DM-electron scattering good probe of light DM ($< 1 \text{ GeV}$)
- Atomic calculation under control – good accuracy & understanding (vs other HF & experiment)
- Molecular calculation more difficult (Coulomb approximation used), but can confidently be used to set bounds (i.e. with mixing)
- Seems promising \rightarrow more sensitive than current bounds / comparable to other proposed experiments
- SPC good probe of light DM-electron scattering !
- Large scale experimental proposal coming soon: DARKSPHERE

Thank you!

Back up - Motivation for Light(er) Dark Matter

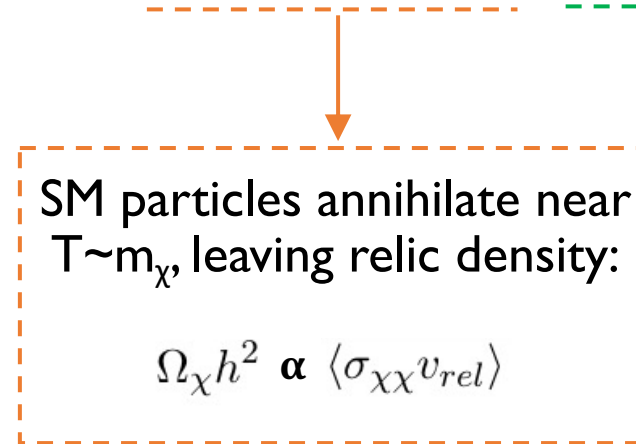
- Dark matter has many ways of appearing in the present day universe:

Thermal Production (e.g. freeze-out WIMP), Freeze-in (FIMP), Inflaton decay, Gravitino...



WIMPs annihilate until $T \sim m_\chi/20$, leaving relic density:

$$\Omega_\chi h^2 \approx \frac{3 \times 10^{-27} \text{ cm}^3 \text{ s}^{-1}}{\langle \sigma_{\chi\chi} v_{rel} \rangle}$$



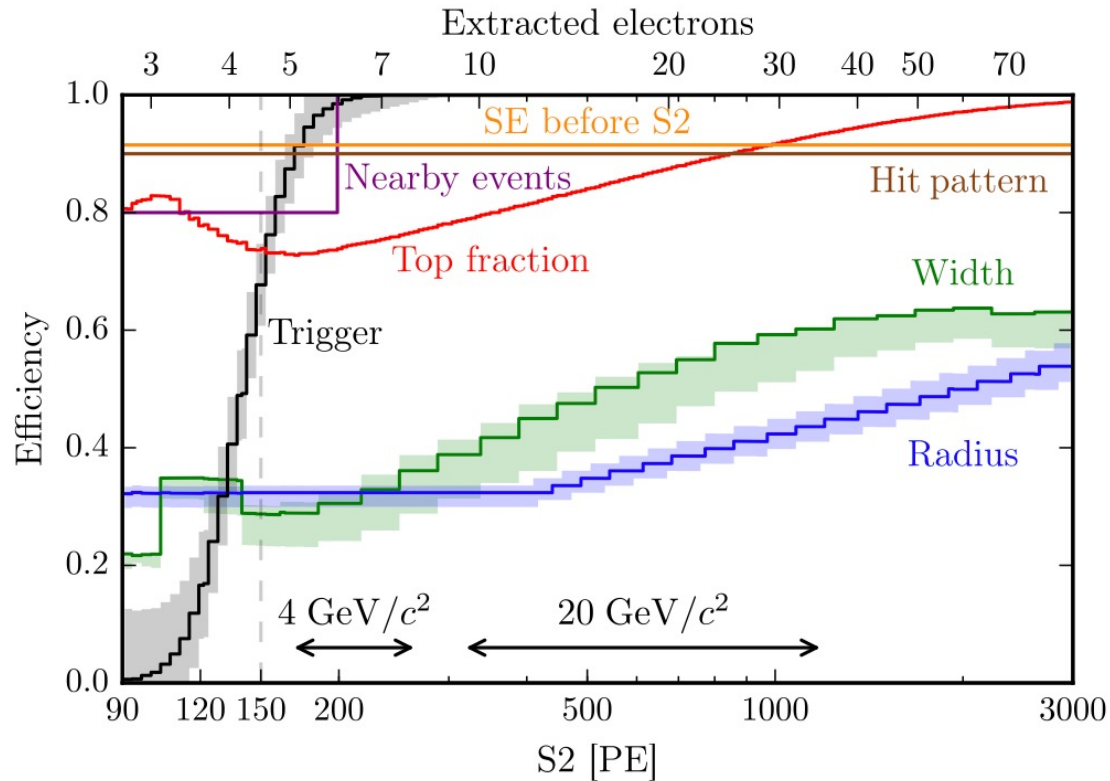
SM particles annihilate near $T \sim m_\chi$, leaving relic density:

$$\Omega_\chi h^2 \propto \langle \sigma_{\chi\chi} v_{rel} \rangle$$

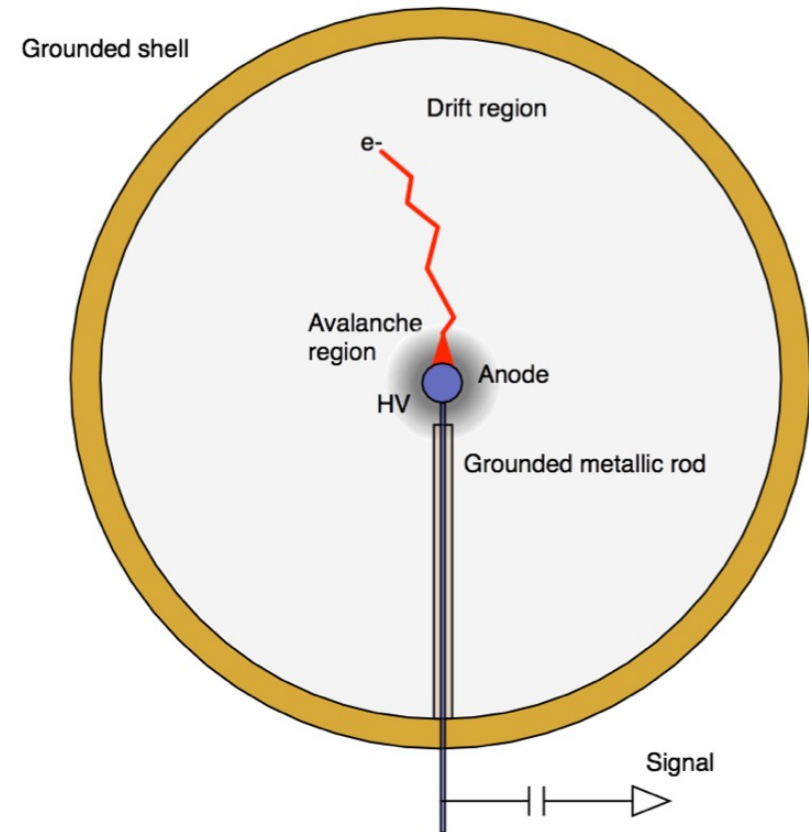
- Individual models of freeze-out and or freeze-in DM can be fully tested (even for unknown details of UV cosmology).
- Without knowledge of T_R , we cannot fully test inflaton decay or gravitino.

Back up – More On Detector

(liquid) Xenon 1T threshold ($>4.5e^-$)

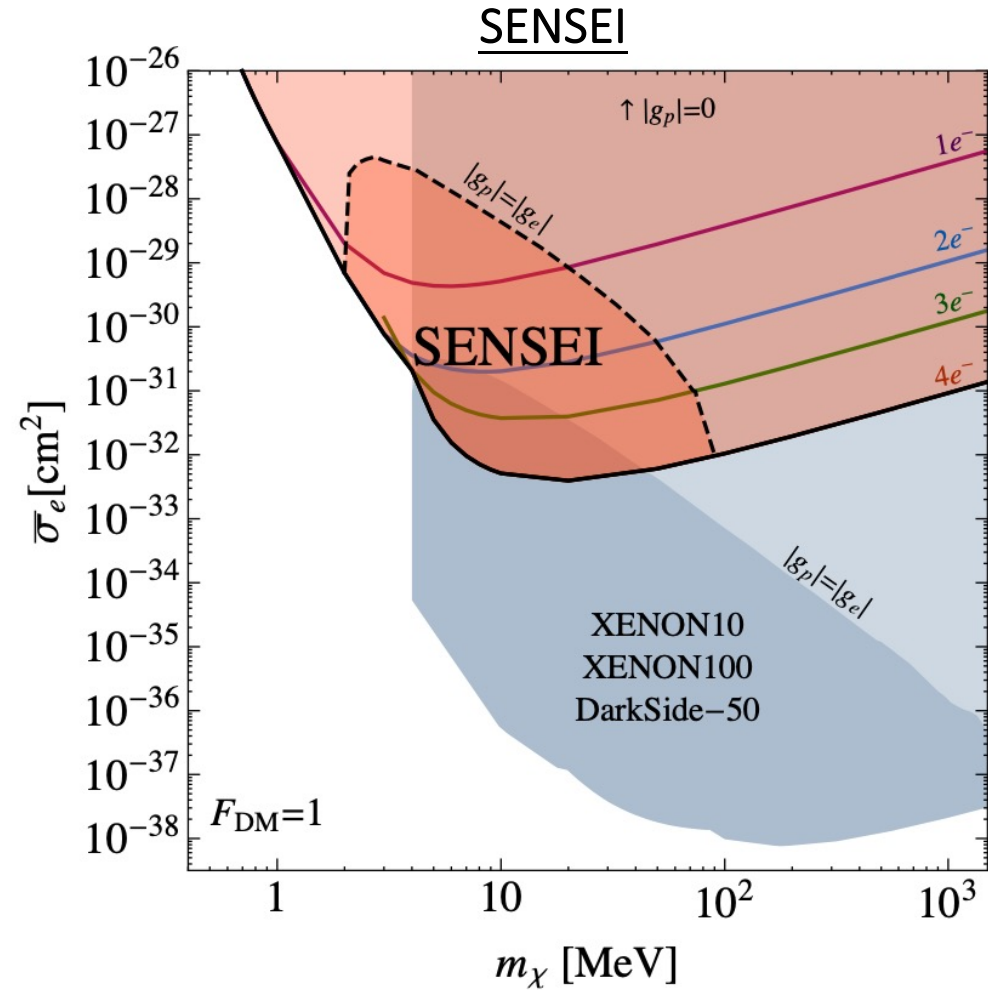
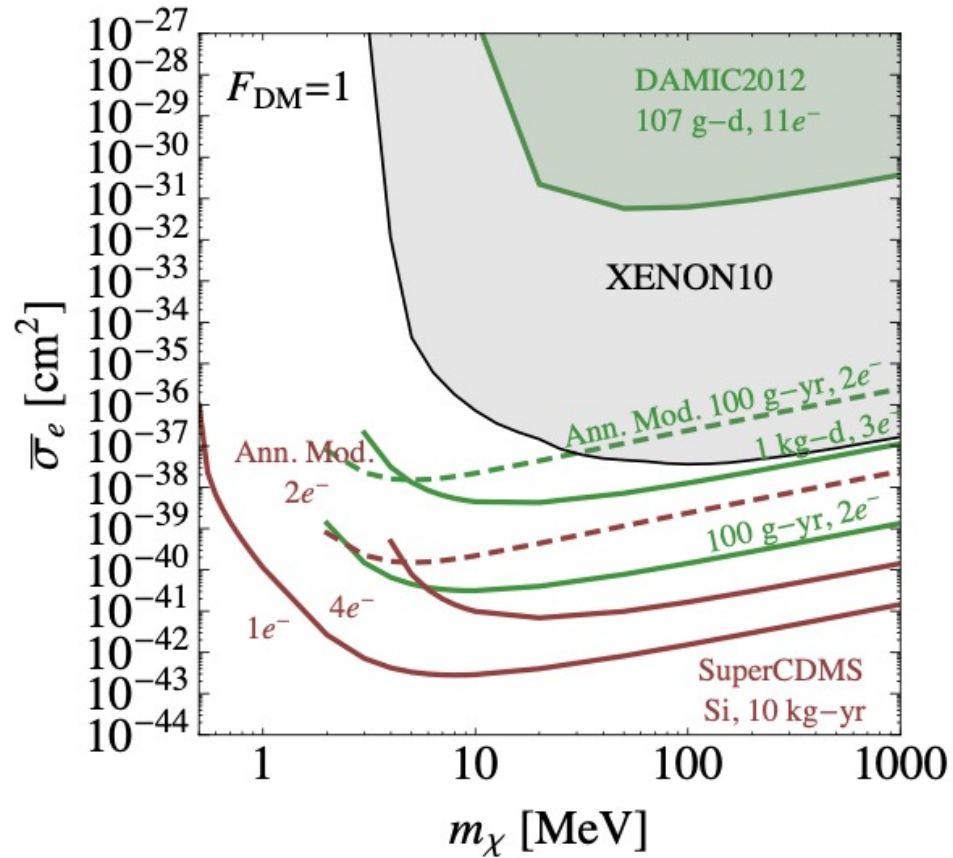


Spherical Proportional Counter (SPC, as proposed in DarkSPHERE)



Back up – More On Constraints

Electron Recoil SuperCDMS projection (Essig et al.)



Back up – Hartree Fock choice

- Hartree-Fock approximation: self-consistent bound states with energies correct to first order:

$$\begin{aligned}
 & -\frac{1}{2} \frac{d^2 P_{n_a l_a}}{dr^2} + \frac{l_a(l_a + 1)}{2r^2} P_{n_a l_a}(r) - \frac{Z}{r} P_{n_a l_a}(r) + \\
 & \sum_{n_b l_b} (4l_b + 2) \left(v_0(n_b l_b, r) P_{n_a l_a}(r) - \sum_l \Lambda_{l_a l_b} v_l(n_b l_b, n_a l_a, r) P_{n_b l_b}(r) \right) \\
 & = \epsilon_{n_a l_a} P_{n_a l_a}(r) + \sum_{n_b \neq n_a} \epsilon_{n_a l_a, n_b l_a} P_{n_b l_a}(r), \quad (3.71)
 \end{aligned}$$

Easy to solve
Hamiltonians for
full shell atoms

Non-Rel approx.
good for
 $q < 200\text{keV}$

Accuracy of
~30% in event
rates/bounds

- Sensitivities of bounds to choices:
 - ~30-50% Gaussian basis choice
 - ~50-100% exchange potential choice, orthogonalization
 - ~10-20% analysis of recoil energy profile vs. deposited energies
 - ~30% astrophysical parameter choices
 - Linear with background

Bound electron wavefunctions (3/4)

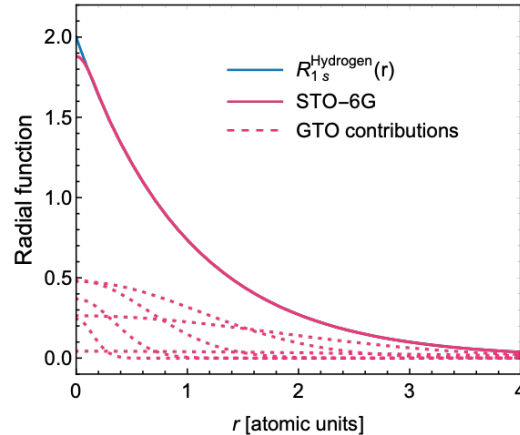
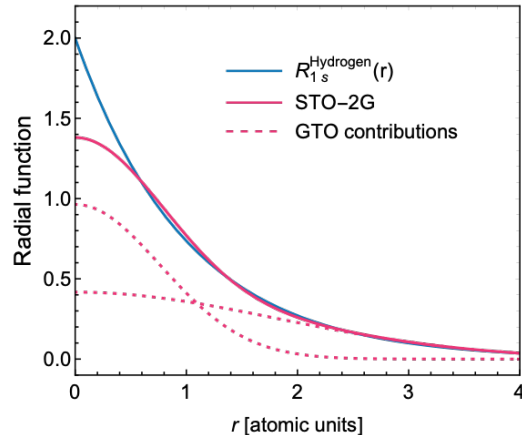
- Hartree-Fock approximation: mean field self consistent bound states:

$$-\frac{1}{2} \frac{d^2 P_{n_a l_a}}{dr^2} + \frac{l_a(l_a + 1)}{2r^2} P_{n_a l_a}(r) - \frac{Z}{r} P_{n_a l_a}(r) + \sum_{n_b l_b} (4l_b + 2) \left(v_0(n_b l_b, r) P_{n_a l_a}(r) - \sum_l \Lambda_{l_a l_b} v_l(n_b l_b, n_a l_a, r) P_{n_b l_b}(r) \right)$$

$$= \epsilon_{n_a l_a} P_{n_a l_a}(r) + \sum_{n_b \neq n_a} \epsilon_{n_a l_a, n_b l_a} P_{n_b l_a}(r)$$

Self-consistent approach required

Expect accuracy of O(30%) in event rates/bounds



Gaussian basis choice important at small/large r



$$AO_{nlm} \propto \sum_i c_i^n \exp(-\alpha_i^n r^2) Y_{lm}(\theta, \phi)$$

$$MO_{nlm} = \bar{v}_{nlm} AO_{nlm}$$

$$H_{ij} = \int AO_i(\mathbf{r}) H_j(\mathbf{r}) AO_j(\mathbf{r}) d^3 \mathbf{r} .$$

Treats atoms and molecules, relativistic treatments, molecular dipoles, and more.

Back up - Bound electron wavefunction symmetry

- Molecular orbitals are no longer eigenfunctions (spherical harmonics) of the SO(3) generators.

$$\psi(\mathbf{r}) = \frac{P(r)}{r} Y_{lm}(\theta, \phi) \rightarrow \psi(x, y, z)$$

- We need new classification of orbitals :

irreducible representations of SO(3) \rightarrow irreducible representation of point group

T_d	E	$8C_3$	$3C_2$	$6S_4$	$6\sigma_d$	linear functions, rotations	quadratic functions	cubic functions
A_1	+1	+1	+1	+1	+1	-	$x^2+y^2+z^2$	xyz
A_2	+1	+1	+1	-1	-1	-	-	-
E	+2	-1	+2	0	0	-	$(2z^2-x^2-y^2, x^2-y^2)$	-
T_1	+3	0	-1	+1	-1	(R_x, R_y, R_z)	-	$[x(z^2-y^2), y(z^2-x^2), z(x^2-y^2)]$
T_2	+3	0	-1	-1	+1	(x, y, z)	(xy, xz, yz)	$(x^3, y^3, z^3) [x(z^2+y^2), y(z^2+x^2), z(x^2+y^2)]$

Tetrahedral group (Methane): T_d

Back-up - bound electron energies

Helium (He)			Neon (Ne)			Methane (CH ₄)			Isobutane (C ₄ H ₁₀)			Xenon (Xe)		
Basis: aug-cc-pV5Z			Basis: aug-cc-pV5Z			Basis: 6-31G(d,p)			Basis: 6-31G(d,p)			Basis: Jorge-QZP		
Total energy: -2.8616			Total energy: -128.5467			Total energy: -40.2016			Total energy: -157.3123			Total energy: -7229.7195		
Orbital	I_{HF}	I_{exp}	Orbital	I_{HF}	I_{exp}	Orbital	I_{HF}	I_{exp}	Orbital	I_{HF}	I_{exp}	Orbital	I_{HF}	I_{exp}
1s ²	24.98	24.6	2p ⁶	23.14	21.7	1t ₂ ⁶	14.80	13.6	6a ₁ ²	12.34	11.13	5p ⁶	12.45	12.7
			2s ²	52.53	48.5	2a ₁ ²	25.66	22.9	5e ⁴	12.44	11.75	5s ²	25.54	23.3
			1s ²	891.79	870.2	1a ₁ ²	304.96	290.8	1a ₂ ²	13.86	12.85	4d ¹⁰	75.72	68.5
									4e ⁴	14.54	13.71	4p ⁶	163.56	146.1
									3e ⁴	16.04	15.03	4s ²	212.69	213.2
									5a ₁ ²	17.15	15.91	3d ¹⁰	711.26	682.7
									4a ₁ ²	20.62	18.58	3p ⁶	958.02	971.4
									2e ⁴	25.17	21.83	3s ²	1087.7	1149
									3a ₁ ²	29.44	24.83	2p ⁶	4839.8	4947
									2a ₁ ²	305.01	—	2s ²	5132.0	5453
									1e ⁴	305.01	—	1s ²	33321	34561
									1a ₁ ²	305.30	—			

Event rates – Back up

- The dark event rate can be calculated using:

$$\boxed{\frac{dR}{dE_e}} = \frac{1}{m_A} \frac{\rho_{\text{DM}}}{m_{\text{DM}}} \sum_{nl} w_{nl} \frac{d\langle \sigma_{\text{ion}}^{nl} v_{\text{DM}} \rangle}{dE_e}, \text{ where } \frac{d\langle \sigma_{\text{ion}}^{nl} v \rangle}{d \ln E_e} = \frac{\sigma_e}{8\mu_e^2} \int_{q_-}^{q_+} q dq |f_{\text{ion}}^{nl}|^2 |F_{\text{DM}}|^2 g(v_{\text{min}}^{nl})$$

- Assume phenomenological background provided by NEWS-G, $F_{\text{DM}}=1$
- 10% Methane (plane wave) contribution

- Likelihood analysis: $\Lambda = \frac{\mathcal{L}(0)}{\mathcal{L}(\sigma_e)} \quad -2\ln(\Lambda) \sim \chi^2_1 \quad \mathcal{L}(\sigma_e) = \prod_{i=0}^{N_{\text{bins}}} \mathcal{P}(N_{\text{obs}}^i | N_{\chi}^i(\sigma_e) + N_{bg}^i)$

- With exposure 5atm.300days in sphere of radius 1.5m DarkSPHERE sensitivity below Xenon IT
- Molecular contribution seen as setting own bounds → potential for molecular bounds on DM-e scattering

Back-up – Sensitivities to higher threshold

